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PROPERTIES OF SOME ENERGY LEVELS  
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PROPERTIES OF SOME  
ENERGY LEVELS IN PHOSPHORUS-30

\* \* \* \* \*

Pharo A. Phelps



PROPERTIES OF SOME  
ENERGY LEVELS IN PHOSPHORUS-30

by

Pharo A. Phelps  
Lieutenant Commander, Civil Engineer Corps  
United States Navy

Submitted in partial fulfillment of  
the requirements for the degree of

DOCTOR OF PHILOSOPHY  
IN  
PHYSICS

United States Naval Postgraduate School  
Monterey, California

1963

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PROPERTIES OF SOME  
ENERGY LEVELS IN PHOSPHORUS-30  
by  
Pharo A. Phelps

This work is accepted as fulfilling  
the dissertation requirements for the degree of  
DOCTOR OF PHILOSOPHY  
IN  
PHYSICS  
from the  
United States Naval Postgraduate School



## ABSTRACT

The excitation function for the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  reaction has been experimentally determined in the proton energy range from 1420 kev through 2160 kev. Unreported resonances in this reaction have been identified at 1975, 2033, 2075, and 2117 kev with estimated errors of  $\pm 4$  kev. Gamma spectra at these resonances are obscured by contaminant radiation but suggest complex decay schemes. Double angular correlation measurements have been made at the 1470-kev resonance indicating a resonance level spin of 2+, or less probably, 2-. The level at 4.50 Mev is shown to probably have a spin of 2+, or less probably, 3+ rather than 0+ as has been predicted.



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## I. INTRODUCTION

As a result of successive refinements and extensions of the collective model of nuclear structure, it has been possible to extend its applicability into the region of intermediate mass nuclei ( $20 \leq A \leq 40$ ). By considering the outer nucleons to be strongly coupled to the nuclear core, excited states in  $\text{Al}^{25}$ ,  $\text{Mg}^{25}$ ,  $\text{Al}^{28}$ , and  $\text{Si}^{29}$  have been interpreted as collective rotational excited states with considerable success.<sup>1-4</sup> More recently states in  $\text{P}^{31}$  and  $\text{Si}^{30}$  have been identified with collective vibrational states of a weakly coupled collective model.<sup>5,6</sup> The experimental collection of information on the states of intermediate mass nuclei serves to verify the theoretical models, to determine the parameters which enter into descriptions of the models, and to permit the extensions and refinements to theory necessary for a more adequate description of nuclei. Of particular relevance to the present investigation is the prediction based on the weakly coupled collective model applied to  $A \approx 30$  nuclei that the 4.50-Mev level of  $\text{P}^{30}$  has a spin-parity of  $0^+$ .<sup>6</sup> In the present work the experimental determination of this value has been the object of double angular correlation measurements performed on the gamma rays observed at the 1470-kev resonance in the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  reaction.

Determinations of resonance energies in the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  reaction have been carried out by numerous investigators since 1941.<sup>7-15</sup> The proton energies which have been employed in these investigations extend over the range 300 to 1950 kev.

In addition to detecting resonances and measuring the energies at which they occur, the allied problem of analyzing the associated gamma-ray spectra has been the object of an increasingly large amount of research commensurate with improvements in gamma-ray spectroscopy techniques and instrumentation.

Investigations of the gamma rays which arise from deexcitation of the excited states of  $\text{P}^{30}$  were begun in 1954 by Endt, Kluyver, and Van der Leun.<sup>16</sup> They studied the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  reaction at resonances occurring at proton energies of 414 kev and 326 kev and assigned spin, parity, and isotopic spin to the ground state and the first excited state of  $\text{P}^{30}$  premised on gamma-transition probabilities, shell model considerations and the isotopic spin multiplet formalism. In 1956 Broude et al.<sup>17</sup>



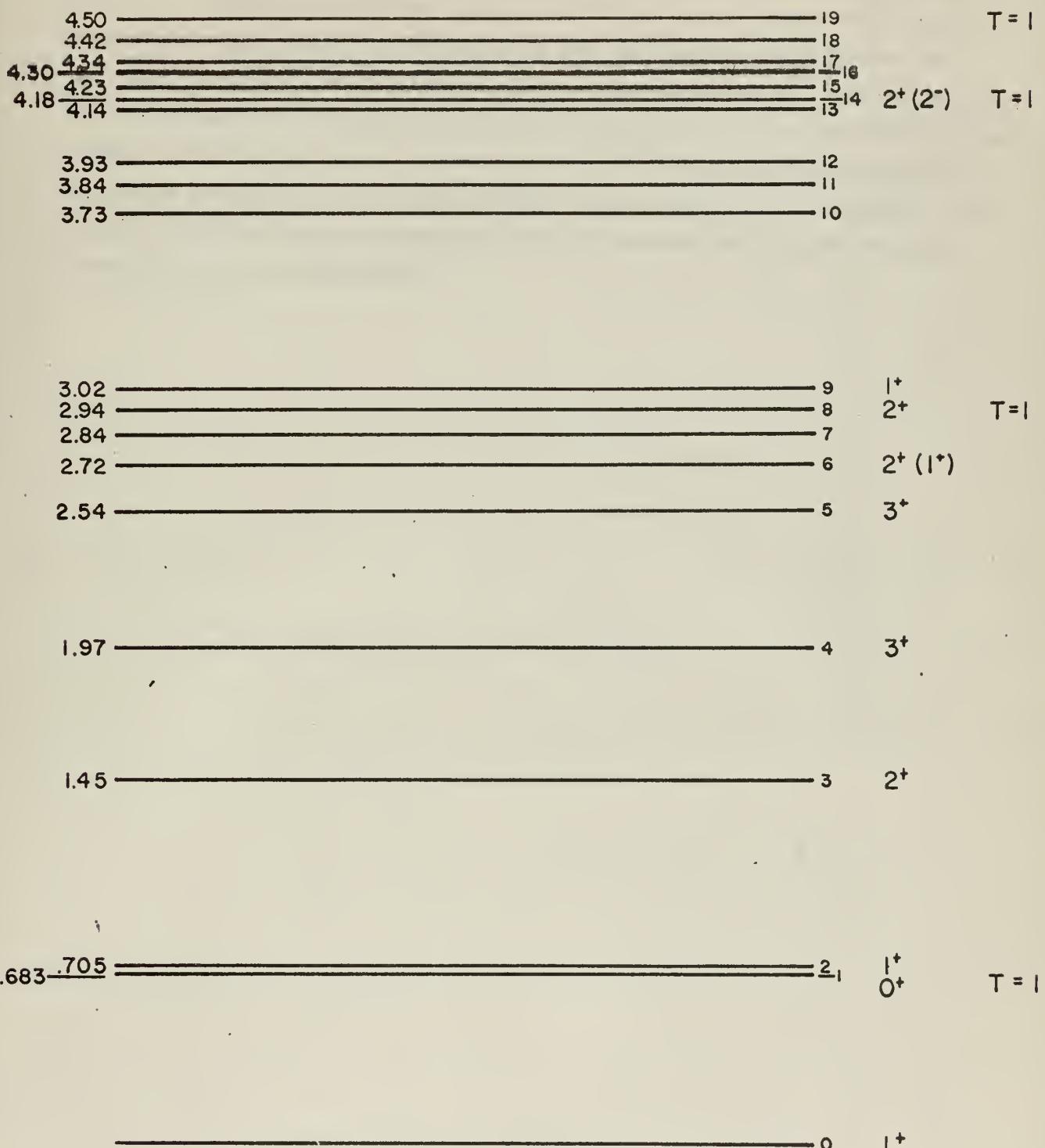
confirmed the assignments to the first excited level by measuring the angular distribution of the gamma-rays at the 414-kev resonance. They also detected five new states in  $P^{30}$  by examining gamma radiation arising from the same reaction at resonances occurring at proton energies of 696 kev and 737 kev.

Investigations by Lee and Mooring<sup>17</sup> and by Endt and Paris<sup>18</sup>, who studied the  $S^{32}(d,\alpha)P^{30}$  reaction, showed the first excited state to be in reality a close doublet. The latter investigators were able to accurately determine the energies of 30 excited levels of  $P^{30}$  up to excitation energies of 5.8 Mev by magnetic analysis of the alpha-particle groups from this reaction.

In 1958 Van der Leun and Endt<sup>19</sup> reinvestigated the four aforementioned resonances, studying the decay schemes by analysis of gamma-ray spectra, coincidence measurements, and angular distribution measurements. From this work came decay schemes, branching ratios, and assignment of spin-parities and isotopic spins to the four resonance levels and to five other excited levels of  $P^{30}$ , including the two states of the above mentioned doublet. During 1962 the gamma-ray spectra associated with the proton resonances at 1330 kev<sup>20</sup> and the resonances at 1375 kev and 1505 kev were investigated.<sup>21</sup> Proposed decay schemes were advanced for these resonances and spin-parity assignments made to three levels of  $P^{30}$ . One of the more comprehensive investigations of the  $Si^{29}(p,\gamma)P^{30}$  reaction is that of Baart et al.<sup>22</sup>, who studied six resonances and proposed decay schemes for them. Firm spin-parities were assigned to six low-lying levels and tentative assignments were made to four more levels. The most recent investigation of the reaction is that by Moore.<sup>23</sup> As a result of this work decay schemes were proposed for resonances at 1302, 1470, 1505, 1640, 1686, 1748, and 1772 kev, and spin-parities were determined for three resonance levels and two intermediate levels. The present knowledge of the quantum states of the lower  $P^{30}$  levels resulting from the above investigations is summarized in Fig. 1. Two compilations which summarize important experimental results of the above mentioned investigations are the Nuclear Data Sheets<sup>24</sup> and the compilation of information on the energy levels of nuclei with  $Z = 11$  to  $Z = 20$  by Endt and Van der Leun.<sup>25</sup>



Energies in MEV, T values are isotopic spins,  
spin-parities in parentheses are less likely alternates.



THE LOWER ENERGY LEVELS OF  $P^{30}$

FIGURE I.



The purpose of the present work has been to investigate resonances in the  $\text{Si}^{29}(\text{p},\gamma)\text{P}^{30}$  reaction in the proton energy range from 1950 kev to 2150 kev which has not previously been reported on, and to investigate the gamma-ray spectrum associated with the known resonance occurring at 1470 kev. In particular, the angular distributions of gamma rays at the latter resonance have been measured in a double angular correlation experiment to obtain information on the 4.50-Mev level in  $\text{P}^{30}$ . Additionally, some information has been obtained from examination of spectra taken at the higher resonances although contamination by  $\text{F}^{19}$  hampered this work. The following sections of this paper describe the experimental equipment, procedure, and data analysis employed, and the results which have been obtained from the investigation.



## II. DESCRIPTION OF EQUIPMENT

### 1. Experimental Layout

Fig. 2 indicates the general arrangement of equipment used in this experiment. Fig. 3 is a photograph of the target area showing the correlation table tilted into a vertical plane.

### 2. Accelerator

The 2-Mev Van de Graaff horizontal electrostatic accelerator of the U.S. Naval Postgraduate School, manufactured by the High Voltage Engineering Corporation, was used in this work as a source of protons. The radio frequency ion source is of the type described by Moak.<sup>26</sup> A proton spin-resonance fluxmeter for accurately measuring the analyzing magnet's field was installed during the course of the work. An error signal from the fluxmeter was used in a feed back loop to the power supply of the beam analyzing magnet to hold the field strength automatically at a set value. This feature simplified the maintenance of a given resonance energy for long bombardment times without "slipping off".

A current integrator of the vacuum tube electrometer type was used to measure beam current. High stability plastic insulated capacitors were used to collect the charge. The voltage across the capacitor in use was applied to the grid of a 954 acorn tube, whose output fired a Schmidt trigger circuit at a preselected voltage level. The trigger output actuated other circuitry to discharge the capacitor and advance a Veeder-Root cycle counter after each cycle.

### 3. Target Assemblies

Two types of targets and target mounts were developed and used in this experiment. These are shown in Fig. 4. For resonance and coincidence measurements, which were made in the plane perpendicular to the beam, the target backing consisted of a one-half-in-diameter flanged copper hemisphere held against a flat copper boss on the mount by a threaded retaining ring. This assembly was designed with the beam direction as a symmetry axis so that it can be used for triple correlation experiments in the perpendicular plane with no perturbations from target geometry. For the double correlations, measured in the horizontal plane containing the beam, target backings were 0.046 cm thick copper strips clamped in a copper mount. Both types of backings and mounts were vacuum



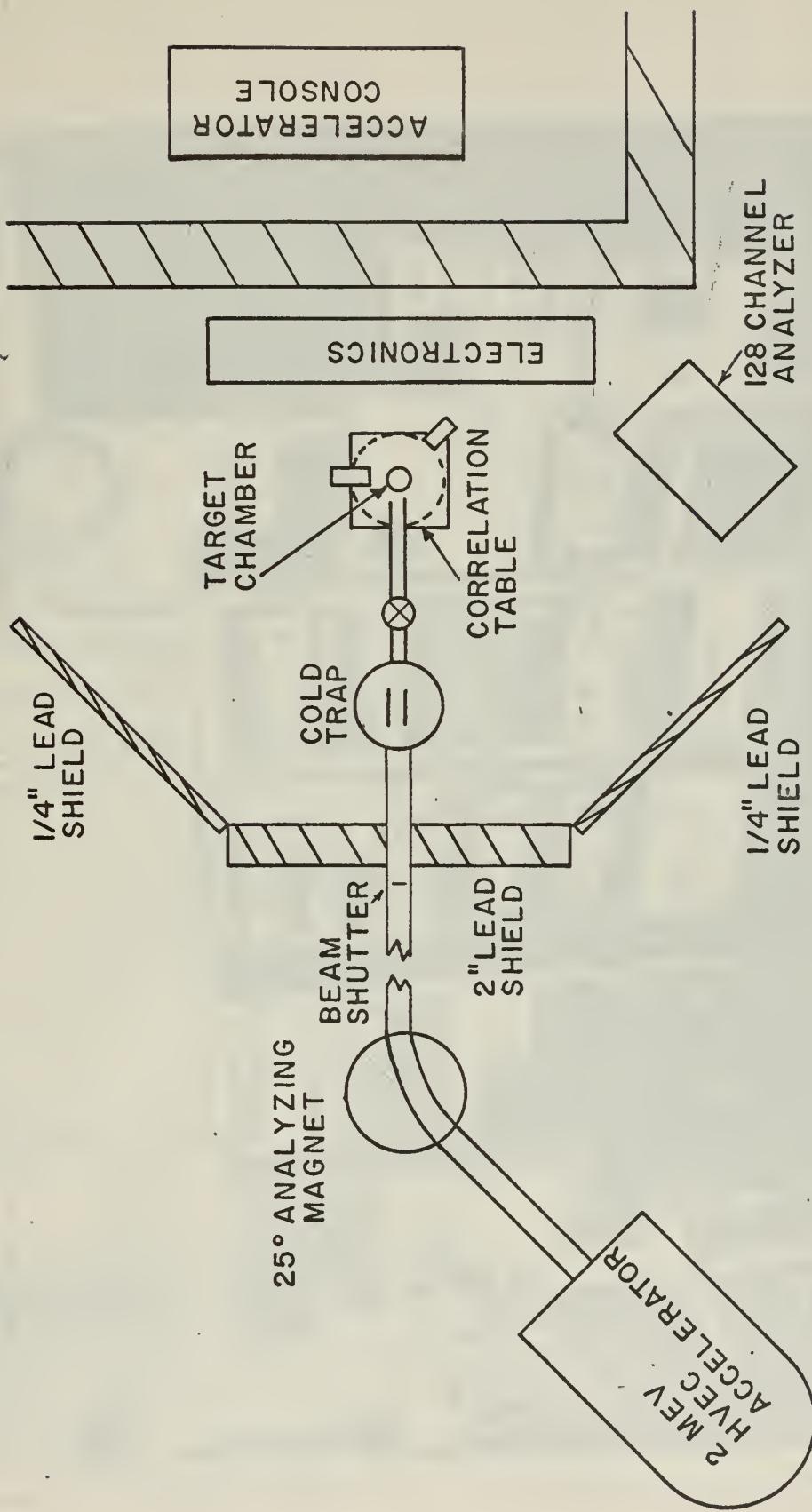


FIGURE 2. EXPERIMENTAL LAYOUT



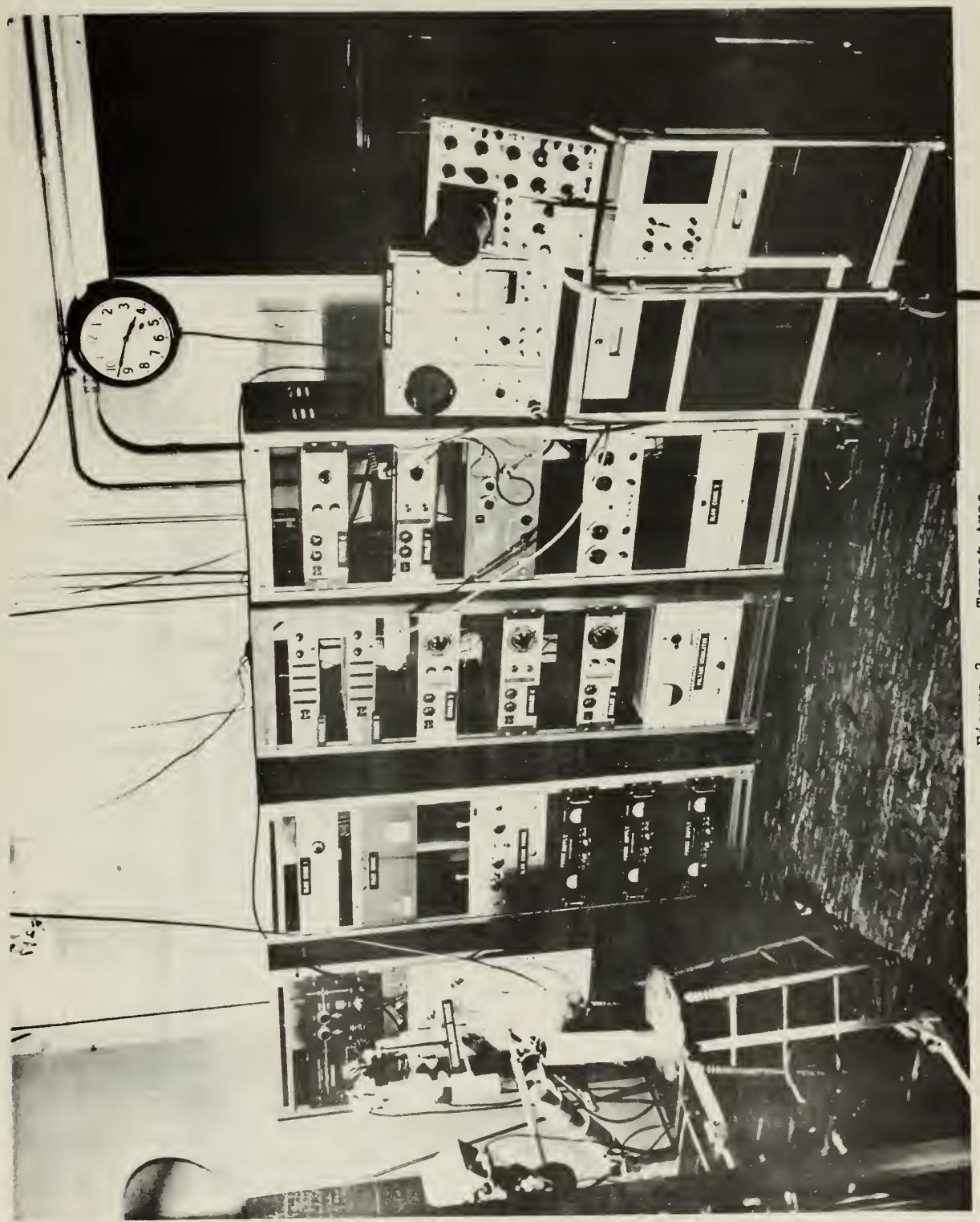


Figure 3 Target Area



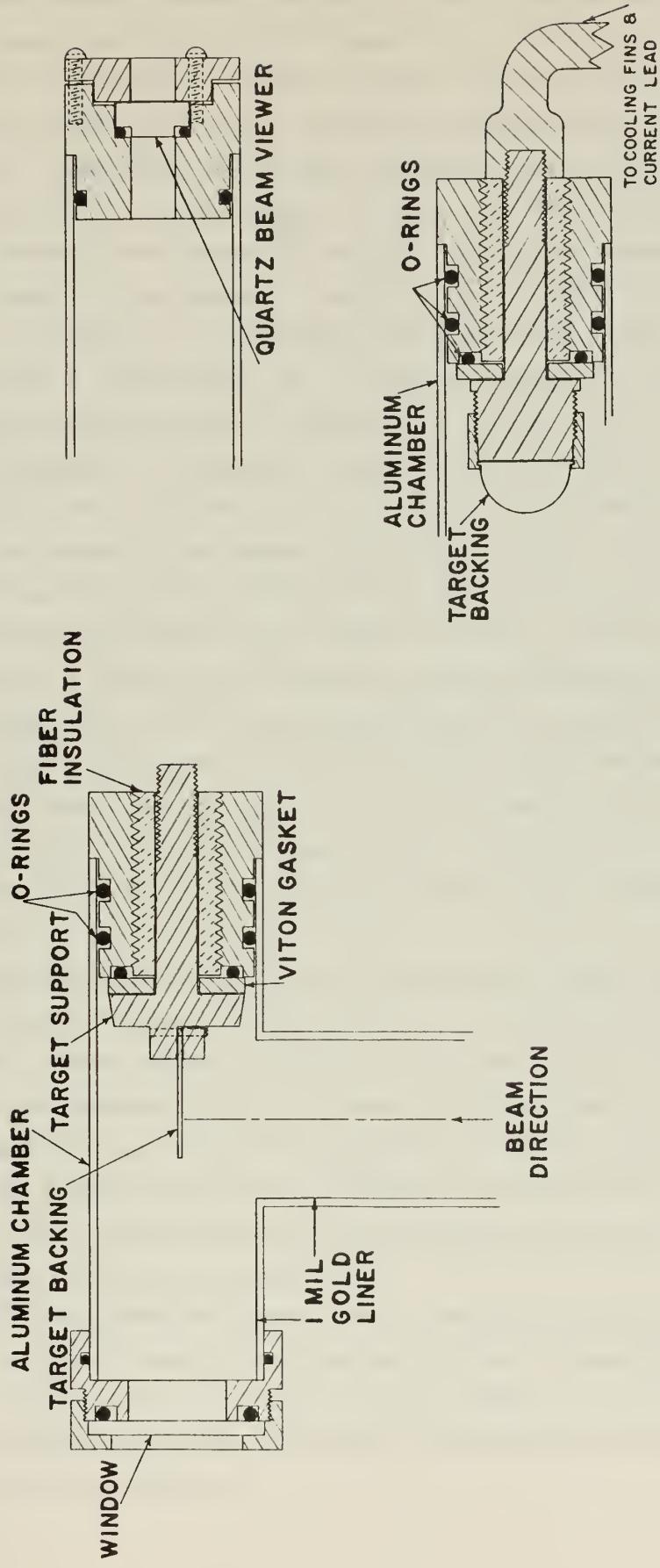


FIGURE 4. TARGET ASSEMBLY DETAILS



plated with a layer of gold thick enough to stop incident protons. For both geometries the target mount was fitted inside one end of a one-in. diameter tubular chamber, the axis of which coincided with the beam path. Both target assemblies were designed so as to keep mass in the target vicinity to a minimum. Competing with this requirement was the need to cool the target in order to prevent any appreciable target material evaporation. Maximum possible beam intensities were desired so as to take advantage of the better counting statistics which accompany high yields. Still another requirement of angular correlation measurements is a small target spot to approximate a point source of gamma rays. A beam collimator consisting of two five-mil tantalum discs separated by 0.4 in. was mounted in a brass holder 38 in. from the target. A 0.096-in. diameter collimating aperture was followed by a 0.098-in. hole in the second disc for the purpose of stopping protons scattered from the edge of the collimating disc. A cylindrical lead shield surrounded the collimator in order to attenuate any gamma rays produced here.

As is described in Chapter III, target material was evaporated directly onto the gold-plated copper backings to permit good heat conduction. Initially, a hollow copper target backing through which coolant could be circulated was tried, but was abandoned for a simpler solid copper backing cooled by conduction through a finned copper extension leading to an external Dewar flask. Liquid air in the Dewar proved to be convenient as a coolant at high beam powers and, below about ten watts, target power dissipation by forced air cooling of the target extension fins proved adequate.

After unsuccessful trials of Lucite and glass target tubes, aluminum was finally adopted for this purpose. An auxiliary target assembly with a quartz window target was used to accurately center the proton beam spot. The target chamber tube, which was insulated from ground at both ends, was maintained at a negative potential of 300 volts relative to ground in order to suppress secondary electrons from the target. In order to prevent recoil protons from exciting aluminum resonances, the target tube was lined with one-mil gold foil.

The hemispherical copper target backings for measurements in the plane perpendicular to the beam were needed in some quantity, so were "mass produced" in a small hand forge made for this purpose. A surface which needed only minor polishing and cleaning before gold plating was obtained in this manner.



A frequently troublesome feature of  $(p,\gamma)$  investigations is the gradual accumulation of carbon on the target as a result of cracking of pump oil vapor by particle bombardment. To reduce this difficulty, advantage was taken of an 8-in. glass cross which was mounted between the target diffusion pump and the target chamber as shown in Fig. 5. A special liquid air cold trap was constructed and installed in this cross to serve a dual purpose. First, it functioned as an ordinary cold trap with a large surface of copper vanes suspended from the coolant reservoir to provide an optically dense path from the diffusion pump thus permitting the vacuum to be routinely maintained close to  $10^{-7}$  Torr. The second feature of the trap was a 1-in. diameter copper pipe piercing the vanes which the proton beam traversed. At high vacuum the proton beam functions effectively as a pump, transporting gases and vapor encountered in its path. When the beam passes through the restricted cold pipe for 6 in. before entering the target chamber, it was reasoned that its drifting retinue of condensable vapor contaminants would be largely trapped by the cold inner wall. The results from use of the trap support this idea since only minor discoloration was observed on targets compared with that which occurred when the trap was not used, and no increase in bombardment energy was found necessary as targets aged.

#### 4. Shielding

To reduce the radiation level at the target resulting from bremsstrahlung and X-rays caused by electrons accelerated toward the high voltage accelerator terminal, a considerable amount of lead shielding was erected. A 2-in.-thick lead wall, 5 ft. high and 4 ft. wide, was placed between the energy control slits and the beam collimator. In addition, two lead screens 1/4-in. thick and each 5 ft. by 4 ft. were bolted onto a caster supported steel framework constructed for this purpose. The mobility of the screens permitted their placement to minimize the radiation level at the target and also allowed easy removal for access to the target area. Although the above-described shielding offered some reduction in natural background radiation, (notably the 1.46-Mev  $K^{40}$  gamma ray) it was not feasible to screen specifically against background without either increasing scattered target radiation at the detectors or utilizing an excessive amount of lead.



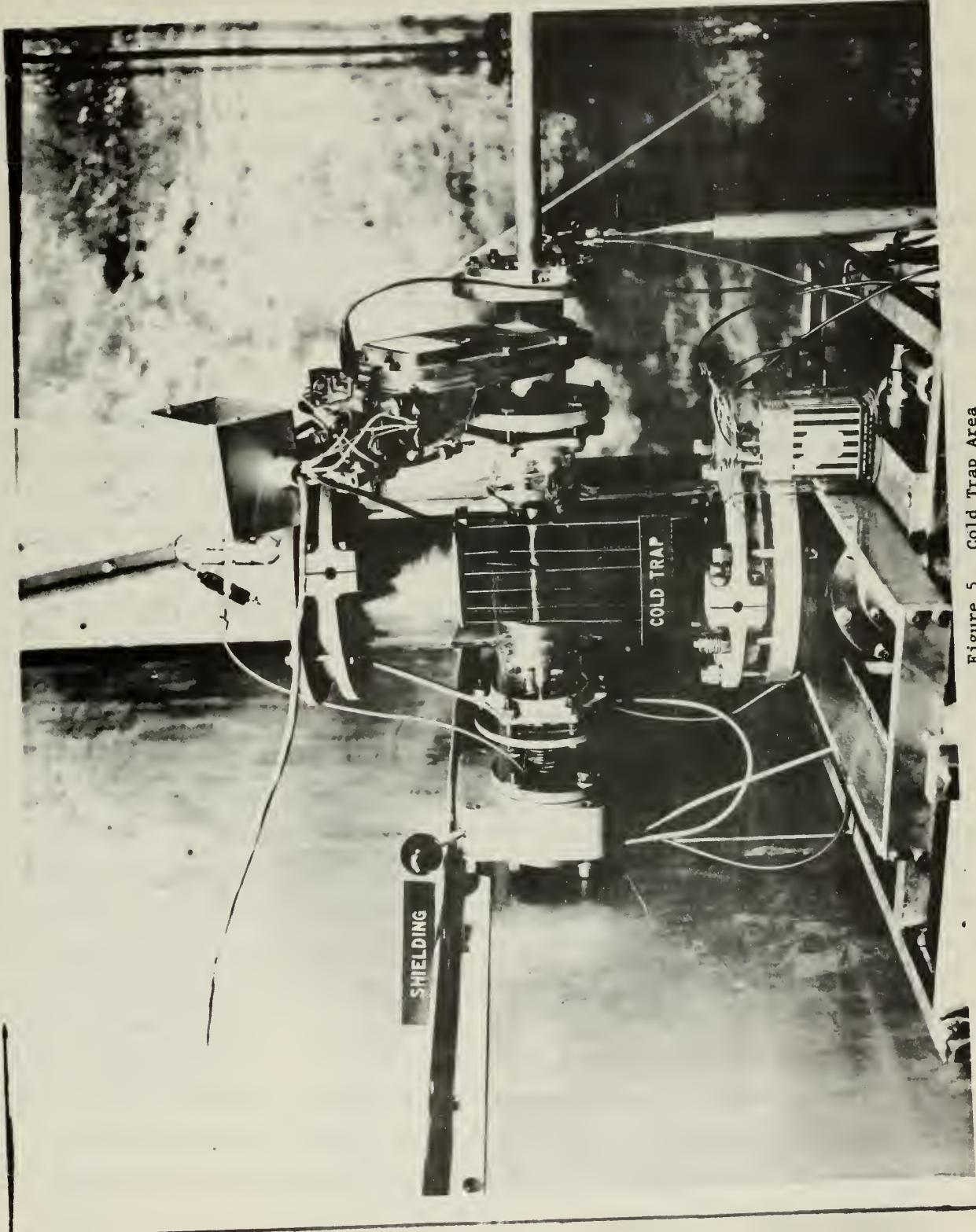


Figure 5 Cold Trap Area



## 5. Correlation table

A correlation table shown in Fig. 6 was built for this experiment to support the gamma-ray detectors and to permit their accurate positioning with respect to the target. A steel base plate mounted on casters supported a hollow magnesium column upon which the table was mounted. Screw jacks in the base plate were extended when the apparatus was in place to hold the table in position and to allow accurate vertical adjustment. The table itself consisted of a movable graduated aluminium disc which was mounted concentrically on a fixed aluminum support plate. Slides mounted on the disc and on the support plate held mounts for the two gamma-ray detection assemblies. The mounts were positioned in the slides by threaded rods which allowed the detectors to be set accurately at distances from 1/2 in. to 6 in. from the target position. The detection assembly mounted on the disc could be positioned in angle with respect to the fixed detector in 12 min. increments of arc by means of an adjusting gear meshing with gear teeth on the periphery of the disc. The entire table assembly was attached to the magnesium support column by a hinge so that it could be tilted into either a horizontal or vertical plane, permitting measurements to be taken in either the horizontal plane containing the proton beam or in the plane perpendicular to the beam. The first named arrangement is designated herein as geometry A and the latter as geometry B.

## 6. Gamma-ray detectors

The gamma-ray detection assemblies were supported with their center-lines 8 3/4 in. above the table. They each consisted of a 3 in. x 3 in. cylindrical sodium iodide (thallium activated) crystal integrally mounted with a Dumont 6363 photomultiplier tube supplied by the Harshaw Chemical Company. In the earlier phases of this work, similar 2 in. x 2 in. crystals mounted with Dumont 6292 tubes were used. The resolution of the 3-in. crystals, measured after the signal had traversed a 12-microsecond delay line input to the pulse height analyzer, was 9% for the 0.662-Mev Cs<sup>137</sup> gamma ray. A 6AK5 cathode follower was constructed into each detector mount to serve as a preamplifier.

## 7. Electronics

The electronic equipment and circuitry for processing signals from the detectors were planned so as to permit considerable flexibility in experimental applications. A block diagram of the basic layout is shown in Fig. 7.



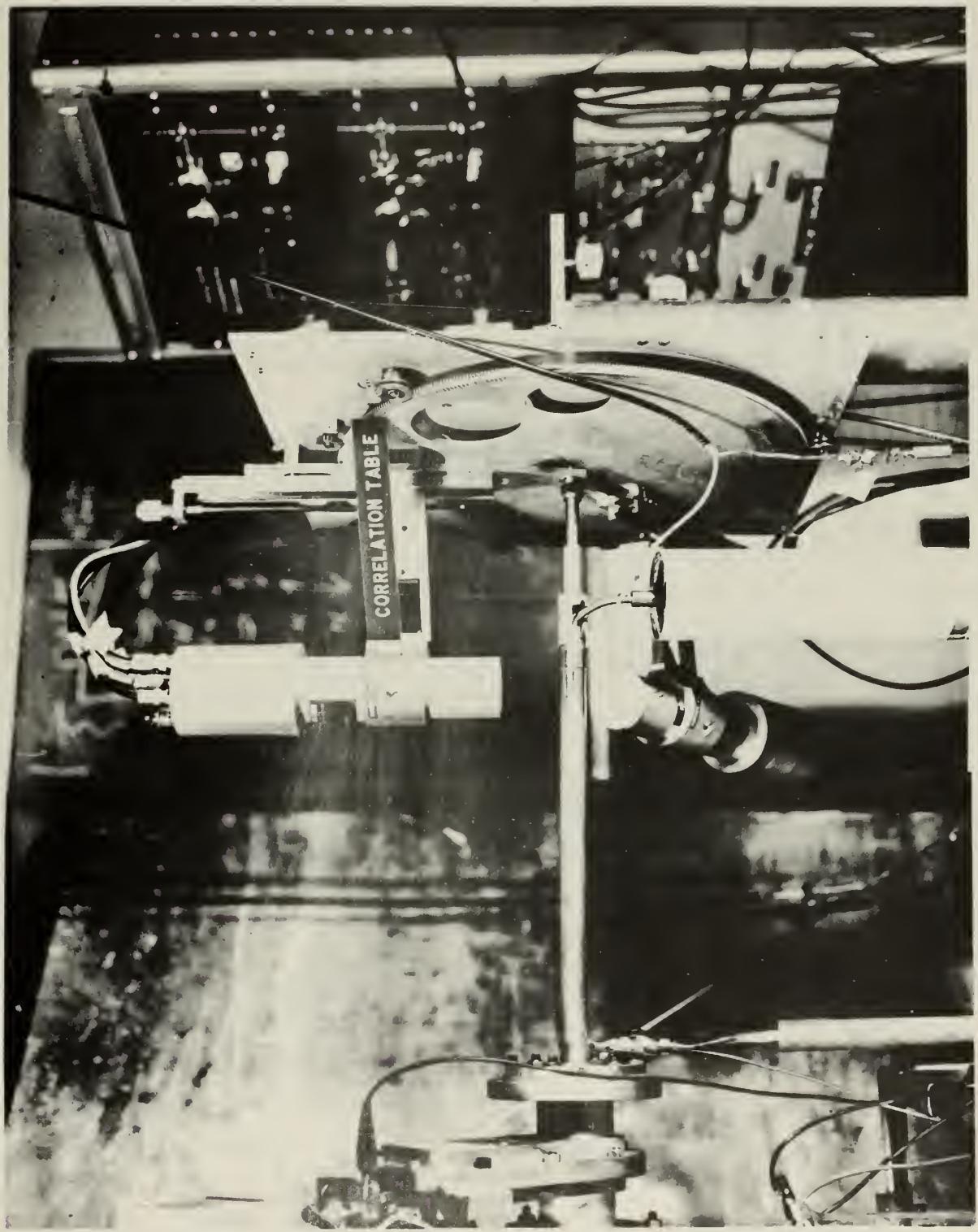


Figure 6  
Correlation Table  
13



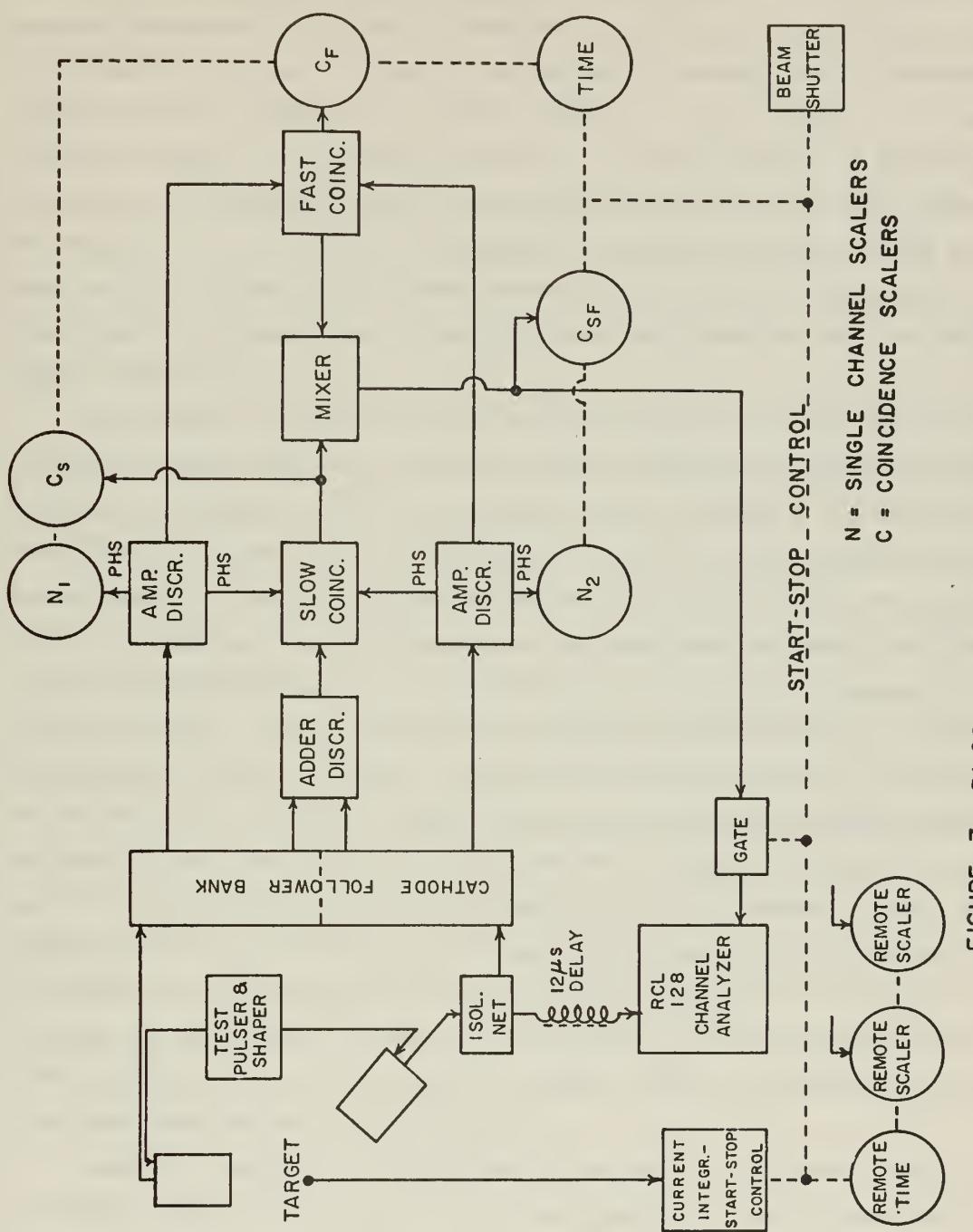


FIGURE 7. BLOCK DIAGRAM OF ELECTRONICS



The system is composed of an RCL 128-channel pulse height analyzer to record pulse height spectra, a slow-fast-sum coincidence analyzer, and scalers for recording various outputs.

It was found that requirements for the detector output signal by several components, including a monitor oscilloscope, tended to overload the preamplifiers and to permit unacceptable interaction of subsequent stages with each other through the coupling of their inputs. To avoid this problem a bank of six cathode followers was constructed. A bussing arrangement made it possible to feed the grids of the followers from either one of two busses or from an individual input, so that up to six pieces of equipment could be driven by the same input pulse with no interaction and with practically no signal loss. Each follower was provided with an individual output gain control.

The output of either detector as desired was also fed directly into the pulse height analyzer through a simple resistance isolation network. From the resistance network the signal passed through a terminated 1350-ohm delay line to allow time for formation of gating pulses needed for coincidence spectra. Several methods of injecting a signal into the analyzer were tried inasmuch as this equipment was designed to accept an input only from a photomultiplier tube connected directly into the preamplifier assembly provided. This was not an acceptable practice because there existed no convenient means of signal extraction for other purposes, and because the preamplifier utilized a high-voltage supply which required running the photomultiplier cathode at a high negative potential with respect to the tube shield with the attendant probability of tube damage. The procedure adopted as most successful was to apply the signal from the delay line to the anode pin of a 14-pin tube base which was connected directly into the analyzer preamplifier. It was also necessary to incorporate a gate pulse-forming network into the analyzer in order to meet the gating pulse specifications required by this equipment.

Gamma-ray spectra were obtained by gating the analyzer with the discriminator output pulses from the linear amplifier fed by the same detector. This procedure excluded a weak precursor pulse which was transmitted by the delay line input, by not gating the analyzer on until a few microseconds before arrival of the delayed pulse. No significant loss in detector resolution resulted from leaving the delay line in the input for



ordinary gamma-ray spectra, and it was found quite convenient to be able to switch from these singles spectra to coincidence spectra without change in analyzer calibration.

For the observation of coincidence spectra, the analyzer was gated by the slow-fast coincidence output. The basic slow-fast coincidence system assembled for this experiment has been described elsewhere.<sup>27</sup> In this system the output from an Atomic Instrument Company slow coincidence unit accepting up to four inputs was combined in a diode mixer with the fast coincidence unit output to obtain the slow-fast coincidence output signal. The fast unit, based on the circuit of Draper and Fleischer,<sup>28</sup> detects coincidences between amplified pulses from the scintillation detectors with a resolution time as low as ten ns. For pulses spanning a large amplitude range, however, more dependable operation was obtained when the resolving time was increased to about 80 ns. This was due to an observed amplitude dependence of the shape of the leading edge of the amplified pulse, to which the fast unit is sensitive.

Several modifications have been made to the coincidence system which do not appear in the previous description. The diode mixer unit for combining the slow and fast coincidence outputs has been modified by inserting a Schmidt trigger between the adding stage and the output stage in order to clip the sum pulse and improve output-pulse characteristics. The fast unit has been modified by providing a variable delay line in the input in order to permit accurate compensation for differences in time delays between the detectors and the coincidence unit. Finally, there has been added to the system a provision for obtaining sum-coincidence spectra as first described by Hoogenboom.<sup>29</sup> In this system the pulses from the two detectors are carefully matched in amplitude calibration, then summed in a pulse adding circuit. These summed pulses are analyzed in a differential discriminator set to accept pulses corresponding to the excitation energy of a resonance level. This output is then used as a third input to the slow coincidence unit so that its output represents coincidence between two completely absorbed gamma rays comprising a double cascade from the resonance level. This signal is then used to trigger the pulse height analyzer gate. The spectrum obtained from pulse height analysis of one detector output thus consists only of full energy peaks due to radiations which are members of a double cascade. The pulse height adder which has been constructed for this purpose employs the circuit used by Kantele and Fink in a somewhat different type of spectrometer.<sup>30</sup> The adder output was analyzed by a conventional linear amplifier and discriminator.



The scalers used to record single channel counts had dead times of one microsecond and those recording coincidences had dead times of approximately eight microseconds. Separate cathode-follower outputs to drive scalers were constructed for all coincidence stages so the relatively low scaler input impedances would not degrade coincidence pulses. The number of scalers provided, which exceeded the requirements for the present experimental measurements, permitted a continuous check on the performance of the coincidence system.

A double pulser with associated pulse shaping networks and an oscilloscope formed an integral part of the equipment and were used for all routine checks and adjustments. An additional slow coincidence unit of variable resolving time<sup>31</sup>, which was constructed as part of the system, was used to check on the operation of the Atomic Instrument Company unit.

Control circuit relays were provided so that the pulse height analyzer and all scalers could be started and stopped simultaneously by the current integrator on-off switch.



### III. EXPERIMENTAL PROCEDURE

#### 1. Target preparation

The target material used in this experiment was isotopically enriched silicon in the dioxide form purchased from the Stable Isotopes Division of Oak Ridge National Laboratory. Relative percentages of the three stable silicon isotopes were 91.85 Si<sup>29</sup>, 7.85 Si<sup>28</sup>, and 0.30 Si<sup>30</sup>. This material was evaporated onto prepared target backings in a vacuum evaporation apparatus under a chamber pressure of 10<sup>-5</sup> to 10<sup>-6</sup> Torr.

Beam power of as much as 20 watts was delivered to a target spot area of about 6 square millimeters during bombardment. Dissipation of this heat by 5-mil tantalum target backings, which had been used in previous work at this laboratory, was found insufficient to prevent local overheating of the target with an accompanying rapid loss of target material. This fact, and the fluorine contamination of the surface of all sheet tantalum examined, led to employment of the heavier target backings of copper described in Chapter II. While vacuum plating these copper backings with 60,000 Å or more of gold (99.999 percent purity) to completely stop 2-Mev protons, it was found that the plating tended to crack and spall off when the copper backings were maintained at ambient temperature during evaporation. This effect was caused chiefly by tensile stresses which have been found to be as much as 850 Kg/cm<sup>2</sup> in 2000 Å gold films deposited on copper.<sup>32</sup> This trouble was overcome by holding the copper backings at 400°C during evaporation by means of a resistance heater which was constructed for this purpose. The heater was made in the form of a 4-in. cylindrical section, to which the target backings were clamped. The cylindrical heater section was suspended in the vacuum coating chamber so that its axis coincided with the evaporant source. Heating the backings in this manner prior to evaporating both gold and SiO<sub>2</sub> also served to remove any residual volatile contaminants leaving a clean surface to receive the coating.

Evaporating boats were fabricated from 5-mil tungsten ribbon and, when cleaned by preheating before use, proved to be satisfactory for the evaporation of both gold and SiO<sub>2</sub>. Temperature of the SiO<sub>2</sub> evaporant was monitored with an optical pyrometer to maintain an indicated evaporating temperature of 2130°C; at this temperature 4.0 mg of SiO<sub>2</sub> evaporated in



about two minutes. The temperature is not corrected for departure of the evaporant from an ideal black body inasmuch as reproducibility and temperature control were desired rather than an accurate measurement of true surface temperature. During the first five to ten seconds of all evaporations the evaporant was intercepted by a shutter, since the first part of an evaporant to boil off is known to carry a higher percentage of any contaminants present. Following evaporation, cooling of the heater and targets was accelerated by admitting pure helium to the chamber through a liquid air trap.

In practice, 2 gm or more of gold were evaporated at a distance of 5 cm from the target backings and exposed parts of target assemblies to produce the desired plating thickness. 4.1 mg of  $\text{SiO}_2$  were evaporated at the same distance to give targets approximately 5-kev thick to 2-Mev protons for use in the excitation function runs. For the targets used to obtain gamma-ray spectra, and for angular correlation runs at 1470 kev, 5.2 mg of  $\text{SiO}_2$  were evaporated to obtain targets 8-kev thick to 1470-kev protons.

Considerable care was used in an attempt to reduce the contamination by  $\text{F}^{19}$ . Evaporation boats were preheated to  $2200^{\circ}\text{C}$  before use, targets were stored in a vacuum (until this was found not to be of value), target chamber pressure was maintained at  $10^{-7}$  Torr during runs, and possible sources of contamination in the vacuum systems such as Teflon were avoided. In preliminary bombardments at the  $\text{F}^{19}$  resonance energy of 1373 kev, the gold-plated copper target backings, which were washed in either pure ethanol or menthanol and rinsed in distilled water just prior to use, were found to be freer from  $\text{F}^{19}$  than backings of either Ta or W which had been cleaned by various methods. In spite of these precautions all targets which were prepared proved to be contaminated to some degree by  $\text{F}^{19}$ .

## 2. Excitation function measurements

The excitation function of the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  reaction has been measured by recording single gamma-ray yield and gamma-gamma coincidence yield at accelerator energy settings over a range of proton energies from 1420 kev to 2170 kev. The 2 in. x 2 in. NaI(Tl) scintillation detectors were used for this purpose. The output of one of the detectors was counted to obtain the gamma-ray yield at each proton energy, and the number of events in which gamma rays were detected simultaneously in each crystal was recorded by the use of the slow-fast coincidence system adjusted to a resolving time of 25 ns.



The two detectors were positioned in geometry B at angles of  $0^\circ$  and  $180^\circ$  relative to the vertical direction and at a distance of 0.6 in. from the beam spot. Integral discriminators in both channels were set so that counts would be recorded for gamma rays with energies exceeding 0.65 Mev, a setting chosen to exclude the 0.51-Mev annihilation radiation arising from the positron decay of  $P^{30}$ , while accepting the full energy peaks of the 0.68 and 0.71 Mev radiations which deexcite the first two excited states in  $P^{30}$ . The 7-microampere proton beam was collected by the current integrator which closed a beam shutter and stopped counters after collection of 84 microcoulombs at each energy step of approximately 3 kev. Background counts were made every 50 kev with the beam intercepted by a shutter in the beam tube outside of the shielded target area.

An auxiliary experiment was conducted to determine the energy calibration for the excitation curves. In this experiment the frequency of the NMR signal was measured to within one part in  $10^{-6}$  as the 992-kev aluminum resonance  $^{33}Al$  was traversed using a freshly prepared target about 10 kev in thickness. Runs with mass 2 beams were also made to check this calibration at a generator energy of 1984 kev. Immediately after calibration, the  $Si^{29}$  resonance energies above 1970 kev and the energy of the resonance at 1506 kev were carefully measured in a similar manner. Using these points as references it was then possible to calibrate the entire excitation curve in terms of magnet current, which was determined to a precision of one part in  $10^{-5}$  at each energy step. This procedure gave a better calibration than that obtained using the fluxmeter logging dial readings, since calibration curves for converting these readings to magnetic field could not be read with adequate precision.

### 3. Measurement of saturation activity

In order to definitely identify resonances due to the  $Si^{29}(p,\gamma)P^{30}$  reaction, measurements were made at the resonances above 1970 kev to determine the half life of decay radiation following bombardment. In these measurements  $Si^{29}$  targets were bombarded for ten minutes, then removed to a small cave of 2-in. lead brick where counting was started within 60 sec after cessation of bombardment. The positron annihilation radiation following beta decay of  $P^{30}$  was detected by a 3 in. x 3 in. crystal. Total counts above the noise threshold were recorded by the pulse height analyzer operated in a gross counting mode. In this mode the channel in which all counts are recorded is automatically shifted



ahead after a preset counting time. After counting for 3.1 sec per channel for a 6.46-minute total counting time, the target was removed from the cave and a background run made with the analyzer in the subtract mode.

Difficulty which was experienced in obtaining gamma-ray spectra recognizably due to Si<sup>29</sup> at the upper resonances as a result of masking radiation from the F<sup>19</sup>(p, $\alpha\gamma$ )O<sup>16</sup> reaction caused some concern that the peaks observed in the excitation curve might have been due to other contaminants. Observance of the 2.5-min half life of P<sup>30</sup> following bombardment could not entirely exclude this possibility since it is conceivable for P<sup>30</sup> to be formed in a nonresonant direct capture process. For this reason it was desirable to compare the activity of the target observed after bombardment at energies corresponding to the peaks in the yield curves to the activity observed after bombardment at off-peak energies. In an experiment to do this, a proton bombardment time of ten minutes to obtain 94 percent of saturation activity was followed by a ten-minute counting period during which the pulse height spectrum was obtained. The target was then allowed to decay for an additional ten minutes before bombardment was resumed at the next energy. The number of counts observed in the 0.51-Mev annihilation radiation peak then served as a measure of the amount of P<sup>30</sup> formed at each energy.

#### 4. Gamma-ray spectra

All gamma-ray spectra were obtained using 3 in. x 3 in. NaI(Tl) crystals with no special shielding nor collimation. For most spectra the pulse height analyzer usable range extended from about channel 7 to channel 127 corresponding to gamma-ray energies from 0.7 to 7.7 Mev. For observation of annihilation radiation a higher gain was used so that these channels corresponded to energies from 0.18 to 1.92 Mev. The analysis of spectra for intensities was performed with a digital computer as described in Appendix 2. The particular method finally used to subtract background radiation from spectra varied according to the resonance being considered. When possible, backgrounds taken slightly above or below resonances were subtracted from spectra. This was not very successful for the resonances above 1950 kev where F<sup>19</sup> contamination gave rise to the F<sup>19</sup>(p, $\alpha\gamma$ )O<sup>16</sup> reaction. This reaction is strongly resonant at 2030 kev with a width at half maximum of 120 kev<sup>34</sup> which effectively blankets the higher proton energy range considered in the present work. The 6-7 Mev



gamma rays occurring in this reaction are present in relative proportions which vary considerably with bombarding energy so that off-resonance background subtraction is not accurate. Bombardment of blank targets which had been processed identically to  $\text{SiO}_2$  targets but using an empty evaporation boat gave a background measurement which was also rather unreliable because of individual variations among targets. A method of subtracting the background measured with the accelerator off was also used, in which case background radiation of specific energies due to contaminants was subtracted out in the spectrum analysis.

In order to circumvent radiation due to target contaminants efforts were exerted to obtain sum-coincidence spectra in the energy region above 1950 kev as a means of getting at decay schemes. The sum-coincidence spectra were, in general, measured with the slow-fast coincidence requirement also imposed, and with a coincidence resolving time of 110 ns. These runs were usually made in geometry B with counters positioned 1.6 in. from the beam spot at a  $90^\circ$  angle relative to each other. A 5/16-in. lead sheet between the counters served to minimize coincidences due to scattering or pair production in one crystal with the lost radiation being captured in the other crystal. The differential discriminator following the pulse adder was set with its lower level at the resonance energy and was adjusted to pass pulses in a band about 0.6 Mev in width. Settings were made by adjusting the discriminator so that the observed count rate of accepted pulses was equal to half the repetition rate of a test pulse. The proper test pulse amplitude was obtained by setting the pulser gain so that counts fell into the pulse height analyzer channel corresponding to the desired energy. Settings of integral discriminators in the two singles channels were adjusted to 0.62 Mev in a similiar manner. The pulse height analyzer was calibrated with  $\text{Cs}^{137}$  and  $\text{Co}^{60}$  sources and with the 5.31-Mev gamma ray which dominates the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  spectrum at the 414-kev resonance. It was found that the Dumont 6363 photomultiplier tubes experienced marked gain shifts as a function of count rate amounting to several percent, an effect which has been attributed to changes in the dynode surfaces<sup>35</sup>. Attempts to allow this shift to stabilize before commencing counts were only partially effective. The overall result of this phenomenon was to vitiate the results of a large number of sum-coincidence spectra.



## 5. Measurements at 1470 kev

Since the 4.50-Mev level in  $P^{30}$  has been shown to be most strongly excited at the 1470-kev resonance, this proton energy was a logical choice for angular correlation measurements which could lead to determination of the spin-parity for the level. The resonance is rather weak so it was decided to take five sets of data to improve the counting statistics and to permit a check of the consistency of results. Each set of data consisted of five spectra taken with the movable detector set to angles of  $0^\circ$ ,  $30^\circ$ ,  $45^\circ$ ,  $60^\circ$ , and  $90^\circ$  relative to the proton beam. In geometry A with the 3 in. x 3 in. movable detector positioned with its face 4.125 in. from the target, runs were made using a 7-microampere beam for counting times of approximately one hour at each angular setting. The same target was used for each set of runs and the order of the angular settings was varied in different sets so as to minimize differences in the gamma-ray spectra due to any contaminant buildup or target erosion. The fixed 3 in. x 3 in. detector was positioned at  $270^\circ$  relative to the beam and 4.125 in. from the target for use as a monitor counter. Each run was terminated when 30,000 monitor counts had been recorded. A 2 in. x 2 in. detector was also used as a secondary monitor counter. The integral discriminator in each monitor channel was adjusted to a valley in the gamma-ray spectrum at 4.75 Mev so that the effect of gain shifts in these channels would be minimized. Before and after each run the discriminator levels were checked and adjusted with the aid of a  $Co^{60}$  source which was precisely and reproducibly located relative to the counters. For this check the gain in the monitor channels was increased by a factor of 4 by means of the gain multiplier switch so that the discriminator setting was equivalent to 1.14 Mev. At this setting, near the 1.17-Mev peak in  $Co^{60}$ , the count rate observed with the  $Co^{60}$  source was quite sensitive to the discriminator setting so that accurate adjustment was possible.

The unfolding of gamma-ray spectra and the determination of angular distribution coefficients was programed and analyzed on a CDC-1604 computer as described in Appendix 2. The radiations which have been used as spectrometer calibration curves in this analysis are listed in Table 1. These spectra were measured in experimental geometry A, backgrounds were subtracted, and any remaining contaminating gamma rays were subtracted off. Each of these monoenergetic spectrograms was plotted, hand smoothed, and redivided into 525 channels to reduce any interpolation errors in the



TABLE 1  
Gamma Rays Used for Spectrometer Calibration

$E_{\gamma}$ (Mev)	Radiation Source	$E_p$ (kev)
0.662	$\text{Cs}^{137}$	-
1.63	$\text{Na}^{23}(\text{p}, \alpha \gamma) \text{Ne}^{20}$	1252
2.37	$\text{C}^{12}(\text{p}, \gamma) \text{N}^{13}$	460
2.75	$\text{Na}^{24}$	-
3.51	$\text{C}^{12}(\text{p}, \gamma) \text{N}^{13}$	1698
4.43	$\text{B}^{11}(\text{p}, \gamma) \text{C}^{12}$	675
5.31	$\text{Si}^{29}(\text{p}, \gamma) \text{P}^{30}$	414
6.135	$\text{F}^{19}(\text{p}, \alpha \gamma) \text{O}^{16}$	1374
8.06	$\text{C}^{13}(\text{p}, \gamma) \text{N}^{14}$	554



subsequent analysis. Energy calibration of each spectrum was made by equating distance in channels between the pair peak and the photopeak to 1.022 Mev. For the three lowest energy calibration curves, for which pair peaks are non-existent or poorly defined, an equivalent calibration was computed based on the pulse height analyzer calibration as determined in an auxiliary experiment.

All spectra were read out of the pulse height analyzer memory photographically. These data were then transcribed to punched cards and printed listings of the cards were carefully proofread against the photographs.

Gain drifts did not prove troublesome at the lower count rates observed during correlation runs at 1470 kev. To compensate for differences in calibration of the pulse height analyzer between successive runs, which were of the order of one percent, each spectrum was self-calibrated by the peaks at 7.00 Mev and at 1.45 Mev.



#### IV EXPERIMENTAL RESULTS

##### 1. Excitation Function

The excitation function for  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  as measured by gamma-ray yield is shown in Fig. 8, and as measured by gamma-gamma coincidence yield in Fig. 9. The energies at which resonances are observed are listed in Table 2, together with some of the results which have been obtained at proton energies below 1900 kev by other investigators.

In arriving at the uncertainty of 0.20% kev for the resonance energies, consideration has been given to errors arising from instability and inhomogeneity of the field of the analyzing magnet, the proton beam spread, possible surface contaminants, calibration drift, and the 0.5-kev estimated error for the 992-kev aluminum calibration point.<sup>33</sup>

Particular attention was given to the resonances above 1900 kev which have not previously been reported. The results of determinations of the half life of the post-bombardment activity of the targets at these energies are given in Table 3. The measurements are seen to agree with values for the half life of  $\text{P}^{30}$  reported in the literature, which range from  $2.51 \pm .02$  min to  $2.62 \pm .02$  min,<sup>24</sup> so that radiative capture by  $\text{Si}^{29}$  is clearly demonstrated. Table 4 shows the intensity of positron annihilation radiation observed after saturation at selected bombardment energies. These intensity values show distinct maxima at 1975, 2033, 2075, and 2117 kev and show that  $\text{P}^{30}$  is formed in a resonant reaction rather than by a non-resonant process. The peaks in the gamma-ray yield curves at these energies can therefore be positively ascribed to resonant proton reaction with  $\text{Si}^{29}$ .

The yield of 0.511-Mev radiation after bombardment at energies between resonances is probably due to a combination of off-resonant reaction with  $\text{Si}^{29}$  and to the 66-sec positron activity of  $\text{F}^{17}$  following the non-resonant  $\text{O}^{16}(\text{p}, \gamma)\text{F}^{17}$  reaction. This hypothesis is supported by half life measurements of post-bombardment activity of the targets at two inter-resonance points at 1955 and 2007 kev. At those energies "half-lives" of the order of  $1.6 \pm .5$  min are observed with the count rate 2 min after bombardment at 1955 kev a factor of ten lower than the count rate 2 min after bombardment at 1975 kev.



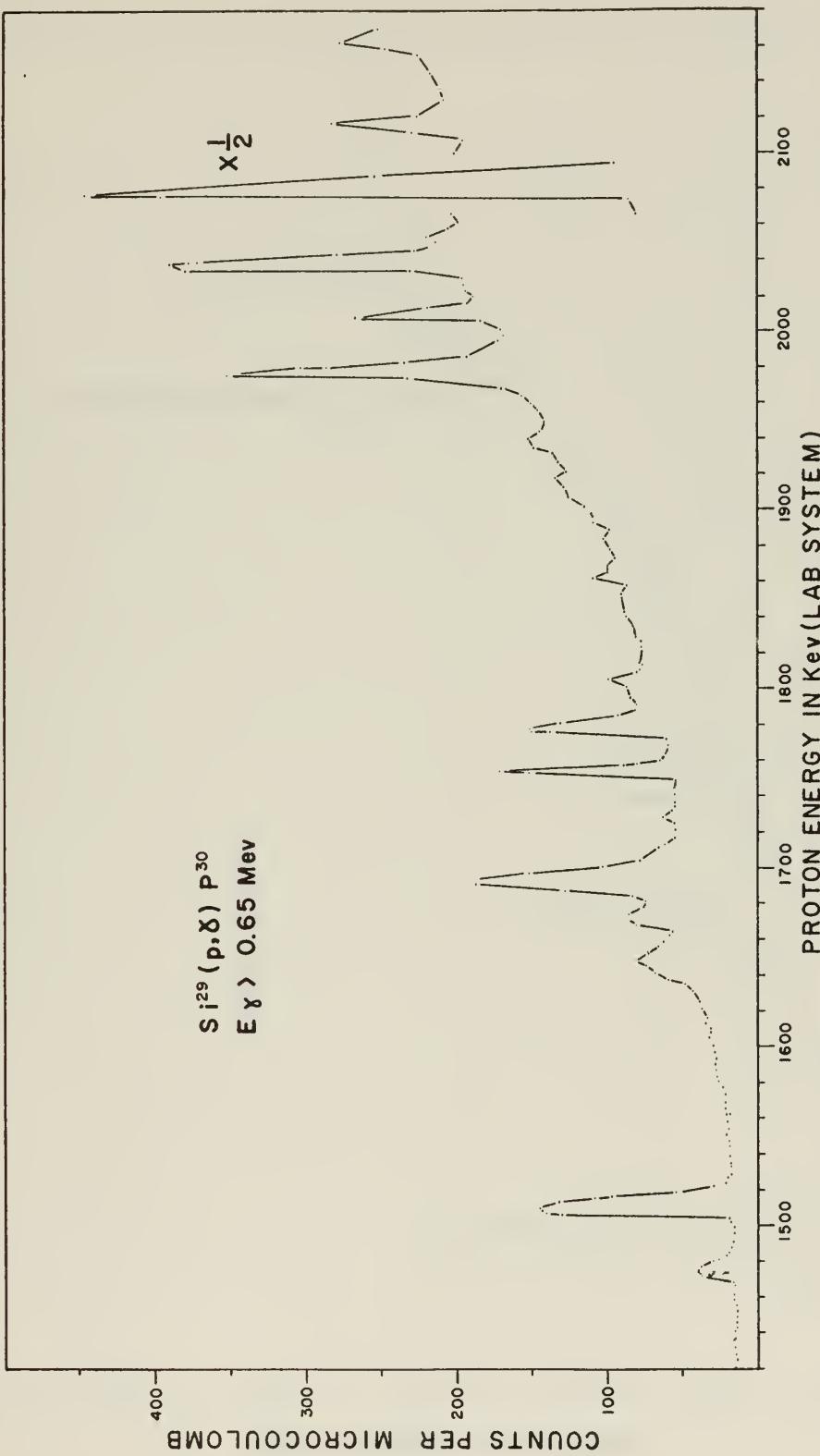


FIGURE 8. GAMMA-RAY YIELD



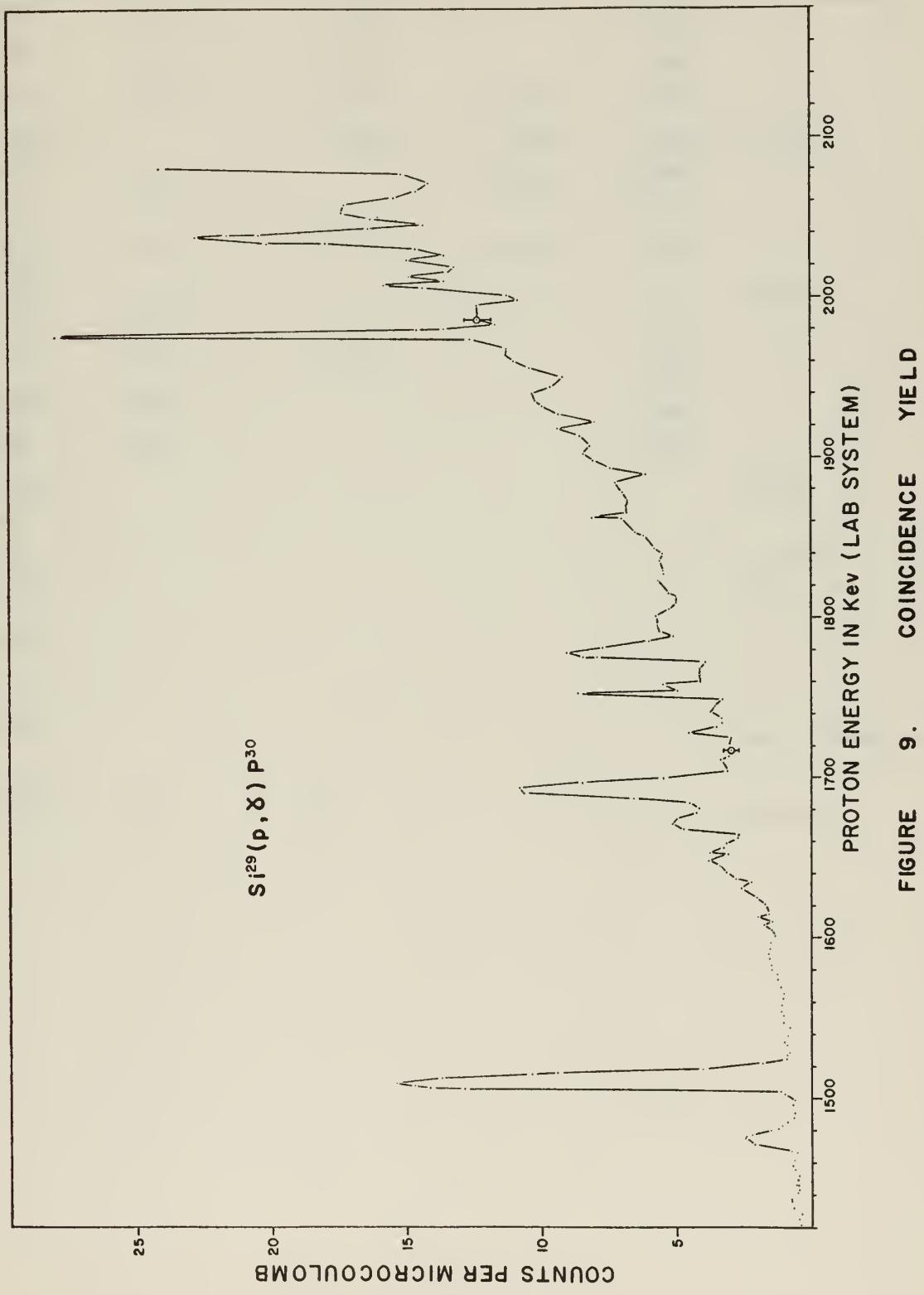




TABLE 2  
Resonances Observed in the  $\text{Si}^{29}(\text{p}, \gamma)\text{P}^{30}$  Reaction

$E_p$  from 1450 to 2170. Values in kev.

This work	Green <sup>11</sup>	Seagondollar <sup>12</sup>	Tsytko <sup>10</sup>	Wild <sup>15</sup>	Remarks
1470	1479	1479		1470	
1506	1515	1512		1505	
1637	1648	1643	1635	1638	
(1646)		(1653)	(1647)	(1650)	$\text{Si}^{28}$
1667	1671		1663	{1665 1670}	
1688	1692	1693	1680	1685	
(1726)					weak
1752	1752	1752		1749	
1775	1777	1775		1772	
(1803)	1811			1808	
1860	1857			1852	
(1933)					weak
1975					
(2007)					$\text{Si}^{30}$
2033					
2075					
2117					
(2159)					not identified
$\pm 0.20\%$	$\pm 2$	$\pm 2$		$\pm 3$	Errors



TABLE 3

Half Life of Post-bombardment Decay of P<sup>30</sup>

Bombardment Energy (kev)	Half life (min)
1975	2.42 $\pm$ .09
2033	2.45 $\pm$ .18
2075	2.51 $\pm$ .06
2117	2.47 $\pm$ .15



TABLE 4

## Intensity of Post-bombardment Annihilation Radiation

Bombarding Energy (kev)	Excitation* Energy in $P^{30}$ (Mev)	Relative Intensity
1506	7.04	19.3
1950	-	1.0
1975	7.49	4.1
1990	-	1.0
2033	7.55	7.9
2060	-	1.8
2075	7.59	12.4
2100	-	1.3
2117	7.63	8.1
2140	-	2.3

\*Based on  $Q = 5.582$ , reference 24.



## 2. Gamma-ray Spectra at 1975, 2033, 2075 and 2117 kev Resonances

Gamma-ray spectra were obtained at the four resonances at proton energies of 1975, 2033, 2075 and 2117 kev. Masking radiation from the resonant reaction with the contaminant F<sup>19</sup> dominates the upper energy region of these spectra so that a detailed analysis of the relatively weaker radiation from the Si<sup>29</sup>(p,γ)P<sup>30</sup> reaction was not possible. However, the gamma-ray energies and intensities which could be determined with reasonable certainty from stripping and least squares analysis for intensities are given in Table 5. The 6-7 Mev gamma rays due to F<sup>19</sup>, which are omitted in Table 5, were accounted for by inclusion in the least squares intensity analysis, since other background subtraction techniques did not work well at these energies as has been discussed in Chapter III. Also included in Table 5 are energies of sum-coincidence spectral peaks which were close to energies observed in the singles spectra. The good spectral resolution and discrimination against background radiation which are obtainable with the sum-coincidence method proved to be of only nominal assistance in analyzing the gamma-ray spectra. This was due in part to the relative weakness of radiative decay of these resonance levels in comparison with inelastic scattering, and in part due to photomultiplier gain drifts which blurred peaks in the sum-coincidence spectra. However, it is apparent from the large number of low intensity peaks observed in the sum-coincidence spectra at the higher-energy resonances that decay schemes from these levels are considerably more complex than decay from the 1752-kev resonance level,<sup>23</sup> where sum-coincidence spectra were recorded for the purpose of comparison.

The strong 1.28-Mev gamma ray observed at 1975, 2033, and 2075 kev is interpreted as deexcitation of the P<sup>30</sup> compound nucleus by proton reemission to the first excited state of Si<sup>29</sup> at 1.277 Mev.<sup>25</sup> The high intensity of this gamma ray is convincing evidence that this inelastic scattering process predominates over radiative capture and that the 128-Mev radiation is primarily responsible for the upper-energy peaks in the gamma-ray yield curve. Rather curiously, however, this gamma ray does not appear in the spectrum at the 2117-kev resonance.

The resonance at 2075 kev lies near the Si<sup>28</sup>(p,γ)P<sup>29</sup> resonance reported at 2090 kev which has a half width of 16 kev. The gamma-ray



TABLE 5

Gamma-ray Intensities at Resonance  
Energies of 1975, 2033, 2075, and 2117 kev

$E_p$ (kev)	Singles Spectra $E_{\gamma}$ (Mev)	Sum-Coincidence $E_{\gamma}$ (Mev)	Intensity (quanta/ $\mu$ C)
1975	4.52	4.58	5
	4.19	4.07	2
	3.48	3.40	2
	3.33		2
	3.00	2.99	3
	2.44	2.50	3
	1.28	-	117
	5.10	-	8.6
	1.28	-	187.3
	7.59	-	14
2033	4.70	4.85	60
	2.98	2.95	55
	2.26	2.23	55
	1.36	-	303
	1.28	-	1360
2117	5.10	-	2
	2.55	-	3

Error in singles spectra  $E_{\gamma}$  estimated to be 1%

Error in sum-coincidence  $E_{\gamma}$  estimated to be 2%

Error in intensities estimated to be 30%



TABLE 6

## Relative Gamma-ray Intensities at the 1470-kev Resonance

E $\gamma$ (Mev)	Transition	Observed* Intensity	Decay Scheme Intensity	Ref 23
7.00	R $\rightarrow$ 0	25.4	25.4	39.6
6.29	R $\rightarrow$ 2	9.6	9.6	6.3
4.50	19 $\rightarrow$ 0	14.6	14.3	13.2
4.14	13 $\rightarrow$ 0	10.7	10.7	11.1
3.98	R $\rightarrow$ 9	10.5	10.8	7.3
3.45	not accounted for	2.0	-	4.0
3.17	R $\rightarrow$ 11	3.0	1.5	2.0
3.17	11 $\rightarrow$ 1		1.5	2.0
3.07	R $\rightarrow$ 12	19.1	17.6	11.7
3.02	9 $\rightarrow$ 0		1.2	1.6
2.86	R $\rightarrow$ 13	13.7	13.7	13.0
2.69	13 $\rightarrow$ 3	2.8	2.8	-
2.50	R $\rightarrow$ 19	38.8	21.6	20.0
2.50	12 $\rightarrow$ 3		17.6	11.7
2.34	9 $\rightarrow$ 1	17.2	16.9	13.4
1.45	19 $\rightarrow$ 9	not analyzed <sup>+</sup>	7.3	6.8
1.45	3 $\rightarrow$ 0	"	20.3	11.7
.71	2 $\rightarrow$ 0	"	9.6	21.8
.68	1 $\rightarrow$ 0	"	18.4	

<sup>+</sup> An intensity of approximately 34 is indicated for the 1.45 Mev radiation by observations at 90° with .6" crystal face to target distance.

\* Relative intensities shown are also approximate absolute intensities in quanta per microcoulomb.



TABLE 7

## Double Angular Correlation Coefficients at 1470-kev Resonance

$E_\delta$ (Mev)	$P_2$ highest term		$P_4$ highest term		$A_2/A_0$	Std. Dev.	$A_4/A_0$	Std. Dev.
	$A_2/A_0$	Std. Dev.	$A_2/A_0$	Std. Dev.				
7.00	-.42	.02	-.43	.03			.05	.03
6.29	-.47	.05	-.47	.06			.01	.06
4.50	.09	.02	.11	.02			-.07	.03
4.14	-.09	.04	-.08	.04			-.01	.05
3.98	-.35	.05	-.37	.06			.02	.07
3.45	.03	.22	—	—			—	—
3.17	-.09	.16	—	—			—	—
3.07	-.04	.07	—	—			—	—
3.02								
2.86	.31	.03	.32	.04			-.03	.05
2.69	.39	.14	.50	.15			-.31	.19
2.50	-.22	.02	-.24	.02			.03	.02
2.34	-.24	.03	-.26	.03			.06	.04

Coefficients shown are corrected for finite angular resolution of detector.



spectrum at this resonance is characterized by a 4.76-Mev radiation to the ground state of  $P^{29}$ .<sup>36</sup> It seems probable, therefore, that the 4.70-Mev gamma ray seen in the spectrum at 2075 kev is due to the  $Si^{28}$  reaction. The observed intensity is in agreement with this explanation. Although the presence of 2.94-Mev and 2.24-Mev gamma rays in this spectrum suggest a cascade through level 8 at 2.94-Mev in  $P^{30}$ , the 1.45-Mev gamma ray known to occur in this cascade<sup>22</sup> is not seen and sum-coincidence work has not been able to confirm this possibility. It is possible, however, that the gamma rays due to  $Si^{29}$  reactions could be responsible for the anisotropy of the radiation from the 4.76-Mev level in  $P^{29}$  which has been observed by Newton,<sup>37</sup> since anisotropy is not expected to result from the spin assignment to this level of  $1/2+$  based on proton scattering experiments.<sup>38</sup>

The remaining weaker radiations seen at the four higher resonances are most probably due to radiative capture by  $Si^{29}$ , but the rather large errors, due chiefly to masking contaminant radiation, do not permit more than tentative postulation of decay modes. Gamma rays observed at the 1975-kev resonance suggest cascades through level 19 at 4.50 Mev and through level 14 at 4.18 Mev. No conclusions can be drawn for the 2033-kev resonance. At the 2075-kev resonance the 7.59-Mev radiation is readily recognized as a ground state transition but the other gamma rays are not identified. Finally, at the 2117-kev resonance, the two radiations observed nicely fit a decay through level 5 at 2.54 Mev which is known to decay 95% to the ground state.<sup>22</sup>

### 3. 1470-kev Resonance Decay and Intensities

The gamma-ray spectrum observed at the 1470-kev resonance with a crystal face to target distance of 0.6 in. and at an observation angle of  $90^\circ$  is shown in Fig. 10. This spectrum has been least squares fitted by gamma rays with energies required by the decay scheme proposed in reference 23. Reanalyses with alternate energy choices in each instance gave a poorer fit in reassurance of the correctness of the decay modes proposed. This decay scheme, with intensities as found in the present analysis, is shown in Fig. 11. (Heavy dots indicate cascades confirmed in reference 23 by coincidence spectroscopy.) The branching ratio shown for the 4.50-Mev level is that given in reference 22. The assumed correctness of this ratio then requires branching ratios of 93:7 for level 9 and 80:20 for level 13 in order to fit the intensities of gamma rays which have been



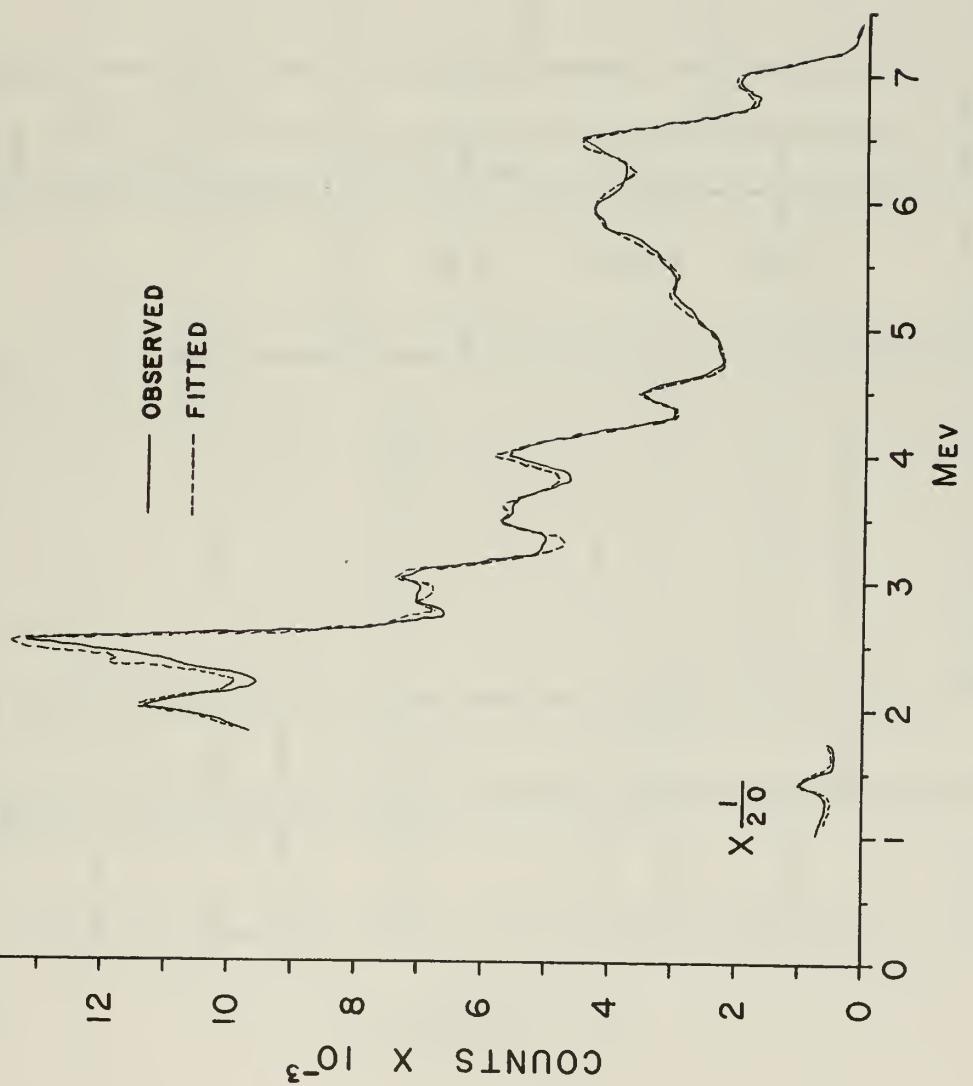


FIGURE 10. GAMMA-RAY SPECTRUM AT 1470 KEV



DECAY SCHEME AT 1470 KEV RESONANCE

$E_\gamma$ (MEV)

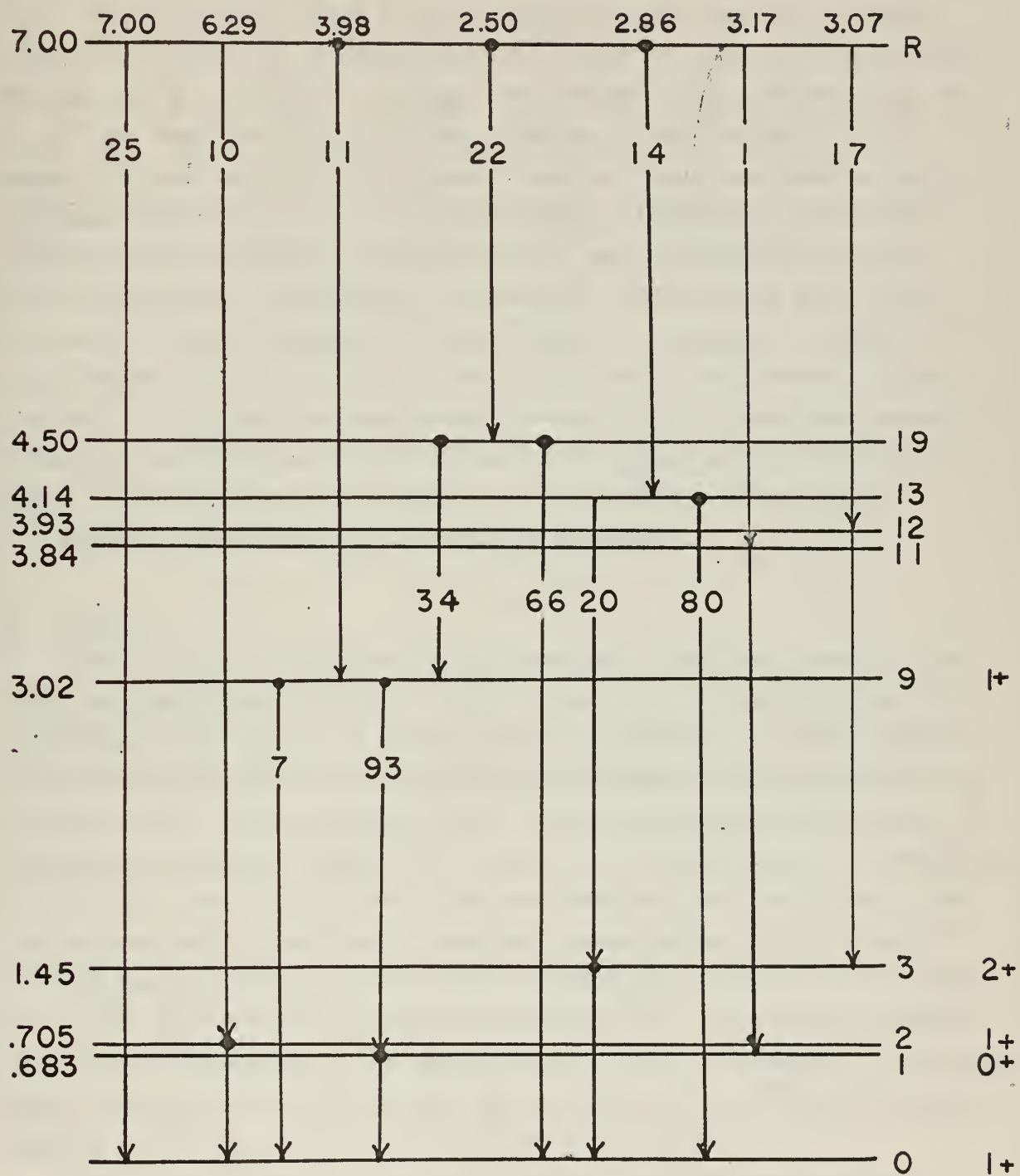


FIGURE 11.



observed. These ratios are in reasonably good agreement with ratios of 89:11 reported for level 9<sup>22</sup> and 88:12 reported for level 13<sup>23</sup>.

Intensities of gamma rays observed in the angular correlation runs are shown in Table 6 along with intensities observed in reference 23 at 55°. The intensities found in the correlation runs have been averaged over  $4\pi$  steradians by weighting the mean values for each observation angle by the sine of that angle. Because the 1.46-Mev background radiation due to K<sup>40</sup> was found to be quite strong relative to the 1.45-Mev gamma ray observed in the correlation runs, no attempt was made to analyze the gamma radiation of this or lower energy. Although the background radiation subtracts out, the statistics of the remaining counts do not allow extraction of significant information. The observed intensities are seen to agree reasonably well with those of reference 23 except for the 7.00-Mev radiation which is found to be considerably weaker in the present work. Since the experimental intensities are based upon reported values of intrinsic peak efficiencies<sup>56</sup> rather than an absolute spectrometer calibration, this discrepancy is not considered significant.

#### 4. Angular Correlations at the 1470-kev Resonance

##### A. General

The correlation function coefficients which have been found in this experiment are given in Table 7. The values given are weighted means of the values found in each of the five sets of data taken. The chi-square values for each correlation coefficient have been calculated and show the different sets to be consistent, under the acceptance requirement that the cumulative chi-square probability be equal to or less than 0.9. Combining intensities from each data set, then analyzing for the coefficients gives values practically identical to the mean values quoted. This agreement to some degree serves as a check that computational errors have not crept in. Each experimentally observed spectrum has also been plotted together with the corresponding fitted spectrum, then visually examined to determine that a reasonable fit was obtained in each case. Fig. 12 shows a typical example of the fit obtained for the correlation runs.

The intensity analysis was performed both subtracting background radiation analytically and subtracting off-resonance background, with essentially the same results being obtained in each case. The standard



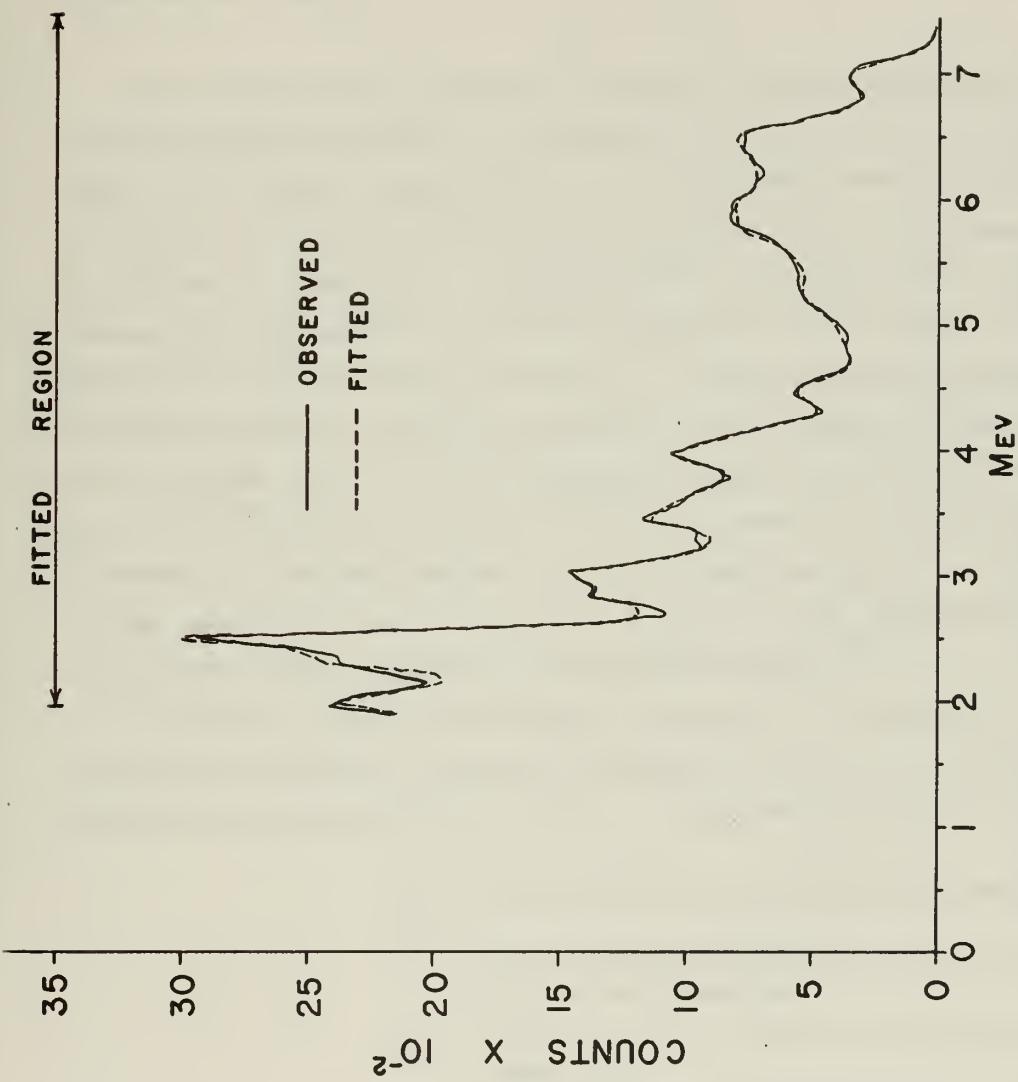


FIGURE 12. GAMMA RAY SPECTRUM FOR CORRELATION RUN 3-5 AT 90°



deviations in the mean coefficient values have been calculated from

$$1/\epsilon^2 = \sum_{i=1}^n 1/\epsilon_i^2$$

and have been compared with the values given by

$$\frac{1}{n(n-1)} \sum_{i=1}^n (x_i - \bar{x})^2$$

which agree closely.

The most obvious feature of the data is the anisotropy of the 4.50-Mev gamma ray which is incompatible with a spin of zero for this level. It is also noteworthy that the 4.50-Mev gamma ray is the only radiation to show a statistically significant  $P_4(\cos \theta)$  term (notation is defined in Appendix 1). Fig. 13 shows the correlation coefficient values for this gamma ray obtained from each set of data. In this figure M denotes mean values and C denotes values obtained by combining intensities from each data set before computing coefficients. Fig. 14 shows the fit to the Legendre series both with and without  $P_4$  terms. Although the existence of the  $P_4$  term is not incontrovertible it permits a markedly better fit to the data and gives a smaller value of the parameter  $\epsilon$  which serves as a guide to the appropriate number of terms which should appear in the Legendre series.<sup>41</sup>

In order to deduce spin-parity values it is necessary to examine the possible particle and radiation mixing ratios for this reaction. The notation used herein to refer to these mixing ratios is as follows:

$\delta_1$  = Proton-partial-wave mixing parameter

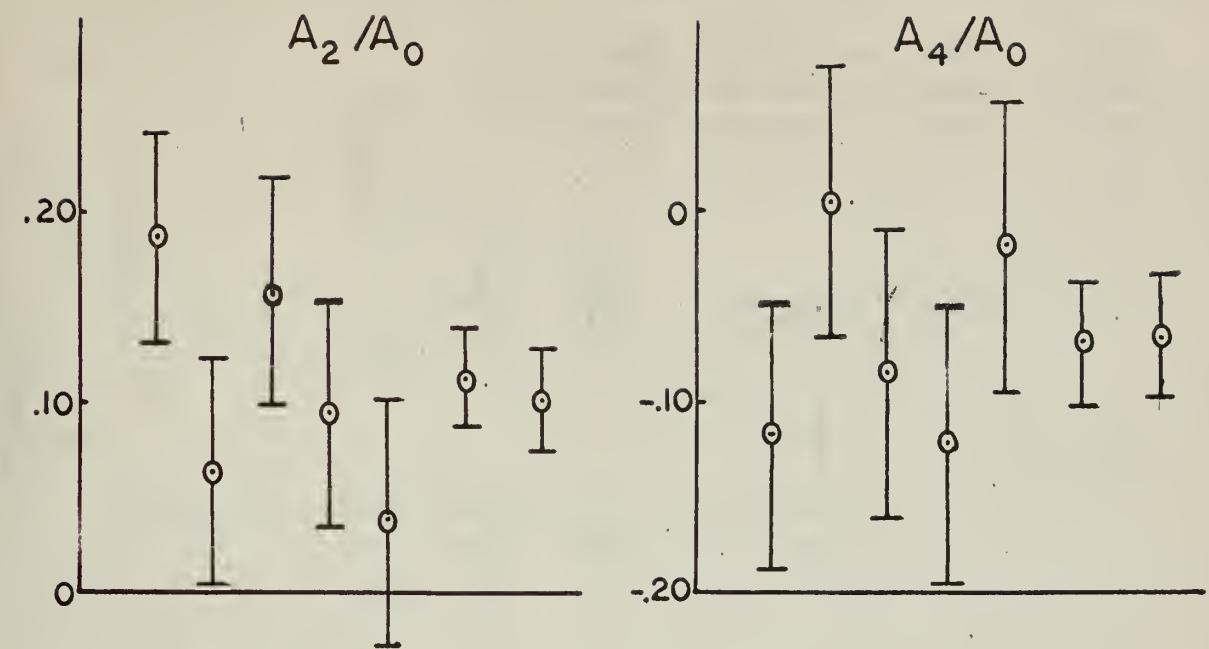
$\delta_2$  = The multipole mixing ratio of an intermediate unobserved gamma ray

$\delta_3$  = The multipole mixing ratio of an observed gamma ray

$\delta_4$  = Channel-spin mixing parameter

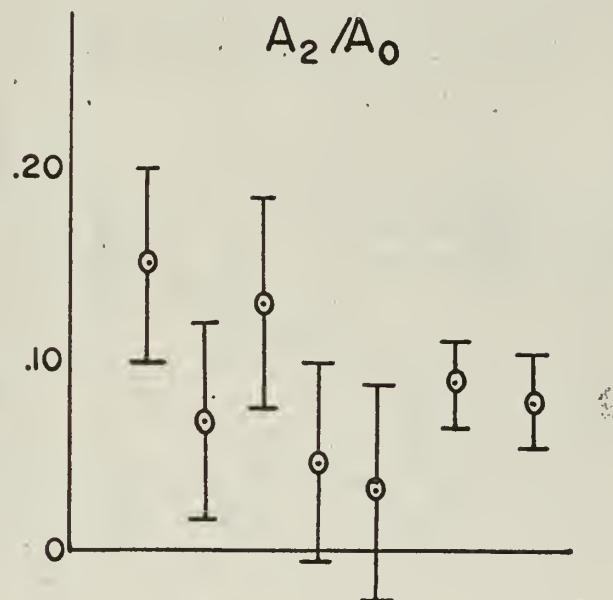
Appendix 1 defines these parameters more fully and describes how they enter into the expression for the theoretical correlation function. For the  $Si^{29}(p, \gamma)P^{30}$  and  $Si^{29}(p, x\gamma)P^{30}$  (intermediate unobserved) reactions,





$P_4$  HIGHEST TERM  $\epsilon = .74$

RUN 1 2 3 4 5 M C



$P_2$  HIGHEST TERM  $\epsilon = 1.30$

FIGURE 13. DISTRIBUTION COEFFICIENTS FOR THE  
4.50 MEV GAMMA RAY



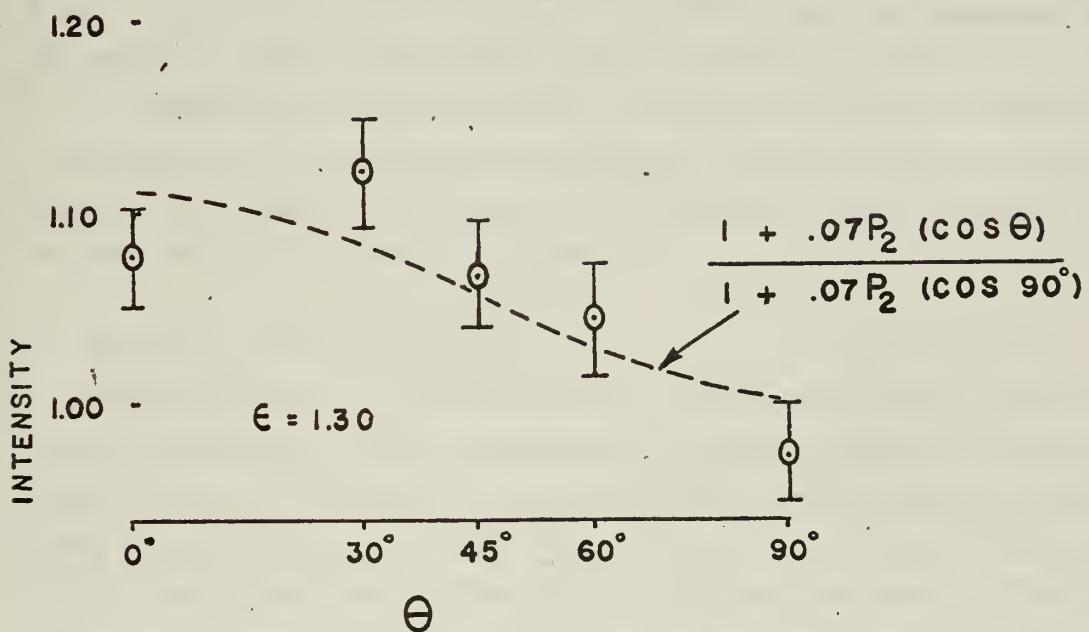
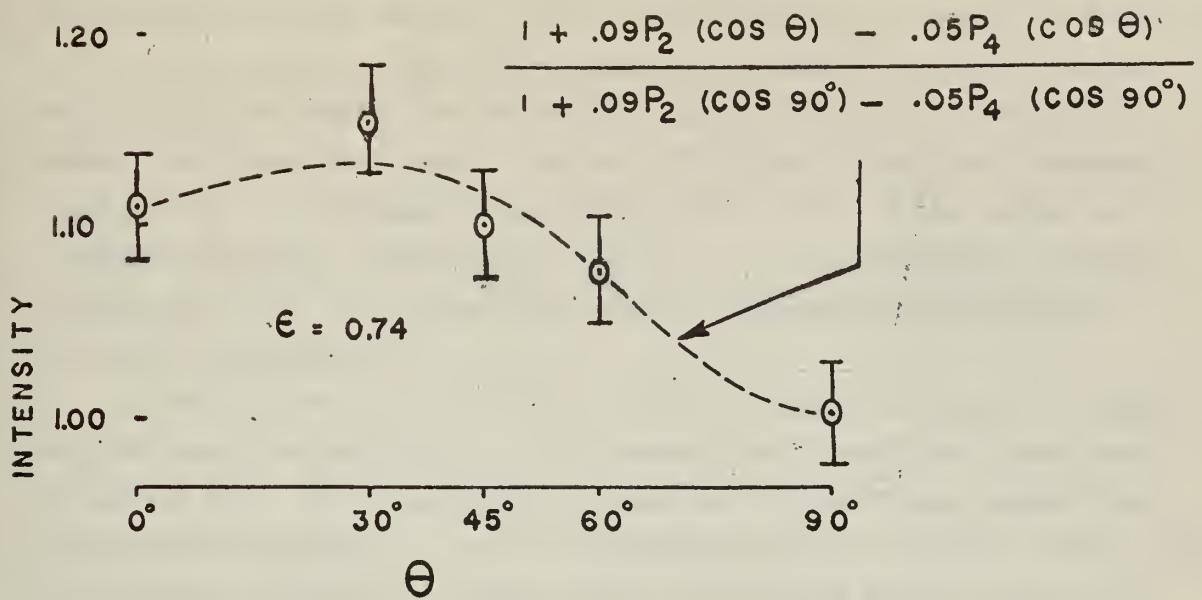


FIGURE 14. FITS OF COMBINED EXPERIMENTAL DATA TO LEGENDRE SERIES - 4.50-MEV GAMMA RAY (UNCORRECTED FOR FINITE SOLID ANGLE)



the correlation coefficients  $A_2/A_0$  and  $A_4/A_0$  are describable as functions of a maximum of three mixing ratios. Level curves, or contours, of constant  $A_i/A_0$  values may be plotted against any two of the mixing ratios. Actually the arctangents of the mixing ratios are used as variables for convenience since they remain finite. If a third mixing ratio appears, changing its value changes these contour plots only to the extent of changing the  $A_i/A_0$  values for all contours by a multiplicative constant. Contour plots for this analysis have been programmed and plotted as described in Appendix 2.

In the following discussion it will be assumed that only E1 radiation or E2/M1 mixtures are likely to be observed for transitions from levels of several Mev. This simplifying assumption is based upon transition probabilities estimated from the independent particle model<sup>42</sup> and is commonly made in analysis of gamma radiation following proton capture in light and intermediate nuclei. Although suppression or inhibition of dipole radiation is possible because of the isotopic spin selection rule,  $\Delta T = \pm 1$  for self conjugate nuclei,<sup>43</sup> there are an adequate number of lower levels of both  $T = 0$  and  $T = 1$  in  $P^{30}$  so that the resonance level can decay by either E1 or E2/M1 radiation without violating this rule.

In analyzing the experimental correlation coefficient values, tables from reference 23 have been quite helpful. This reference tabulates the maximum and minimum values which may be assumed by  $A_2/A_0$  and  $A_4/A_0$  for a variety of postulated spin sequences.

#### B. Resonance level spin-parity

The observed  $A_2/A_0$  terms of about -0.40 for the 7.00, 6.29, and 3.98 Mev gamma rays, which all proceed to known  $1^+$  levels, are found to be most useful in determining the spin-parity of the resonance level. Among possible values through 4 for this level, it is found that only  $1^+$ ,  $2^+$ , and  $2^-$  can lead to the  $A_2/A_0$  coefficient measured for the 7.00-Mev ground state transition. The same conclusion is supported by the  $A_2/A_0$  values observed for the 6.29 and 3.98 Mev radiations.

If one inspects the contour plot, Fig. 15, for the  $1^+ \rightarrow 1^+$  transition it is seen that a d-s proton-angular-momentum mixing ratio of over 0.90 is required to produce the observed  $A_2/A_0$  value for the 7.00-Mev gamma ray. Since the relative penetrability of d waves to s waves is about  $3 \times 10^{-2}$ , such a mixture is not a likely choice. The  $A_4/A_0$  term observed



for the 4.50-Mev radiation supports this conclusion since  $A_4/A_0$  terms do not result from resonance spins smaller than 2. Finally, the value of  $A_2/A_0$  observed for the 2.34-Mev radiation will be shown to be incompatible with a 1+ resonance level spin. This radiation proceeds to the known 0+ level 1, at 0.68 Mev, from the known 1+ level 9, at 3.02 Mev. Level 9 is fed more than 50% from the resonance level and, secondarily, by the intermediate 1.48-Mev radiation from level 19. The  $A_2/A_0$  value observed for this primary 3.98-Mev feeder restricts the ranges of  $\delta_1$  and  $\delta_2$  for the  $R \rightarrow 9 \rightarrow 1$  transition. When these ranges are delineated on a contour plot for a  $1+ \rightarrow 1+ \rightarrow 0+$  transition, it is found that  $A_2/A_0$  for the second radiation is restricted to a range between -0.01 and +0.24. By an extension<sup>48</sup> of the formalism for intermediate unobserved cascades given in Appendix 1 it is possible to calculate the angular distribution coefficients for the  $9 \rightarrow 1$  radiation which result from the  $R \rightarrow 19 \rightarrow 9 \rightarrow 1$  cascade. It is found that, for a 1+ resonance level, the minimum possible value of  $A_2/A_0$  occurs for an assumed spin of 2 for level 19. This minimum value has been calculated to be -0.175. Thus a 1+ resonance level leads to a minimum possible value for  $A_2/A_0$  of about -0.093. Comparing this with the observed value of  $-0.24 \pm .03$  for the 2.34-Mev gamma ray, 1+ can be ruled out as a possible spin-parity of the resonance level.

Having narrowed the possible resonance spin values to 2+ or 2-, the contour plots may be examined for the  $2+ \rightarrow 1+$  and  $2- \rightarrow 1+$  transitions, applicable to the 7.00-Mev radiation to the ground state. In the first case, channel-spin mixing may occur while in the second case proton-partial-wave mixing is permitted. For the  $2+ \rightarrow 1+$  case Fig. 16a shows the contour plot with the experimental value of  $A_2/A_0$  for the 7.00-Mev gamma ray delineated thereon by cross-hatching. Fig. 16b is a similar plot for  $A_4/A_0$ . The darker shading in Fig. 16b indicates the superposition of cross-hatching on both plots and defines the ranges and combinations of mixing parameters which can give rise to the observed coefficient values. For  $\delta_3 \approx 0$  (i.e. M1 radiation),  $\delta_4$  is constrained to be less than  $\approx \tan 50^\circ$ . This must not be taken too seriously, however, since the  $A_4/A_0$  contours are widely spaced at the top and bottom boundaries of the permitted region, so that a very slight increase in the error for the experimental  $A_4/A_0$  value would allow  $\delta_4$  to range up to  $90^\circ$ . The small region at  $\delta_3 \approx \tan(-67^\circ)$  and  $\delta_4 \approx \tan 49^\circ$  also fits the observed corre-



lation. The plots for the  $2- \rightarrow 1+$  case, shown in Figs. 17a and 17b, are quite similar to the  $2+ \rightarrow 1+$  case and show that there is no restriction on f-p proton-wave mixing for  $\delta_3 \approx 0$  (E1 radiation). The two small permitted regions at  $\delta_3 \approx \tan 67^\circ$  may be ruled out since they require an M2/E1 mixture. It is seen that an unlikely but small M2/E1 mixture with  $\delta_3 \approx 0.026$  or greater is required for this transition for the case of resonance formation by pure p waves. As will be shown, the  $2-$  resonance level assignment is also less probable than  $2+$  based on mixing ratios which are required by experimentally observed coefficients for the 4.50-Mev radiation.

As a result of the above considerations, a spin-parity of  $2+$  is proposed for the resonance level with  $2-$  as a less likely possibility.

#### C. $R \rightarrow 13 \rightarrow 0$ cascade

There are only two cascades which allow cross checking of mixing parameters as a method of eliminating possible spin sequences. These are the cascade through level 19 and that through level 13 both of which levels have undetermined spins. The latter cascade will be considered first. The existence of an  $A_4/A_0$  term which has been observed in the distribution of the 4.14-Mev radiation from level 13 at the 1686-kev resonance<sup>23</sup> is evidence that level 13 must have a spin greater than or equal to 2. Of the likely possible spins for this level,  $2\pm$ ,  $3\pm$ , and  $4-$ , only  $2+$  and  $2-$  can result in the negative value of  $A_2/A_0$  observed for the 4.14-Mev gamma ray. An examination of the contour plots for the four sequences  $2\pm \rightarrow 2\pm \rightarrow 1+$  for the  $R \rightarrow 13 \rightarrow 0$  cascade shows compatibility in all cases with the observed coefficients for the 2.86-Mev gamma ray and the 4.14-Mev gamma ray. For the sequence  $2+ \rightarrow 2- \rightarrow 1+$  the parameter  $\delta_4$  is constrained to be larger than  $\tan 70^\circ$  but the other possible sequences are compatible with a full range of values for either  $\delta_1$  or  $\delta_4$  as the case may be. The only finding is that, for all cases, the  $R \rightarrow 13$  transition must be relatively pure dipole radiation. The observed coefficient values for the weak 2.69-Mev gamma ray, which also is believed to deexcite level 13, are also compatible with all four possible spin sequences with reasonable choices for  $\delta_3$ .

#### D. $R \rightarrow 19 \rightarrow 0$ cascade

Since no requirements on parameters applicable to this cascade are imposed by transitions already considered it becomes necessary to consider



the  $R \rightarrow 19 \rightarrow 0$  sequence with no prior restrictions on the deltas involved. We find, furthermore, that no information is forthcoming from the  $R \rightarrow 19$ , 2.50-Mev gamma-ray correlation coefficient since this radiation is indistinguishable from the  $12 \rightarrow 3$ , 2.48-Mev gamma ray.

A survey of maxima and minima of  $A_2/A_0$  coefficients shows, if we allow 2- as a possible resonance level spin, that the following sequences for the  $R \rightarrow 19 \rightarrow 0$  transition are consistent with the experimental coefficients for the 4.50-Mev radiation:

2+ 1+ 1+	2+ 1- 1+
2- 1+ 1+	2- 1- 1+
2+ 2+ 1+	2- 2- 1+
2- 2+ 1+	2- 3- 1+
2+ 3+ 1+	2+ 3- 1+

It has been found in the analysis that a better fit (see Fig. 14) to the data is given by the series including a  $P_4$  term. Then, provided this term is not spurious, which is not to be excluded as a possibility, the list can be reduced by eliminating sequences not giving rise to  $A_4/A_0$  terms.

The following possibilities then remain:

2+ 2+ 1+
2- 2+ 1+
2+ 3+ 1+
2+ 3- 1+
2- 3- 1+

The last two transitions shown require E3/M2 radiation so, adhering to the hypothesized E1 and E2/M1 radiative transitions as practical candidates, the spin of the 4.50-Mev level is limited to either 2+ or 3+. Of the three remaining possibilities, the  $2- \rightarrow 2+ \rightarrow 1+$  sequence is perhaps least likely. Figs. 19a and 19b show that this sequence for  $R \rightarrow 19 \rightarrow 0$  requires an f to p wave mixing parameter of over 0.10 with a large E2/M1 mixing ratio of nearly 12.0. For a  $2+ \rightarrow 3+ \rightarrow 1+$  sequence, Figs. 20a and 20b show that an E2/M1 ratio of over 1.7 is required for the 4.50-Mev



radiation in order to fit the data. Because large E2/M1 mixing parameters are not commonly seen in the lighter nuclei, this sequence, though more likely than  $2+ \rightarrow 2+ \rightarrow 1+$ , does not appear as probable as a  $2+ \rightarrow 2+ \rightarrow 1+$  sequence, as will be seen. For the  $2+ \rightarrow 2+ \rightarrow 1+$  sequence Figs. 18a and 18b show that for an assumed channel-spin mixing ratio of zero, an E2/M1 mixing ratio of 0.27 will fit the observed coefficients for the 4.50-Mev gamma ray with a considerable latitude permitted for  $\delta_2$ . Alternate choices for  $\delta_4$  will also permit acceptable choices of mixing parameters for this sequence though in general  $\delta_2$  and  $\delta_3$  will be required to be larger.

Summarizing, level 19 is found to have a spin of 2+ or 3+ under the assumptions which have been made, with 2+ being more probable.

#### E. Comparison with other results

Concurrently with the present work, an investigation of triple angular correlations at the 1470-kev resonance has been undertaken at Wright-Patterson Aeronautical Laboratories.<sup>40</sup> The angular distribution of the 7.00-Mev gamma ray was also examined in this experiment. An  $A_2/A_0$  value of  $-0.38 \pm .02$  was observed for this gamma ray in good agreement with the value shown in Table 7. Preliminary results of this experiment have indicated restriction of the resonance level spin to a value of 1 or 2 and show that a zero spin for the 4.50-Mev level is not compatible with such a restriction.<sup>62</sup> The preliminary analysis of the triple-correlation experiment is seen to be in agreement with the present work. It is to be hoped that the complete analysis of this work which is now in progress, will lead to a firmer indication of the spin-parity of the 4.50-Mev level since the 2+ (3+) value proposed herein depends upon the somewhat debatable  $A_4/A_0$  term observed and upon the hypothesis that only E1 or E2/M1 radiation is present.

#### F. T = 1 character of level 19

The prediction of 0+ for the spin-parity of level 19 is based on the premise that this level is the T = 1 analog of the 3.80-Mev level of Si<sup>30</sup> in the isotopic spin multiplet formalism.<sup>42</sup> A spin of 0 has been suggested for the 3.80-Mev level in Si<sup>30</sup> based on observations of cascade gamma rays following the Si<sup>30</sup>(p,p')Si<sup>30</sup> reaction.<sup>63</sup> A spin of 0+ results also from



calculations made by Thankappan and Pandya of collective vibrational states in a weakly-coupled collective model of the nucleus.<sup>6</sup> The spin-parity of level 19 in P<sup>30</sup> is thus expected both on experimental and theoretical grounds. The finding of non 0+ spin-parity for this level then appears to be good evidence that either the level is complex (e.g. a doublet) or that it is not the isotopic spin multiplet analog of the 4.80-Mev state in Si<sup>30</sup> as has been accepted. Some evidence has been reported that the 3.80-Mev level in Si<sup>30</sup> is a doublet, based upon observations of protons from the reactions Si<sup>28</sup>(t,p)Si<sup>30</sup> and Si<sup>30</sup>(p,p')Si<sup>30</sup>.<sup>59,44</sup> It is possible, therefore, that the analogous 4.50-Mev level in P<sup>30</sup> is a doublet, although no other evidence so suggests. The other possible explanation is that a nearby level in P<sup>30</sup>, such as that at 4.42 Mev, may be the T = 1 level rather than the 4.50 Mev level. This does not appear unlikely, since the T = 1 character attributed to the 4.50-Mev level is a probable rather than a firm assignment which was deduced from the investigation of the S<sup>32</sup>(d, $\alpha$ )P<sup>30</sup> reaction by Endt and Paris.<sup>18</sup> It is based upon a correction of the 3.80-Mev level in Si<sup>30</sup> for the Coulomb energy and neutron-proton mass difference which yields a predicted energy of 4.47 Mev for the excitation energy in the P<sup>30</sup> nuclide. The assignment is strengthened by the fact that the alpha-particle group from S<sup>32</sup>(d, $\alpha$ )P<sup>30</sup> corresponding to a residual excitation in P<sup>30</sup> of 4.501± .010 Mev has only about half the intensity of neighboring groups. This has been interpreted as being due to operation of the selection rule conserving isotopic spin in this reaction.

On the basis of the energy predicted for the T = 1 level in P<sup>30</sup> the 4.42-Mev level is seen to be nearly as likely a candidate as the 4.50-Mev level, particularly since the energy prediction is not expected to be exact. Similarly in the T conservation argument, the observed lower intensity of the alpha group for the 4.50-Mev level can hardly be considered conclusive evidence of the T = 1 character of the level, since some T = 0 levels are associated with alpha-particle groups of even lower intensity (e.g. the 3.84-Mev level). Unfortunately, the 4.42-Mev level in P<sup>30</sup> is not known to be excited in radiative capture, so that spin-parity of this level will probably have to be determined from a different reaction.

##### 5. Summary of Results

Excitation curves for the Si<sup>29</sup>(p, $\gamma$ )P<sup>30</sup> reaction have been obtained over the proton energy region from 1420 through 2160 kev. Four previously unreported



resonances have been detected and confirmed at proton energies of 1975, 2033, 2075, and 2117 kev with experimental uncertainties of  $\pm 4$  kev. Gamma-ray spectra at these resonances show that inelastic scattering predominates over radiative capture, except at 2117 kev where no inelastic scattering is evident. Some information concerning the radiative decay schemes at the 1975, 2075, and 2117 kev resonances was obtained by spectral analysis with the assistance of sum-coincidence spectra.

The gamma-ray spectrum at the 1470-kev resonance has been analyzed and found to be in agreement with the decay scheme proposed by Moore.<sup>23</sup> The angular distributions of gamma rays observed at this resonance have been measured in a series of five experimental observations, each consisting of gamma-ray spectra obtained at five angles relative to the proton beam. Analysis of this data, under the assumption that E1 and E2/M1 radiation deexcite  $P^{30}$  energy levels of several Mev, has shown that the 1470-kev resonance level has a spin-parity of 2+ or, less probably, 2-. The spin-parity of level 13 at 4.14 Mev has also been found to be limited to 2+ or 2-. Level 19 at 4.50 Mev was found not to have a spin-parity of 0+ which had been predicted.<sup>6</sup> Under the above-mentioned assumption, an apparent  $A_4/A_0$  term observed in the angular distribution of this radiation leads to the conclusion that the spin-parity for level 19 is 2+ or, less probably, 3+.



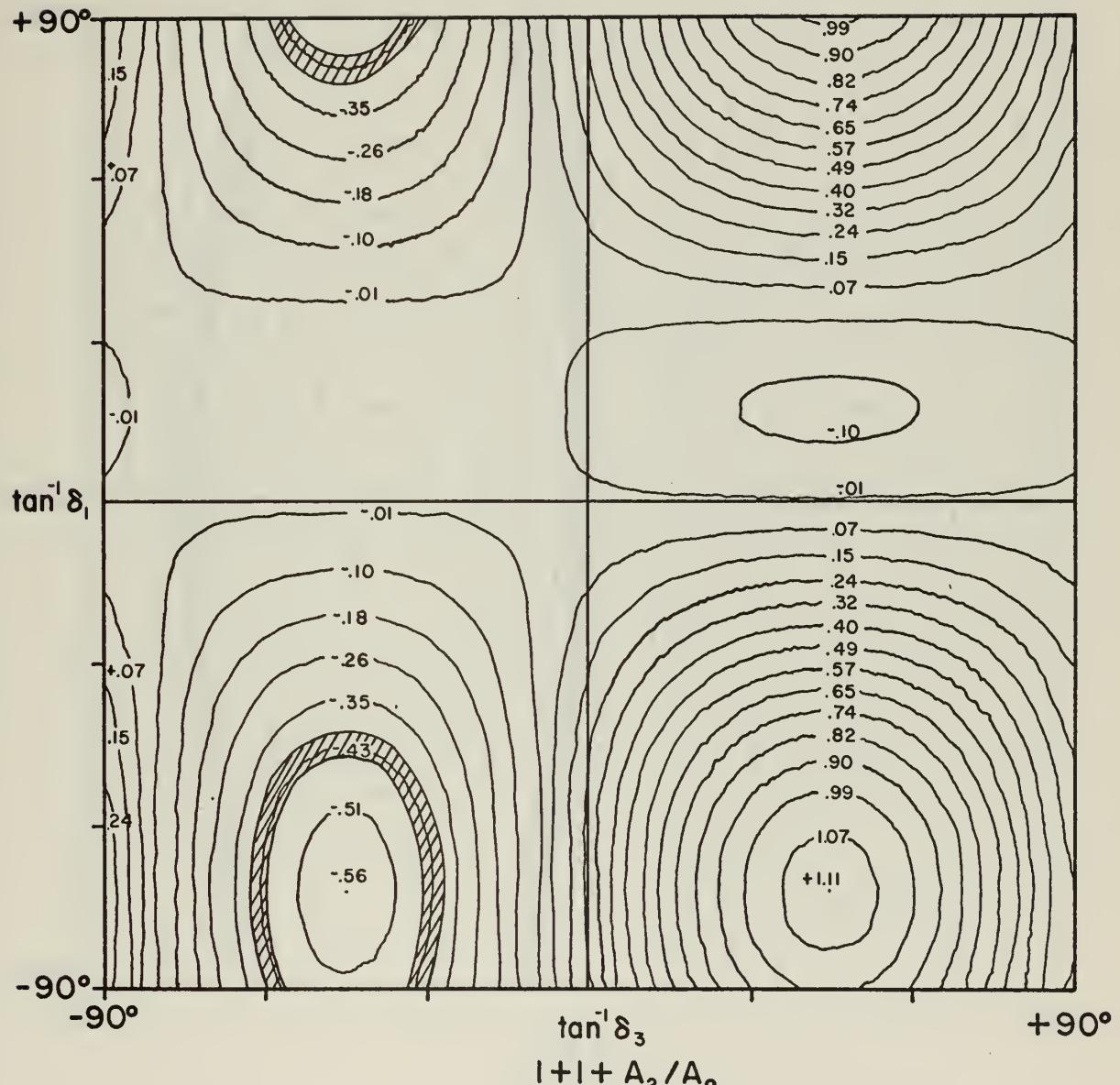
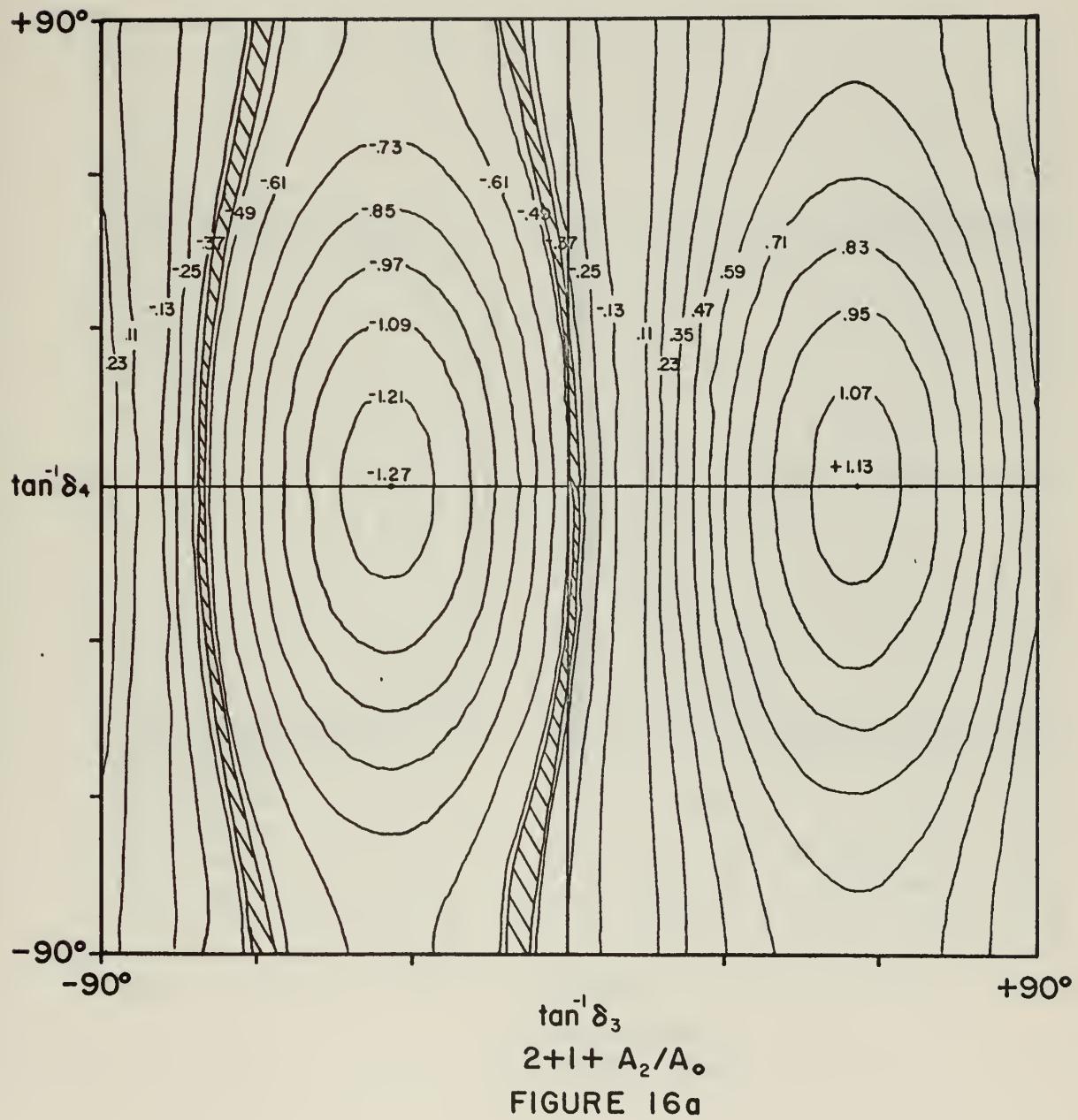
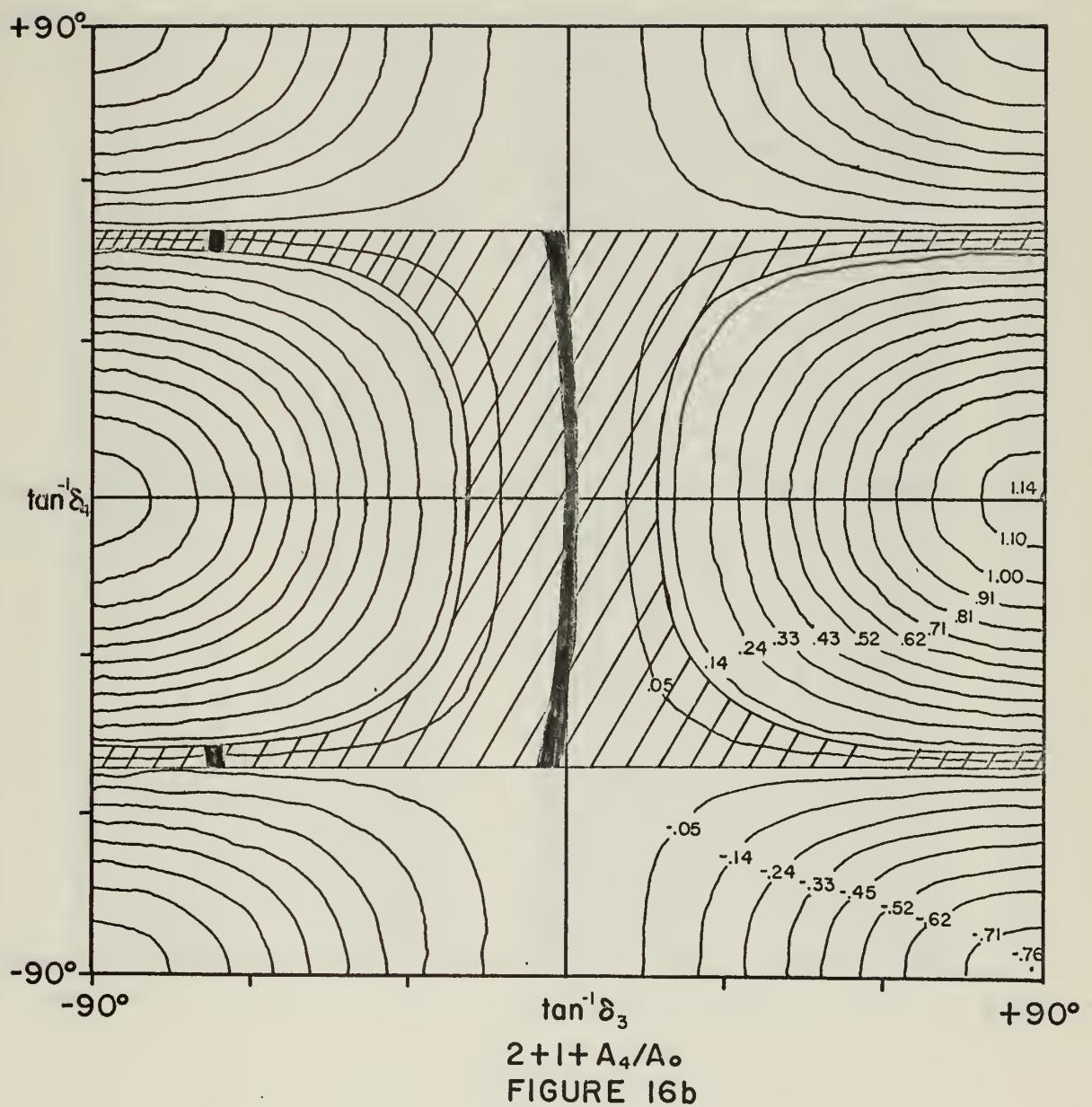


FIGURE 15

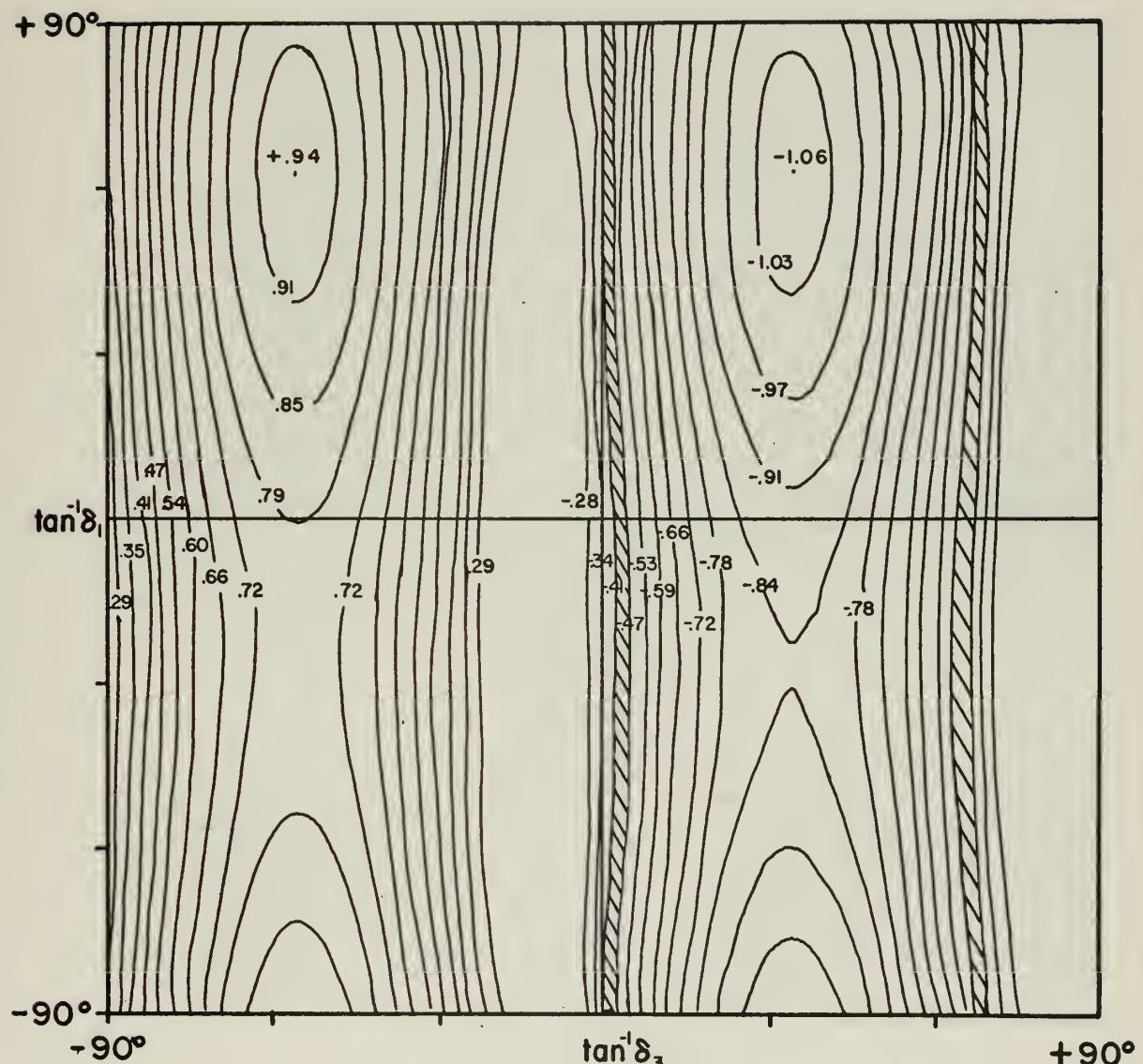






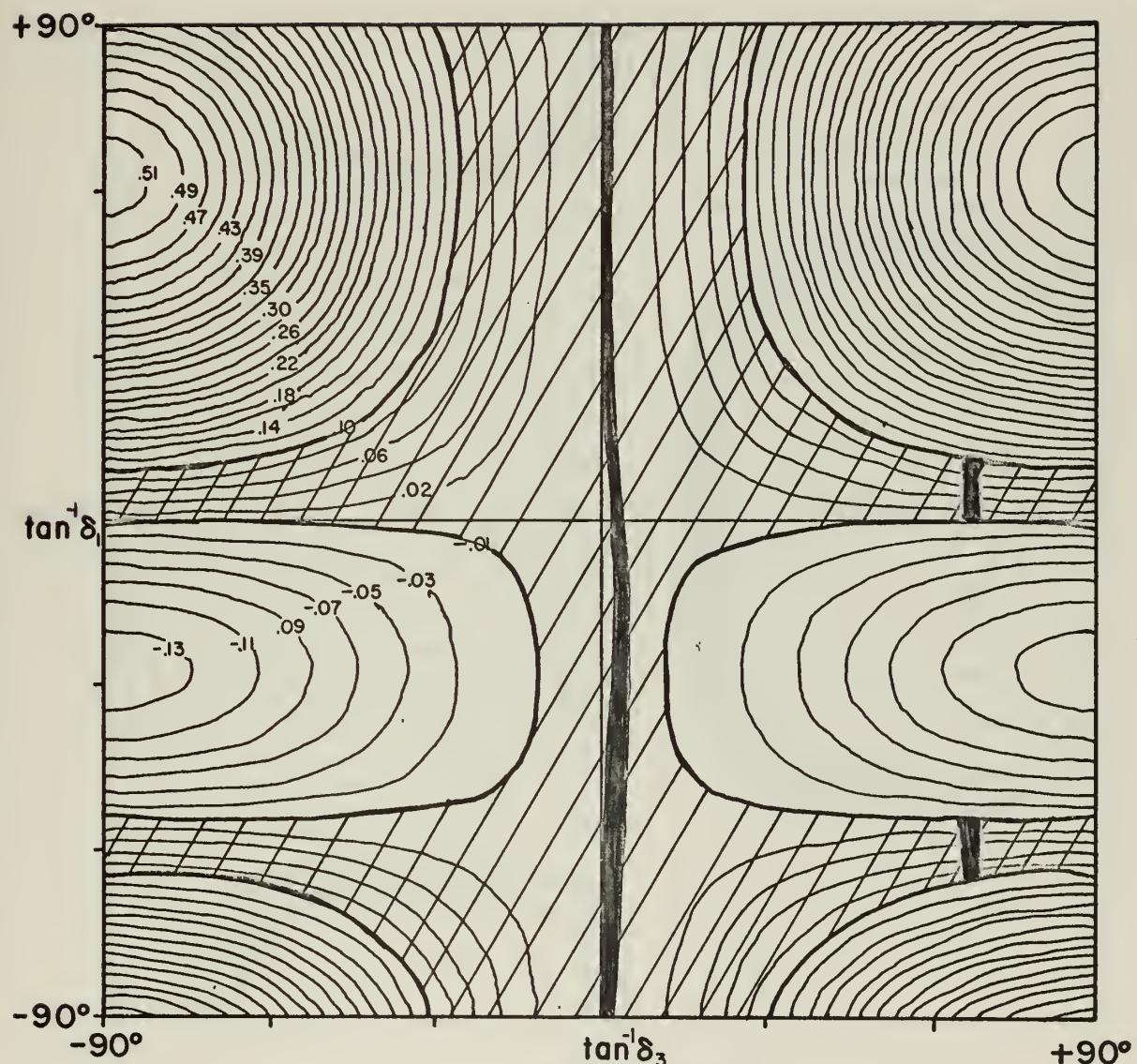






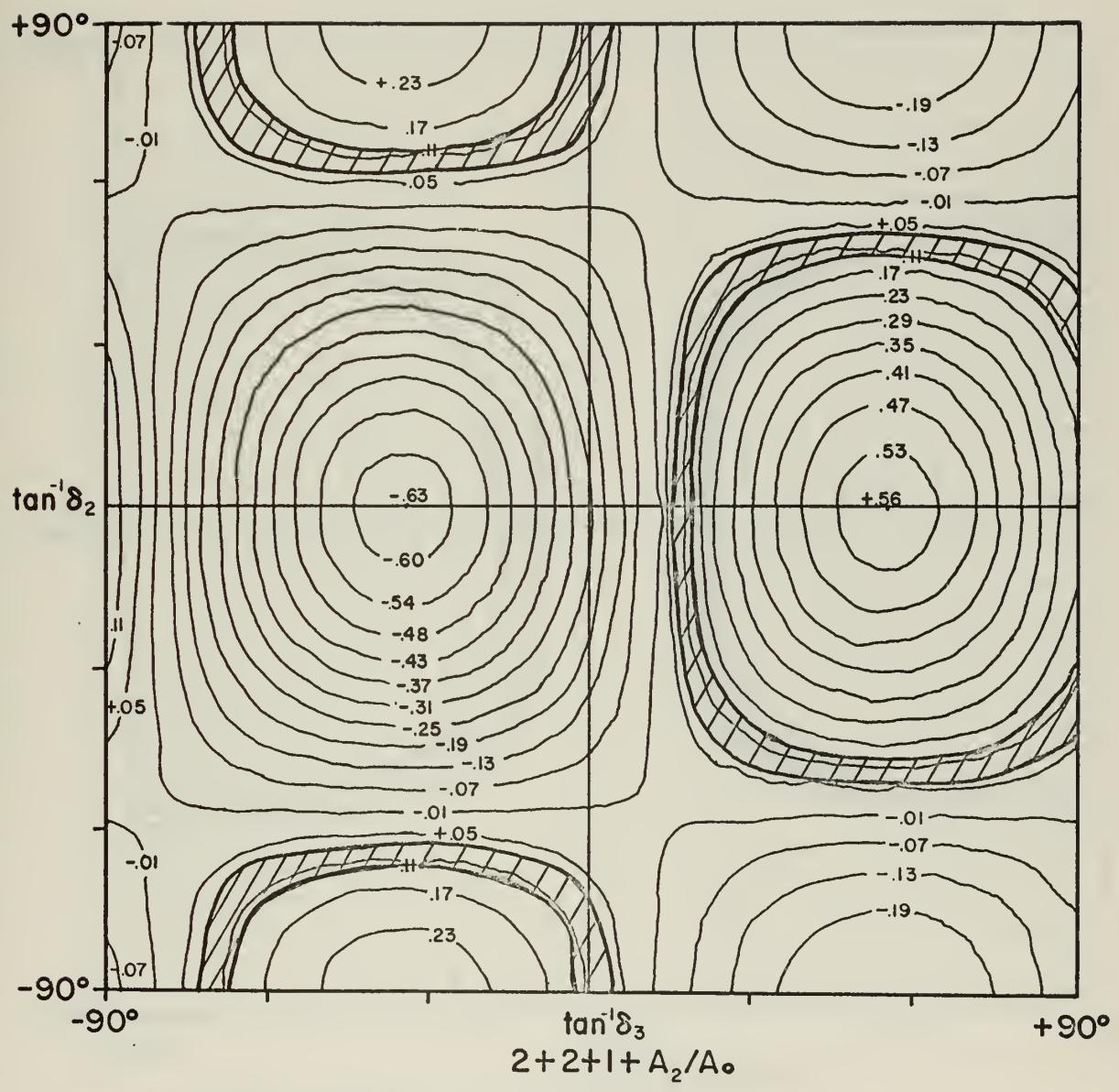
$2 - 1 + A_2/A_0$   
FIGURE 17a





$2 - 1 + A_4/A_0$   
FIGURE 17b

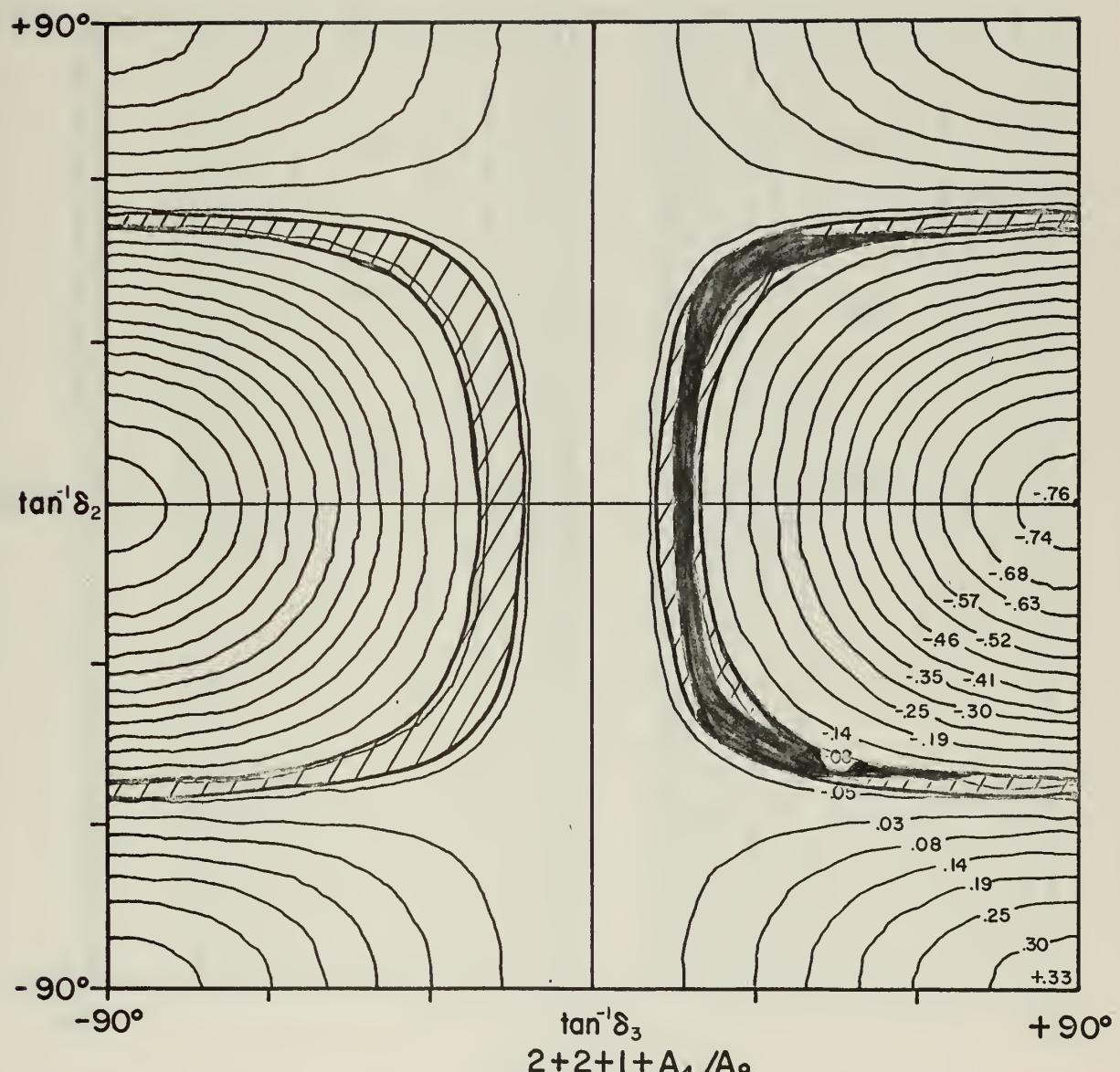




CASE SHOWN FOR  $\delta_4 = 0$   
 CONTOUR MULTIPLIER =  $.75 + .25 \sin(2 \tan^{-1} \delta_4 + 90^\circ)$

FIGURE 18a

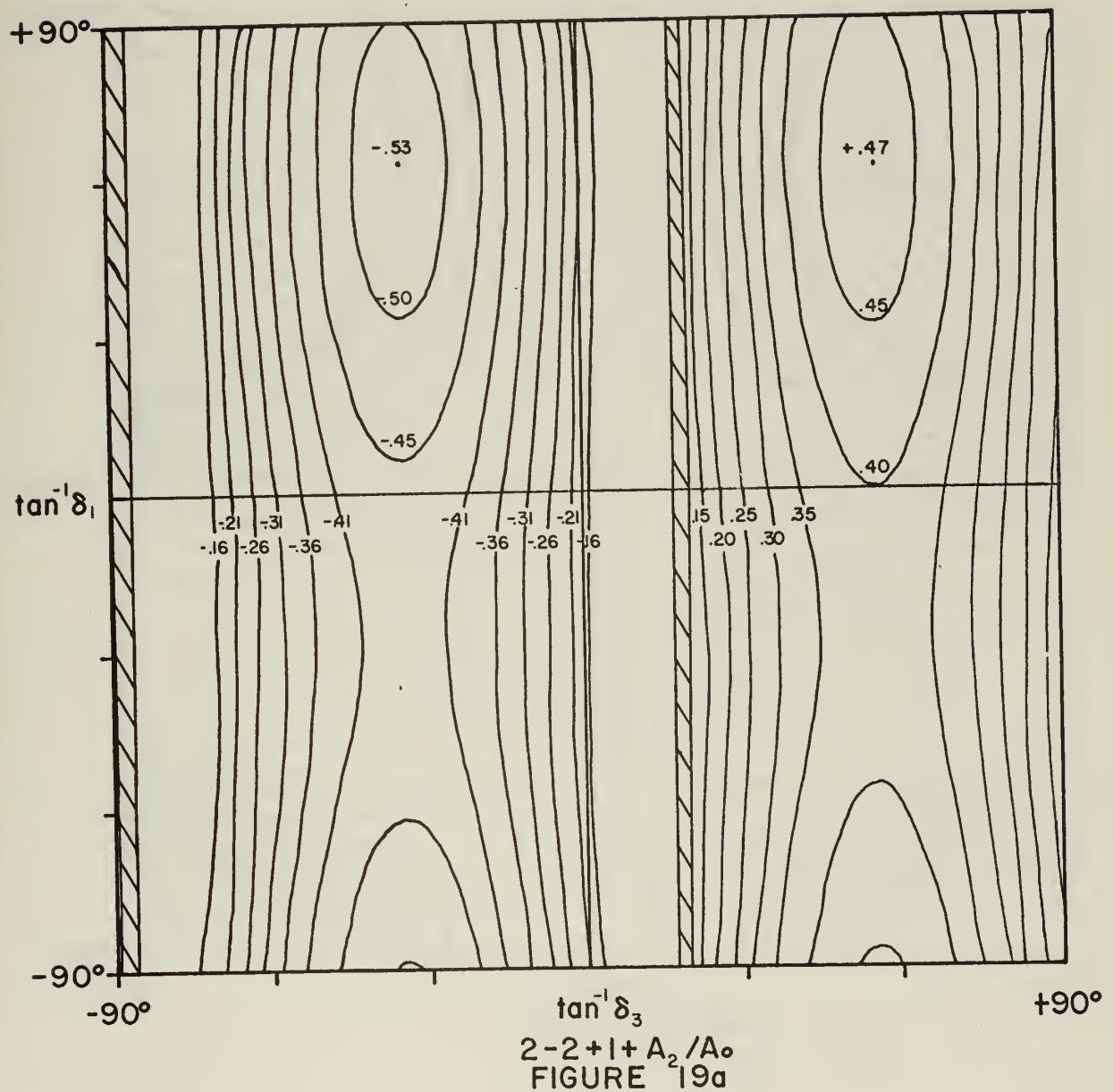




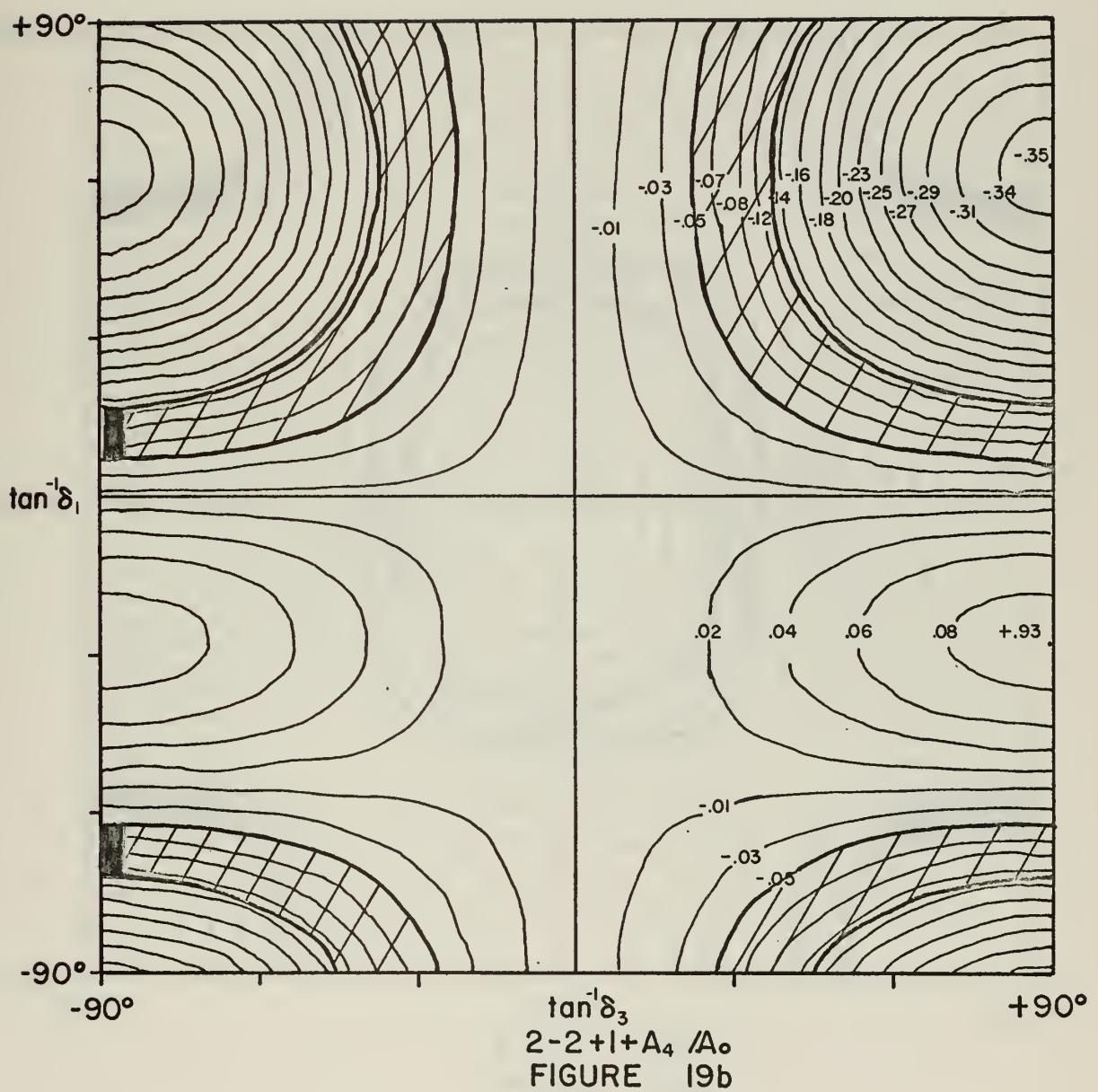
CASE SHOWN FOR  $\delta_4 = 0$   
 CONTOUR MULTIPLIER =  $-.127 + .635 \sin(2\tan^{-1}\delta_4 - 90^\circ)$

FIGURE 18b

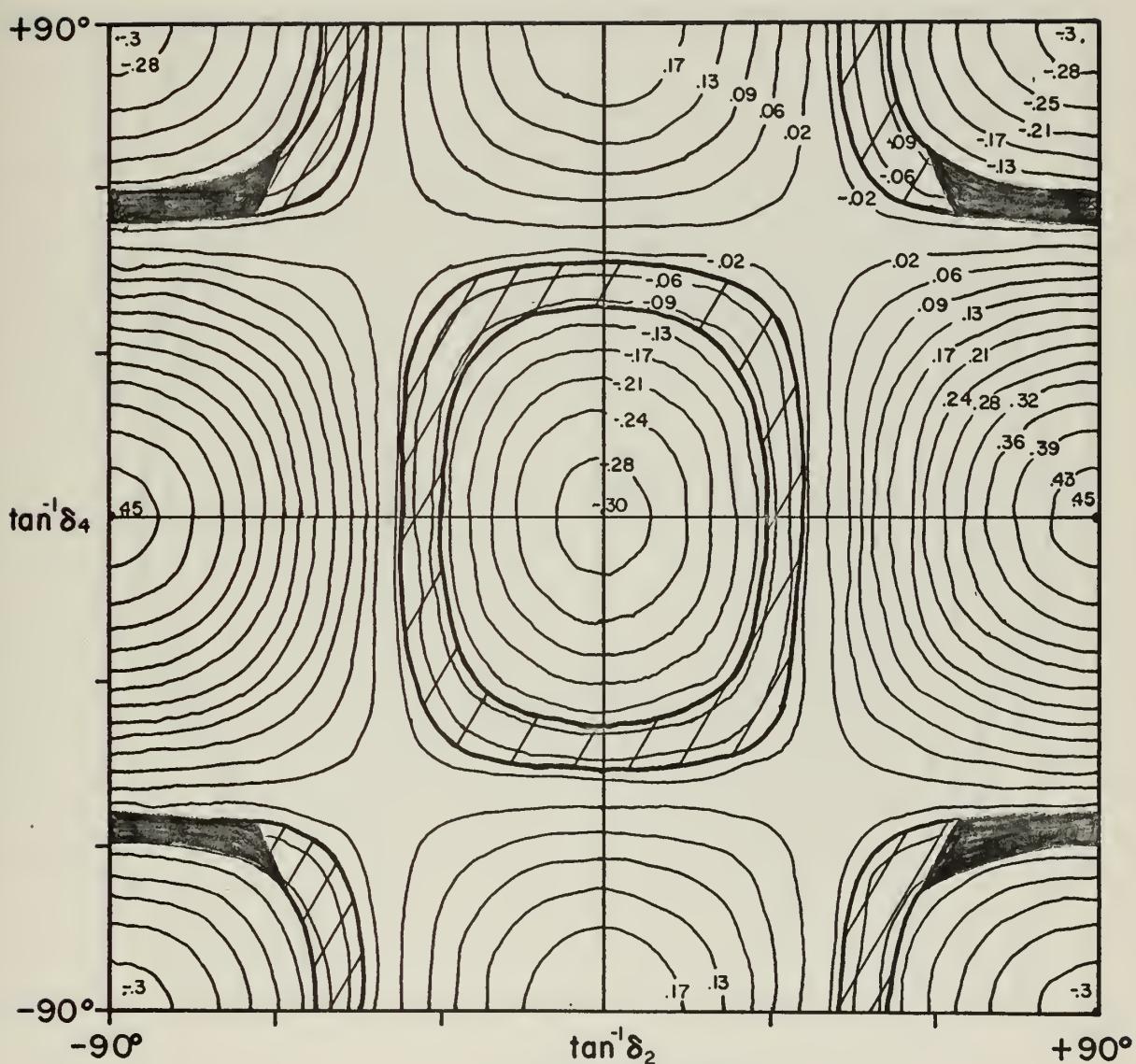






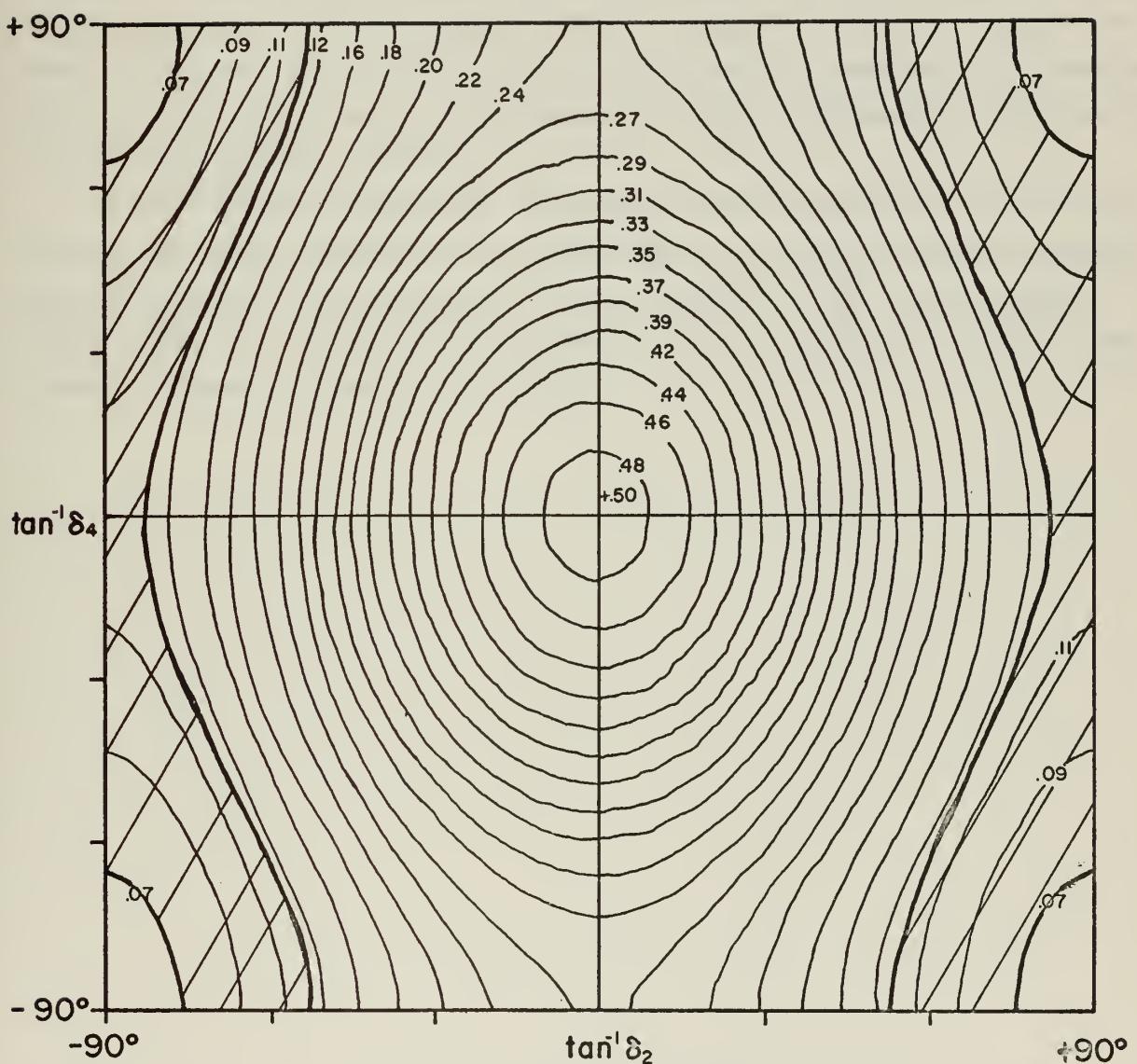






$\tan^{-1} \delta_2$   
 $2+3+1+A_4/A_0$   
 FIGURE 20a





$2+3+1+A_2/A_0$   
**FIGURE 20b**



## V ACKNOWLEDGEMENTS

I wish to thank my advisors, Professor H.E. Handler and Professor E.A. Milne for their encouragement, guidance, and advice offered during the course of this work. I also am most grateful to Dr. L.W. Seagondollar and Dr. R.A. Moore of the University of Kansas for pointing out the significance of the 4.50-Mev level spin determination and to Dr. Moore for furnishing results of his experimental work. Among the many other people who have helped make this experiment possible, I wish to thank Mr. R.C. Moeller, Mr. M.A. Brillhart, and Mr. R.A. Garcia for their indispensable contributions to the experimental equipment.

I also wish to acknowledge the help and understanding of my wife Lois during the years consumed by this undertaking. I wish, finally, to express my appreciation of the cooperation and confidence of the Civil Engineer Corps of the U.S. Navy which have made possible my absence from more normal duties in order to complete this work.



## Appendix 1 - Theoretical Description of Double Angular Correlations

## A. General

There has been abundant literature coverage of the theory of angular correlations of successive nuclear radiations since Hamilton first predicted the effect and developed theoretical expressions for successive emission of quanta.<sup>45</sup> A comprehensive review of the subject has been published by Biedenharn and Rose<sup>46</sup> and recent articles useful in experimental applications have been published by Gove<sup>47</sup> and by Biedenharn.<sup>48</sup> The following notation is essentially that of Biedenharn. In all cases a primed symbol represents a larger value than the same symbol without a prime.

In the case of a ( $p, \gamma$ ) reaction in which the direction of a primary gamma ray deexciting a resonance level is observed relative to the direction of an incident proton beam, the angular correlation is referred to as a double correlation. The function describing the relative probability of the gamma decay as a function of the angle between the beam direction and the emitted gamma ray can be written as an expression in the even Legendre polynomials:

$$W(\theta) = \sum_{\nu=0}^{\nu_{\max}} A_{\nu}(1) A_{\nu}(2) P_{\nu}(\cos\theta), \quad \text{even} \quad (1)$$

$$\begin{aligned} \text{where } A_{\nu}(i) &= a_{\nu}(L_i L_i) F_{\nu}(L_i L_i j_i J) + 2\delta_h a_{\nu}(L_i L_i') F_{\nu}(L_i L_i' j_i J) \\ &+ \delta_h^2 a_{\nu}(L_i' L_i') F_{\nu}(L_i' L_i' j_i J) \end{aligned}$$



and  $F_2(LL'j_1J) = (-1)^{j_1-J-1} [(2L+1)(2L'+1)(2J+1)]^{\frac{1}{2}} \cdot$

$$(LL'1 -1|\nu 0)W(JJLL';\nu j_i)$$

The index h is related to the index i by:

$$h = i^2 - i + 1$$

$(LL'1 -1|\nu 0)$  and  $W(JJLL';\nu j_i)$  are respectively the

Clebsch-Gordon coefficients and the Racah coefficients which are commonly used to describe systems involving coupling of angular momenta. The F functions, for which an extensive tabulation is available,<sup>49</sup> are defined in terms of a Clebsch-Gordon and a Racah coefficient by Biedenharn and Rose.<sup>46</sup> Arguments of the coefficients are angular momenta of the sequence:

$$j_1 \begin{pmatrix} L'_1 \\ L_1 \end{pmatrix} J \begin{pmatrix} L'_2 \\ L_2 \end{pmatrix} j_2 \quad (2)$$

$j_1$  = channel spin = proton spin  $\oplus$  target spin

$L'_1, L_1$  = interfering proton orbital angular momenta

$J$  = spin of resonance level

$L'_2, L_2$  = interfering gamma ray multipolarities

$j_2$  = spin of final state (not necessarily a ground state)

$\delta_1$  = mixing parameter representing the ratio of the reduced matrix element for the transition involving  $L'_1$  proton waves to that for  $L_1$  waves

$\delta_3$  = mixing parameter representing the ratio of the amplitude of gamma rays of multipolarity  $L'_2$  to those of multipolarity  $L_2$



$$a_z(LL') = \cos(\xi_L - \xi_{L'}) \cdot \frac{2[L(L+1)L'(L'+1)]^{\frac{1}{2}}}{L(L+1) + L'(L'+1) - z(\nu+1)} \quad (3)$$

is a particle parameter in which  $\xi_L$  and  $\xi_{L'}$  are the phases at the nuclear surface of the particle partial wave functions of multipolarity  $L$  and  $L'$ . The coherent interference between partial waves is expressed by the  $\cos(\xi_L - \xi_{L'})$  term. The factor  $\xi$  includes the coulomb phase shift due to electrostatic potential for charged particles and the phase shift occurring in "hard sphere" scattering. The  $a_z(L_2L'_2)$  terms for photons as opposed to particles are equal to unity. In the case where mixing of channel spins  $j_1$  and  $j'_1$  occur,  $W(\theta)$  may be expressed as an incoherent mixture:

$$W(\theta) = w_{j_1}(\theta) + \delta_4^2 w_{j'_1}(\theta) \quad (4)$$

where  $\delta_4^2$  is a mixing parameter expressing the ratio of probability of channel spin  $j'_1$  entering the reaction to that of channel spin  $j_1$ , ( $j'_1 j_1$ ). The maximum order of Legendre polynomials appearing in (1),  $\nu_{\max}$ , is determined by conditions imposed on the  $W$  (abcd; ef) coefficients by conservation of angular momentum which may be expressed by the so-called triad relationships. These require each of the triads of values (acf), (bdf), (abe), and (cde) to form a possible triangle. Restriction to the case of intermediate states of well defined parity (i.e. no interfering intermediate states of opposite parities) leads to the further simplification that  $\nu$  may take on only even values.

In the case of the reaction  $(p, x\gamma)$ , in which  $x$  represents an unobserved intermediate gamma ray, the correlation is expressed as an incoherent mixture of multipoles occurring in the unobserved transition, i.e. in



the sequence:

$$j_0 \left( \begin{array}{c} L_0 \\ L_0 \end{array} \right) J_1 \left( \begin{array}{c} L_1 \\ L_1 \end{array} \right) J_2 \left( \begin{array}{c} L_2 \\ L_2 \end{array} \right) j_2 \quad (5)$$

$$W(\theta) \text{ can be expressed: } W(\theta) = W_{L_1}(\theta) + \delta_2^2 W_{L_1'}(\theta) \quad (6)$$

where:

$j_0$  = channel spin

$L_0, L_0'$  = interfering proton orbital angular momenta

$J_1$  = spin of resonance level

$L_1, L_1'$  = interfering gamma ray multipolarities of unobserved radiation

$J_2$  = spin of intermediate level

$L_2, L_2'$  = interfering gamma ray multipolarities of observed radiation

$j_2$  = spin of final state

$\delta_2^2$  = ratio of intensity of the unobserved gamma ray of higher multipolarity to the intensity of the lower multipolarity gamma ray.

The correlation function now contains an additional part expressing the connection between the two intermediate levels by the unobserved radiation as follows:

$$W(\theta) \sum_{\nu}^{N_{\max}} A_{\nu}(0) C_{\nu}(1) A_{\nu}(2) P_{\nu}(\cos\theta) \quad (7)$$

where the  $A_{\nu}(i)$  are the same as in (1) and

$$C_{\nu}(1) = (-1)^{\nu} [(2J_1+1)(2J_2+1)]^{\frac{1}{2}} W(J_1 - L_1 J_2; J_1 J_2)$$

It is seen that for the general double correlation case four mixing parameters may occur (barring the further complication of mixed intermediate states).



In the present investigation the  $\text{Si}^{29}$  target nucleus has a ground-state spin of  $1/2+$ , with the simplifying result that channel-spin mixing and proton-partial-wave mixing may not occur simultaneously. The number of  $\delta$ 's is thus limited to a maximum of three. Explicitly, for resonance levels of  $1+, 2-, 3+, 4-$ , etc. orbital angular momentum mixing of two partial waves may occur but channel-spin mixing is prohibited. For the resonance level spins of opposite parities from those listed only channel-spin mixing is permitted.

When three mixing parameters occur in the general case of intermediate unobserved radiation ( $p, x\gamma$ ) with a spin sequence

$$\begin{pmatrix} j'_0 \\ j_0 \end{pmatrix} L_0 J_1 \begin{pmatrix} L'_1 \\ L_1 \end{pmatrix} J_2 \begin{pmatrix} L'_2 \\ L_2 \end{pmatrix} j_2 \quad \text{or} \quad j_0 \begin{pmatrix} L'_0 \\ L_0 \end{pmatrix} J_1 \begin{pmatrix} L'_1 \\ L_1 \end{pmatrix} J_2 \begin{pmatrix} L'_2 \\ L_2 \end{pmatrix} j_2 \quad (8)$$

the normalized correlation function can be expressed as

$$W(\theta) = 1/A_0 \sum_z A_z P_z(\cos\theta) \quad (9)$$

where each  $A_z/A_0$  is a product of the form:

$$A_z/A_0 = \prod_{\mu} g_{z\mu}(\delta_{\mu}) \quad (10)$$

with  $\mu = 1, 2, 3$  or  $2, 3, 4$

and 
$$g_{z\mu} = \frac{a_{z\mu} + \delta_{\mu} b_{z\mu} + \delta_{\mu}^2 c_{z\mu}}{(1 + \delta_{\mu}^2)}$$

(It is to be noted that the coefficient  $A_z$  appearing in (9) is a conventional representation of the more complete forms shown in (1) and (7). A distinction must also be made, of course, between the correlation function  $W(\theta)$  and the Racah coefficient  $W(abcd; ef)$ .



The coefficients a, b, and c can be expressed as follows:

$$\left\{ \begin{array}{l} \text{proton} \\ \text{wave} \\ \text{mixing} \end{array} \right. \left\{ \begin{array}{l} a_{\nu 1} = \frac{2L_0(L_0 + 1)}{2L_0(L_0 + 1) - \nu(\nu + 1)} F_\nu(L_0 J_0 J_1) \\ b_{\nu 1} = \frac{4[L_0(L_0 + 1)L_0^!(L_0^! + 1)]^{1/2}}{L_0(L_0 + 1) + L_0^!(L_0^! + 1) - \nu(\nu + 1)} F_\nu(L_0 L_0^! J_0 J_1) \\ c_{\nu 1} = \frac{2L_0^!(L_0^! + 1)}{2L_0^!(L_0^! + 1) - \nu(\nu + 1)} F_\nu(L_0^! J_0 J_1) \end{array} \right. \quad (11)$$

$$\left\{ \begin{array}{l} \text{unobs.} \\ \text{gamma} \\ \text{mixing} \end{array} \right. \left\{ \begin{array}{l} a_{\nu 2} = (-1)^\nu [(2J_1 + 1)(2J_2 + 1)]^{1/2} W(J_1 \nu L_1 J_2; J_1 J_2) \\ b_{\nu 2} = 0 \\ c_{\nu 2} = (-1)^\nu [(2J_1 + 1)(2J_2 + 1)]^{1/2} W(J_1 \nu L_1^! J_2; J_1 J_2) \end{array} \right.$$

$$\left\{ \begin{array}{l} \text{gamma} \\ \text{mixing} \end{array} \right. \left\{ \begin{array}{l} a_{\nu 3} = F_\nu(L_2 J_2 J_2) \\ b_{\nu 3} = 2F_\nu(L_2 L_2^! J_2 J_2) \\ c_{\nu 3} = F_\nu(L_2^! J_2 J_2) \end{array} \right.$$

$$\left\{ \begin{array}{l} \text{chan.} \\ \text{spin} \\ \text{mixing} \end{array} \right. \left\{ \begin{array}{l} a_{\nu 4} = \frac{2L_0(L_0 + 1)}{2L_0(L_0 + 1) - \nu(\nu + 1)} F_\nu(L_0 J_0 J_1) \\ b_{\nu 4} = 0 \\ c_{\nu 4} = \frac{2L_0(L_0 + 1)}{2L_0(L_0 + 1) - \nu(\nu + 1)} F_\nu(L_0 J_0^! J_1) \end{array} \right.$$

$$\varphi_1 = \cos(\xi_{L_0} - \xi_{L_0^!})$$

$$\varphi_2 = \varphi_3 = \varphi_4 = 1$$



For the  $(p, \gamma)$  case the only changes required are:  $a_{\nu 2} = 1$ ,  $b_{\nu 2} = 0$ ,  $c_{\nu 2} = 0$  and  $J_1 = J_2 = J$ .

The angular correlation function, as an alternative, may be expressed in terms of the coefficients  $Z$  and  $Z_1$ , instead of the  $F$  coefficients:

$$\begin{array}{ll}
 \text{proton} & \left\{ \begin{array}{l} a_{\nu 1} = (-1)^{j_0} Z(L_0 J_1 L_0 J_1, j_0 \nu) \\ b_{\nu 1} = 2(-1)^{j_0} Z(L_0 J_1 L_0^! J_1, j_0 \nu) \\ c_{\nu 1} = (-1)^{j_0} Z(L_0^! J_1 L_0^! J_1, j_0 \nu) \end{array} \right. \\
 \text{wave} & \\
 \text{mixing} & \qquad \qquad \qquad (12)^*
 \end{array}$$
  

$$\begin{array}{ll}
 \text{unobs.} & \left\{ \begin{array}{l} a_{\nu 2} = (-1)^{-L_1} W(J_1 J_2 J_1 J_2; L_1 \nu) \\ b_{\nu 2} = 0 \\ c_{\nu 2} = (-1)^{-L_1^!} W(J_1 J_2 J_1 J_2; L_1^! \nu) \end{array} \right. \\
 \text{gamma} & \\
 \text{mixing} & 
 \end{array}$$
  

$$\begin{array}{ll}
 \text{gamma} & \left\{ \begin{array}{l} a_{\nu 3} = (-1)^{j_2} Z_1(L_2 J_2 L_2 J_2, j_2 \nu) \\ b_{\nu 3} = 2(-1)^{j_2} Z_1(L_2 J_2^! L_2^! J_2, j_2 \nu) \\ c_{\nu 3} = (-1)^{j_2} Z_1(L_2^! J_2 L_2^! J_2, j_2 \nu) \end{array} \right. \\
 \text{mixing} & 
 \end{array}$$
  

$$\begin{array}{ll}
 \text{chan.} & \left\{ \begin{array}{l} a_{\nu 4} = (-1)^{j_0} Z(L_0 J_1 L_0 J_1, j_0 \nu) \\ b_{\nu 4} = 0 \\ c_{\nu 4} = (-1)^{j_0^!} Z(L_0 J_1 L_0^! J_1, j_0^! \nu) \end{array} \right. \\
 \text{spin} & \\
 \text{mixing} & 
 \end{array}$$



For  $(p, \gamma)$  reactions  $J_1 = J_2 = J$ ,  $a_{22} = (2J+1)^{-1}$ ,  $b_{22} = 0$ ,  $c_{22} = 0$

\* (The two formalisms (11) and (12) are to some extent incompatible.

Not only may the coherent mixing parameter phases differ by  $\pm 1$ <sup>50,51</sup>

but also for  $L_0 = 0$  the parameter  $\delta$ , doesn't represent the same matrix element ratio in both notations. The notation above, (12), has been used for most calculations in the present work. These expressions are in agreement with Sharp<sup>52</sup> except that the sign of  $b_{23}$ , for the  $(p,x\gamma)$  case only, has been changed to agree with the part of Huby's correction applying to the observed gamma ray.<sup>50,51</sup> For the "unnatural" radiation mixtures ( $M_1/E_2$ ,  $E_3/M_2$ , etc.) the sign of  $b_{23}$  should be reversed.

This sign convention permits the use of Moore's tabulated values<sup>23</sup>

where the same convention is used.)

The  $Z$  and  $Z_1$  coefficients and Racah and Clebsch-Gordon coefficients have been tabulated by Sharp et al.<sup>52</sup> Using this tabulation Moore<sup>23</sup> has prepared a convenient table of the  $a$ ,  $b$ , and  $c$  coefficients applicable to radiative proton capture by  $Si^{29}$ .

In order to fit experimental data on angular correlation coefficients to theory, it is desirable to first eliminate as many spin sequences as possible on the basis of any evidence which may be available.

In cases where only two mixing parameters enter, contour plots<sup>47</sup> showing contours of constant theoretical values of the coefficients  $A_2/A_0$  can be plotted against the inverse tangents of the two mixing parameters for all possible spin sequences. If three mixing parameters enter, one of them may be held constant for the purpose of plotting the correlation coefficient contours against the inverse tangents of the other two parameters. It is seen from (10) that the contour values for different values of the fixed parameter differ only by a multiplicative constant.



When experimental values of correlation coefficients are delineated on the contour plot, some spin sequences can be eliminated and the possible combinations of mixing parameters are restricted for other sequences. The observed values of several correlation coefficients in one decay scheme may then fix or highly restrict values of the spin-parities of resonance and intermediate levels, and determine the values of the mixing parameters.

#### B. Application to 1470-kev resonance

For the resonance occurring at a laboratory proton energy of 1470 kev, the phase factor  $\cos(\xi_L - \zeta_L)$  has been calculated using the expression  $\xi_L = \zeta_L + \phi_L$  where  $\zeta_L$  represents the phase shift due to the point coulomb potential and  $\phi_L$  is the "hard sphere" phase shift.

$$\zeta_L \text{ is defined } ^{42} \text{ by } \exp(2i\zeta_L) = \frac{\Gamma(L+1+i\eta)}{\Gamma(L+1-i\eta)} \quad (13)$$

where  $\eta = \frac{ZZ' e^2}{\hbar v}$  which leads to:

$$\zeta_{L+2} - \zeta_L = \tan^{-1}\left(\frac{\eta}{L+2}\right) + \tan^{-1}\left(\frac{\eta}{L+1}\right) \quad (14)$$

The value for  $\phi_L$  defined by

$$\exp(2i\phi_L) = \frac{G_L(\eta, kR) - iF_L(\eta, kR)}{G_L(\eta, kR) + iF_L(\eta, kR)} \quad (15)$$

where F and G are the regular and irregular solutions to the radial wave equation for the coulomb potential. The value for  $\phi_L$  can be found from the graphs of Sharp et al.<sup>53</sup> (these graphs actually list  $-\phi_L$ ) and employing  $R = 1.5A^{1/3}f.$  as an estimate of the nuclear radius. These graphs have also been used to calculate the relative penetration factors

$$v_L/v_{L-2} = \frac{F_{L-2} + G_{L-2}}{F_L^2 + G_L^2} \quad (16)$$

which are useful as a guide to the expected mixing parameter of interfering partial proton waves.<sup>42</sup> Results of these calculations appear in Tables 8 and 9.



TABLE 8

Relative Penetration Factors for 1470-Kev Protons

d/s	f/p	g/d
$3.1 \times 10^{-2}$	$5.0 \times 10^{-3}$	$1.3 \times 10^{-3}$

TABLE 9

Coulomb Phase Factors for 1470-Kev Protons

$\xi_2 - \xi_0$	$\xi_3 - \xi_1$	$\xi_4 - \xi_2$
$104^\circ$	$73^\circ$	$56^\circ$



## Appendix 2 - Computer Programs and Analysis

### A. General

In order to analyze experimental data in the form of gamma-ray spectra recorded by the 128-channel analyzer, two computer programs were developed. The first and more extensive program entitled program SPECFIT computes experimental correlation coefficients, while program CONTOUR is designed to utilize a California Computer Products Model 565 graph plotter as an output medium to automatically plot contours of constant angular correlation coefficient values against the inverse tangents of two mixing parameters. Both programs are written in fortran language and were run on a CDC-1604 digital computer.

### B. Program SPECFIT

The output pulses of a photomultiplier tube viewing a scintillation crystal upon which monoenergetic photons impinge do not all have the same voltage. Instead, the three fundamental processes of photoelectric absorption, Compton scattering and pair production, which occur in the crystal with probabilities dependent upon gamma-ray energy, lead to a response pulse-spectrum which is characteristic of each particular gamma-ray energy. Experimental geometry, crystal size, shielding, various line-broadening effects, and characteristics of individual crystal-photomultiplier tube combinations influence the response also, so that response functions are in general unique for each spectrometer in a given experimental arrangement.

An experimentally observed spectrum is, of course, usually composed of a number of gamma rays, the spectrum being the sum of the constituent radiations each multiplied by its intensity. The problem of deciphering such an observed spectrum to obtain the energies and intensities of component gamma rays is of fundamental importance in gamma ray spectroscopy and has not yet been satisfactorily solved in complete generality.

Using notation, which is appropriate for the discrete pulse height intervals into which multi-channel analyzers sort pulses, and also appropriate for the discrete number of gamma ray energies composing resonance deexcitation spectra, one can write:

$$S_i = \sum_{j=1}^m R_{ij} I_j$$



where  $s_i$  = observed number of counts in channel  $i$

$R_{ij}$  = probability that a photon of energy  $E_j$   
originating at a particular source  
will produce a pulse falling in pulse  
height channel  $i$  (usually referred to  
as the response matrix)

$I_j$  = the number of photons of energy  $E_j$   
originating at the source

$m$  = the number of gamma ray energies which  
occur in the spectrum under consideration

A rather fundamental and obvious approach to this problem is to assume no knowledge of gamma-ray energies is available, divide the gamma ray energy range into a number of discrete intervals equal to the number of channels, then solve for:

$$I_i = \sum_j \sum_k R_{ik}^{-1} R_{kj} I_j = \sum_k R_{ik}^{-1} s_k$$

In fact, the first attempt to unfold experimental spectra in the present work utilized a modification of a program<sup>54</sup> which employed an iterative technique developed by Schofield<sup>55</sup> to solve the above equation. This attempt failed for two reasons which are rather instructive. First, the analytical expression used to approximate the response matrix did not adequately represent the true spectrometer response, and second, the technique proved inadequate with experimental counting statistics attainable to handle the analysis of approximately 100 channels.

In order to avoid the first problem resort was had to the technique of interpolating between experimentally-determined monoenergetic calibration curves in order to obtain the response matrix. In order to overcome the second objection it was decided to use information on the gamma ray energies available from preliminary "stripping" of spectra and from other sources. Having thus reduced the number of gamma rays to be considered, a linear combination of the selected response matrix elements was then fitted to the observed spectrum by a least squares process. This procedure was followed by another least squares analysis to obtain



correlation function coefficients from the gamma ray intensities calculated for a series of runs at different experimental angles. The steps in the computation procedure in program SPECFIT as finally developed are as follows:

1. Load data. (See section D of this Appendix)
2. Subtract any constant background from calibration spectra. This background represents the Compton "tails" of higher energy gamma rays contaminating certain of the calibration gamma rays.
3. Normalize each calibration spectrum to a photopeak height of one unit.
4. Include a straight line in the response matrix if this option is selected. This step is essentially a refinement of step 2.
5. Correct the input and background spectra for dead time of the pulse height analyzer. For this purpose the dead time given by the manufacturer is  $38 + 0.5i$  microseconds where  $i$  is the pulse height channel number. This leads to an expression for corrected counts:

$$N'_j = \frac{N_j}{1 - 1/T \sum_{i=2}^{127} N_i (38 + 0.5i) \cdot 10^{-6}}$$

where  $N_i$  = the number of counts observed in channel  $i$

$T$  = counting time in seconds

6. Calculate energies in Mev corresponding to each channel number of the input spectrum.
7. Adjust energy scale and number of channels of background spectrum to match the input spectrum. Multiply by ratio of run times to get the proper background to be subtracted from the input spectrum and perform the subtraction.
8. Using the monoenergetic calibration spectra, calculate a matrix of response curves for the gamma rays which the spectrum is to be decomposed into. This calculation is performed in a subroutine entitled RESPONS as follows:
  - a. Find the two calibration curves bracketing a gamma ray energy desired as a response matrix element.



- b. Shift the lower-energy calibration curve found in a. so that its photopeak falls in the same channel as the higher energy photopeak. Fill in bottom of shifted curve as determined by the slope between the bottom two unshifted channels. Match energy calibration of the lower-energy curve to the upper. (For this step, and elsewhere where interpolation is necessary, cubic interpolation is performed using the four nearest points to the desired value.)
- c. Interpolate linearly between each pair of points of the bracketing calibration curves to obtain a response curve for the desired gamma-ray energy. (A "curve" in this description is meant to imply the vector consisting of a  $1 \times n$  array of values,  $n$  being the number of channels used in the description).
- d. Using the input spectrum channel energies calculated in 6., interpolate to get a response curve with channels and channel energies corresponding to those of the input data.
- e. Normalize the response curve for the given gamma-ray energy so that the ordinate corresponding to each channel represents the probability that one photon leaving the source will produce a count in that channel. For this calculation the experimentally-determined intrinsic peak efficiency data of Lazer<sup>56</sup> have been used to derive the following approximate expression for intrinsic peak efficiency as a function of source distance:

$$\epsilon = AE^{-x}$$

where  $A = .165 + .0067h$

$$x = .9699 - .00918h$$

$h$  = source to crystal distance in centimeters

$E$  = gamma ray energy in Mev

Assuming an accurately gaussian photopeak, and that the crystal resolution is approximately inversely proportional to the square root of gamma ray energy, (an assumption which was experimentally verified) the ratio of photopeak height to area can be expressed:

$$\alpha = \frac{1}{6\sqrt{2\pi}} = \frac{2}{WR} \sqrt{\frac{\ln 2}{.662\pi E}} = \frac{1}{.866WR \sqrt{E}}$$



where R is the usual resolution definition of peak width at one half maximum amplitude divided by energy for the Cs<sup>137</sup> gamma ray, and W is the spectrometer calibration in pulse height channels per Mev. When  $\alpha$  is multiplied by the appropriate solid angle fraction and attenuation coefficient, the value obtained represents the probability of recording one count in the photopeak channel when one photon originates at the source.

f. Return to main routine.

9. Fit the experimental spectrum with a linear combination of the selected monoenergetic response curves, the coefficients representing absolute intensities because of the method of normalizing the response matrix. The weighted sum of the squares of the differences between observed channel counts and fitted counts is minimized in order to arrive at this fit.

The normal equations resulting from the minimization are:

$$\sum_j S_j R_{jk} w_j = \sum_{ij} R_{ji} R_{jk} I_i w_j \approx 0 \quad i < j$$

where  $S_j$  = Observed counts in channel j

$R_{jk}$  = Response matrix element giving probability of a count in channel j due to one photon of energy  $E_k$  originating at the source

$w_j$  = Weight assigned to the squared error in channel j

$I_i$  = Number of photons of energy  $E_i$  which leave the source during the counting period

j summations are over the channels for which a fit is desired

i summation is over the number of gamma ray energies  $E_i$  used to fit the observed spectrum

For calculation the  $I_i$  are expressed:  $I_i = \sum_k A_{ki}^{-1} B_k$

with  $A_{ki} = \sum_j R_{ji} R_{jk} w_j$  and  $B_k = \sum_j S_j R_{jk} w_j$



A standard subroutine (GAUSS3) is used to invert A. Mention should be made of the weights  $w_j$  which have been used. Normally, for a Poisson distribution the weight is taken as  $1/\sigma^2$  where the number of observed counts in a channel are used as an estimate of the variance  $\sigma^2$ . Ferguson,<sup>57</sup> who has pointed out that this procedure overweights statistically low counts, identifies the Poisson distribution variance, which rightly is the mean number of counts, with the fitted number of counts for each channel. This procedure is adopted here with a modification to account for statistical error arising from the subtraction of background counts. We have  $S'_i = S_i - b_i$ , where  $S_i$  are observed channel counts corrected for dead time and  $b_i$  are background counts to be subtracted. Then,

$$\sigma^2(S'_i) = \sigma^2(S_i) + \sigma^2(b_i)$$

$$\approx S_i + (T/T_b)b_i = S'_i + (1 + T/T_b)b_i$$

where  $T$  = counting time for data and  $T_b$  = counting time for background run. Identifying  $S'_i$  with the computed number of counts,  $S''_i$ , then gives  $S''_i + (1 + T/T_b)b_i$  for the channel variances. To avoid overweighting of very small values of fitted counts an arbitrary minimum value of  $\sigma^2$  is chosen to be two per cent of the average value of  $\sigma_i^2$  over the fitted channels. Since no fitted count values are initially available, an initial weight of one is assigned to each channel and the solution is then iterated to convergence, stopping when

$$\sum_i \left( \frac{\delta I_i}{I_i} \right)^2 \text{ is smaller than } 10^{-6}.$$

(This series of iterations, usually five in the present work, accounts for most of the program running time and can be abbreviated for preliminary analyses; see reference 57 in this regard.)  $\delta I_i/I_i$  is the fractional change in the computed intensity of the  $i^{\text{th}}$  gamma ray between successive iterations.

10. The computed intensities are examined for algebraic sign. Since only positive intensities have a physical basis any negative intensities are discarded and the problem resolved to this point.

11. For each gamma ray calculate the fitted total number of counts in the channels above the channel representing 40 per cent of the photopeak energy.



(This is an arbitrary calculation used only for checking results obtained in manual unfolding). The fractional error assigned to these counts is set equal to the fractional error in the corresponding intensity. The statistical variance estimators of the intensities have been taken to be the diagonal elements of the inverse of the matrix A in 9.<sup>41</sup> In addition to statistical intensity errors there is an error due to gain and discriminator drifts in the monitor counts channel which must be accounted for as well as error due to eccentricity of the beam spot on the target. These fractional errors have been estimated to be 3.3 per cent and 1.2 per cent respectively in the present experiment. The three errors are combined according to:

$$1/\zeta^2 = \sum_j 1/\zeta_j^2.$$

Two other possible sources of error should be noted here. First, the errors in the experimentally determined response matrix, which can considerably complicate the error calculations for the intensities, are kept to a minimum by obtaining good counting statistics in the calibration spectra. The other source of error is gain drift which tends to smear spectral peaks. Fortunately, the main effect of such drift seems to be an increase in the mean squared errors in the curve fitting process and intensity determinations are apparently not too sensitive to the effect.

12. The intensities and their errors calculated in 11 are read out and stored for use in the subsequent determination of correlation coefficients and mean values.

13. The absolute values of differences between observed and fitted counts are summed over fitted channels and divided by total input spectrum counts over the same range to give a measure of the fit which has been achieved in the analysis. This measure is then read out.

14. The input and fitted spectra are plotted in a graph output.

15. The above process is repeated for each run in the data set until the data for all the experimental angles has been processed.

16. When the preceding analysis is completed for a set of data a subroutine entitled LEASTSQ may be called, if appropriate, to compute the correlation coefficients. For each gamma ray in the spectrum a least squares fit of the counts computed at the experimental angles is made to the usual series:

$$W(\theta) = \sum_i^1 A_i P_i(\cos \theta), \quad i \text{ even}$$



where  $P_i$  are the even Legendre polynomials. The fit is first made to  $A_0 + A_2 P_2 + A_4 P_4$  then  $A_0 + A_2 P_2$ . (Provision has been made for fitting to the series terminating with  $P_6$  by simply changing the initial value of NP to 4; however, see reference 61 concerning accuracy of  $P_6$  terms). The computation, which follows the procedure given by Rose<sup>41</sup>, proceeds as follows:

- a. Table gamma ray absorption coefficients<sup>58</sup> in copper and corresponding energies in Mev. Correct intensities and sigmas for attenuation in copper target backing.
- b. Set up matrix  $P_{ij}$  of Legendre polynomials for experimental angles.
- c. Do least squares calculation to find coefficients using the reciprocal of the estimated intensity variances from 11 as weights.
- d. Normalize coefficients and correct for the effect of the finite solid angle subtended by the scintillation crystal. These latter corrections are precomputed in an auxiliary program using the expressions given by Rose.<sup>41</sup> Estimated variances of the coefficient ratios,  $A_2/A_0$ , etc., are computed from the expression given by Reich et al.<sup>61</sup> (After computation variances were increased by a factor  $1/\epsilon^2$  in all cases where the computed value of  $\epsilon^2$  was less than unity.)
- e. Calculate anisotropy.
- f. Compute the criterion  $\epsilon^2$ . This quantity, defined by Rose's analysis, is a measure of non-statistical errors in the data. If such errors are eliminated the value of  $\epsilon^2$  serves to indicate whether the appropriate number of polynomials has been used in the correlation function expression, with values of  $\epsilon^2$  greater than unity indicating an inappropriate choice. It is pertinent here that non-statistical errors have already been introduced into the problem in 11 above so that this criterion is valid only to the extent that an accurate estimate of the true variance of the counts has been achieved. The computed values of  $\epsilon^2$  in general are near unity and seem to verify that the estimated variances can be considered to be reasonably good.



- g. Selected values of the normalized calculated correlation function, and values of a comparison theoretical function if desired, are read out, and similarly normalized experimental points are read out. (Experimental values are uncorrected for solid angle, the correction being applied to the theoretical function for comparison with experimental results.)
- h. Iterate for each gamma-ray energy.
- i. Reduce number of polynomials if orders higher than  $P_2$  have been used and iterate solution. (The program also calculates coefficients for the cosine series:

$$W(\theta) = 1/A_0 \sum_{n=1}^m A_{2n} [\cos(\theta)]^{2n}$$

- j. Return to main routine.

- 16. Repeat entire analysis for as many sets of data as indicated.
- 17. Combine coefficients to get weighted mean values and Chi square values if more than one set of data is analyzed. Read out these values.
- 18. Combine intensities at each angle from all sets of data and re-analyze for coefficients using this combined data set.

C. Glossary of names of selected variables and constants appearing in program SPECFIT. An asterisk indicates an input quantity.

A(I,J)	The matrix product of the response matrix with its transpose, weighted with the WT values.
AINV(I,J)	The inverse of A(I,J).
ANISO(I)	Anisotropy for Legendre series solution (I = 1), and for cosine series solution (I = 2).
ARGP(I)	Vector containing cosines of the angles B3(I).
*B(I)	A vector representing the product of the inverse of the response matrix and the observed spectrum.
*BACK(I)	A number representing background counts to be subtracted from calibration curve I.
BACKGND(I)	The calculated background spectrum to be subtracted from an input spectrum.
*BCHAN	The lower of the channel pair described under TCHAN



**FIN(I,J,K)** Storage for FINTENS(I), data set J, run K.  
**FMU(I)** Mass absorption coefficient in copper for EMU(I)  
 Mev gamma ray Units in cm<sup>2</sup>/gm.  
**\*HCM** Distance from source to crystal face in centimeters.  
**HOLD(I)** A temporary storage vector.  
**LPKS** Same as NPKS but not including any straight line component. (See NBAOPT)  
**\*MAXCHAN** An upper channel limit in loading CALB data. Between MAXCHAN and MINT, CALB(I) values are filled in with the value of CALB(MAXCHAN).  
**\*MESH** The number of channels in input spectrum.  
**\*MINT** Number of channels in calibration spectra.  
**\*MINTB** Number of background channels.  
**MPKS** Initial value of NPKS for restoring the value.  
**MPKS2** Initial value of LPKS for restoring the value.  
**\*NBAOPT** Background option selector.  
 NBAOPT = 0 No background read in nor subtracted from SPEC  
 NBAOPT = 1 Background spectrum is read in and SPEC is corrected for this background  
 NBAOPT = 2 Same as NBAOPT = 1 and in addition a straight line is included in the response matrix.  
**\*NCUT** A channel number below which the graph output values are divided by a scale factor SCALE.  
**NERFL** An output of matrix inversion subroutine. A value of 1 indicates matrix is non singular, and 2 indicates a singular matrix.  
**\*NHIEST** The upper channel of the range over which a fit is made in SPECFIT.  
**\*NLOWEST** The lower channel of the range over which a fit is made in SPECFIT.  
**\*NOCALB** Number of calibration spectra.  
**\*NPHOTO** The photograph number of the observed spectrum; for identification only.



*NPKS	The number of spectral components with which the observed spectrum is to fitted. Program will increase this value by one if option is chosen to include a straight line in R (See NBAOPT).
*NRUNS	Number of runs in a set of data.
NP	Number of Legendre polynomials to be fitted in first pass through LEASTSQ.
*NPRINT	<p>Print suppression index</p> <p>NPRINT = 0 All printouts and graphs are made</p> <p>NPRINT = 1 Suppress printouts of spectrum values and cosine series coefficients</p> <p>NPRINT = 2 Same as 1 and also suppress graph outputs</p> <p>NPRINT = 3 Same as 2 with briefer more compact arrangement of correlation coefficients</p> <p>NPRINT = 4 Same as 3 but graphs plotted</p>
*NSETS	Number of sets of data.
OMEGA(I,J)	Used for temporary storage in SPECFIT and diagonal elements used for weights in LEASTSQ.
*PKSEP(I)	The number of channels corresponding to 1.022 Mev for energy calibration of CALB curves. Taken as channel distance from photo peak to pair peak for higher energy calibration gammas.
*PHOT(I)	The channel of calibration curve I in which the photopeak occurs. (May be an interpolation between two integral channels numbers).
*PKEN(I)	Energy in Mev of a spectral component.
POLY(I,J)	The matrix of Legendre polynomials of order 2(J-1) evaluated at angle B3(I).
R(I,J)	The response matrix.
RED	A normalizing factor for 3" x 3" unshielded NaI(Tl) crystal representing the probability that one photon at the source will produce one count in the photopeak channel.
*RES	The proton energy in kev at the resonance being investigated (for identification only).
SAN(I)	Estimate of standard deviation of ANISO(I).
SAVE(I,J)	An auxiliary matrix.



SBB(I)	Coefficient for renormalizing input spectrum with counts from an auxiliary monitor (optional; if not entered values are unity).
*SCALE	A scale factor for graph output. See NCUT.
SCOEFT(I)	The standard deviation in COEFT(I).
SIGCTS(I)	The calculated error in COUNTS(I).
SIGTENS(I)	The error for calculated spectral intensity INTENS(I).
SIN(I,J,K,)	Storage for errors in FIN(I,J,K)
*SPEC(I)	The vector representing the counts in the observed spectrum.
STC(I,J,K)	Storage for accumulating calculated correlation coefficients of order I, for gamma ray J and data set K.
STORE(I,J)	A temporary storage matrix for A while it is being inverted.
STSC(I,J,K)	Standard deviation of STC(I,J,K) .
STW(I,J)	Temporary storage.
STSW(I,J)	Temporary storage.
SUM7(I)	Temporary storage.
*TBACK	Duration of background run in units of seconds x 60
TC(I)	The weighted mean value of correlation coefficient of order I.
*TCHAN	The upper of a pair of input spectrum channels to be used for sensing energy calibration of the input spectrum.
TEMP(I)	Temporary storage.
*TOPEN	Energy of TCHAN in Mev.
TS(I)	Standard deviation of TC(I)
*TX60SEC	Duration of a run in units of seconds x 60.
WT(I)	The weight assigned to channel I in least squares solution for intensities.
XI(I)	The matric product of POLY and B1 weighted with the diagonal elements of OMEGA.



XS(I) A vector representing the energy in Mev of each channel of the input spectrum.

Z(I) A temporary storage vector.

\*ZERO The number which if any diagonal element of matrix A is smaller than, A will be considered singular.



D. DATA INPUT SEQUENCE FOR PROGRAM SPECFIT

Input	Number of sets	Number of cards in set	Input name	Remarks
1	1	1	NSETS, NPRINT	4
2	1	1	EL, ECC, HCM	
3	1	1	NOCALB, MINT	
4	NOCALB	1	BACK(J), PHOT(J), ENCALB(J), MAXCHAN, PKSEP(J)	
		MAXCHAN/7	CALB(I,J)	
5	1	1	NPKS, NBAOPT	
6	1	NPKS	PKEN(I)	
7	1	1	BSL, BEN, BPHOT, MINTB, TBACK	1
8	1	MINTB/7	BTGT(I)	
9	1	1	NRUNS, CTSNORA, NLOWEST, NHDEST, NCUT, SCALE, ZERO	2
10	1	1	RES, NPHOTO	3
11	1	1	MESH, TCHAN, TOPEN, BCHAN, BOTEN, B3(NRUNDEX), T60XSEC	
12	1	1	CTSNORB	
13	1	MESH/7	SPEC(I)	
14	1	4	Graph data cards	5

REMARKS:

- 1.) Read in only for NBAOPT greater than 1.
- 2.) Input resumes here for each successive data set.
- 3.) Input resumes here for each successive run within a data set.
- 4.) All loading formats are multiples of F10.3 and I10 except that ZERO is E10.1 and TBACK is F20.3.
- 5.) Blank cards may be loaded if graph option not used.



## E. Description of program CONTOUR

As described in appendix 1 the correlation function coefficients for  $(p, \gamma)$  and  $(p, x\gamma)$  intermediate unobserved reactions for spin 1/2 target nuclides can be described as a product of, at most, three functions of the form:

$$g_{ik} = \frac{a_{ik} + b_{ik} \delta_k + c_{ik} \delta_k^2}{1 + \delta_k^2}$$

where  $i$  denotes the order of the coefficient and  $k$  is the index of the mixing parameter  $\delta$ . For computation the substitution  $\delta = \tan Y$  is commonly used, and leads<sup>60</sup> to  $g = \alpha + \beta \sin(2Y + \gamma)$

$$\text{with } \alpha = \frac{a + c}{2}$$

$$\beta = + \frac{s}{2} \left[ b^2 + (a - c)^2 \right]^{1/2}$$

$$\gamma = \tan^{-1} \left( \frac{a - c}{b} \right)$$

$$s = b / |b|$$

where indices have been dropped for clarity.

This program computes sets of possible values of two mixing parameters for various fixed values of correlation coefficients. Using a graph plotter output medium available through a locally developed subroutine, these sets of values are plotted to form a set of contours along which the value of the correlation coefficient is constant.

Steps in the computation are as follows:

1. Read in data for one possible spin sequence to be examined. Compute values of  $\alpha$ ,  $\beta$  and  $\gamma$  for the lowest order coefficient which occurs. If three mixing parameters occur simultaneously, read out the expression for the fixed  $g$  as a function of the fixed mixing parameter. (All loading formats are standardized to be F10.3 and I10 for floating and fixed point variables respectively.)
2. Find and read out the maximum and minimum possible values of the coefficient and the corresponding values of the  $Y$ 's.



3. Fix the coefficient at an initial value within its range. Assign values to one of the Y's in equal increments throughout the range  $-90^\circ$  to  $+90^\circ$  and solve for the corresponding values of the remaining Y, which in general will be double valued over the region selected.
4. Read out pairs of Y values for the selected coefficient value both to the print medium and to the graph plotter.
5. Repeat steps three and four for successively larger values of the coefficient until a specified number of contours has been plotted.
6. Repeat steps two through five until contour plots for coefficients of all orders permitted by the given spin sequence have been produced.

#### F. Input variables for CONTOUR

NSETS	Number of spin sequences for which plots are to be made.
JFIND	Index of the mixing parameter to be solved for. (Y values on graph). <ul style="list-style-type: none"> <li>1 = proton partial wave mixing</li> <li>2 = unobserved gamma multipolarity mixing</li> <li>3 = observed gamma multipolarity mixing</li> <li>4 = channel spin mixing</li> </ul>
JSTEP	Index of mixing parameter assigned successive values (x values on graph) See JFIND.
JFIX	Index of mixing parameter to be held constant when three parameters occur simultaneously. (Set to five for $(p, \gamma)$ reaction where only two parameters occur.)
IMAX	Index specifying the highest order coefficient occurring in the spin sequence. (Order = $2 \cdot IMAX$ )
JMAX	Largest value which the index of mixing ratio assumes for use in reading input data. (Usually set at four.)
NCONTS	Number of contours to be plotted on each graph.
NPTS	Number of pairs of points to be computed for each branch of one contour.
DELTA(JFIX)	Fixed value assigned to the third mixing parameter during a calculation.



RHO

Energy dependent Coulomb phase factor to be multiplied  
into the b coefficients for proton-partial-wave  
mixing.

A(I,J)

B(I,J)

C(I,J)

The a, b, and c factors occurring in the expression for  
g. I denotes the coefficient order (order = 2I) and  
J the mixing parameter index. See FIND.



G. Fortran Listing of Program SPECFIT

```

•• JOB * PHELPS NO GRAPHS - 35 MIN RUNNING TIME
PROGRAM SPECFIT
DIMENSION SPEC(130), BACK(130), CALB(530), PHOT(09), ENCALB(09)
1  TEMP(530), R(130), HOLD(530), PKEN(20), PKCTS
2  E(20), R(130), A(20), INV(20), STCR
3  B(20), COUNTS(20), SIGCTS(20), WT(130), SA...
4  CEE(4), COEE(4), CEE(4), CEE(4), PART1(9),
5  XE(4), COEFT(4), COEFT(4), COEFT(4), PART1(9),
6  SUM7(4), COANISO(2), SIGTENS(20), COEEINV(4,4), PART1(9),
7  FMU(20), EMU(20), PKSEP(10), BACKGND(130), BTGT(530),
8  STSC(4,20,6), TC(4), TS(4), FIN(20,5,5), SIN(20,5,5)
C  COMMON SPEC, XS, CALB, PHOT, BACK, HOLD, Z, DIFTS,
1  COKEN, PKCTS, R, FINTENS, EA, AINV, STORE, SIGTENS, SIGCTS,
2  COUNTS, WT, SAVE, FINTENS, PART1, CEE, CEEINV, XI, COEFT,
3  COEFT, ARGP, B1, B2, B3, SUM7, ANISO, SBB, STSW, BTGT,
4  MESH, MINT, GAME, ZERO, NERFL, NRUNS, PKSEP, LPKS,
5  NSETDEX, STC, STSC, NPRINT, HCM
NSETDEX = 1
C
LOAD DATA
READ INPUT TAPE 2, 831, NSETS, NPRINT
READ INPUT TAPE 2, 820, EL, ECC, HCM
READ INPUT TAPE 2, 820, NOCALB, MINT
DO 830 J=1, NOCALB
READ INPUT TAPE 2, 835, BACK(J), PHOT(J), ENCALB(J), MAXCHAN,
1  PKSEP(J)
1  READ INPUT TAPE 2, 815, (CALB(I,J), I=1, MAXCHAN)
DO 830C I = MAXCHAN, MINT
CALB(I,J) = CALB(MAXCHAN,J)
SUBTRACT CONSTANT BACKGROUND FROM CALIBRATION SPECTRA
C  12 DO 13 J = 1, MINT
DO 13 I = 1, MINT
13 CALB(I,J) = CALB(I,J) - BACK(J)
NORMALIZE CALIBRATION SPECTRA TO PHOTOPEAK HEIGHTS
C  505 DO 2C J = 1, NOCALB
NPHX = PHOT(J)
IF (CALB(NPHX + 1, J) - CALB(NPHX, J)) 507, 507, 509
507 CB = CALB(NPHX, J)
GO TO 510
509 CB = CALB( NPHX + 1, J)
510 DO 20 I = 1, MINT
20 CALB(I,J) = CALB(I,J) / CB
READ INPUT TAPE 2, 840, NPKS, NBAOPT
LPKS = NPKS
MPKS = NPKS

```



```

MPKS2 = LPKS
DO 845 I=1, NPKS
  READ INPUT TAPE 2, 825, PKEN(I), PKCTS(I)
C   E(I) = PKEN(I)
C   INCLUDE CONST VECTOR IN RESPONSE FOR NBAOPT OVER 1
  IF(NBAOPT - 1)846, 90
  NPKS = NPKS + 1
  MPKS = NPKS
  PKEN(NPKS) = 9.999
  E(NPKS) = PKEN(NPKS)
  DO 91 I = 1, 130
    R(I, NPKS) = 0.002
C   LOAD BACKGROUND FOR NBAOPT OVER 0
  91  READ INPUT TAPE 2, 848, BSL(BEN, BPHOT, MINIB, BACK
C   READ INPUT TAPE 2, 816, (BTGT(I), I = 1, MINIB)
  846  FMINTB = MINIB
  IF(FMINTB - 200.0) 847, 847, 849
  847 XDEAD = 0.50
  GO TO 850
  849 XDEAD = 0.125
  SUMD = 0.0
  DO 850 I = 2, MINTB
    FLI = I
    SUMD = SUMD + BTGT(I) * (38.0 + XDEAD * FLI)
  851  SUMD = SUMD * (6.0E-05 / BACK)
  DO 852 I = 1, MINTB
    BTGT(I) = BTGT(I) / (1.0 - SUMD)
  852  WRITE OUTPUT TAPE 3, 832, NSETDX
  806  WRITE OUTPUT TAPE 3, 1073
  LOAD DATA INPUT TAPE 2, 842, NRUNS, CTSNORA, NLWEST, NHDEST, NCUT,
C   READ INPUT TAPE 2, 805, RES, NPHOTO
  1  SCALE, ZERO
  NCUT = 23
  NRUNDEX = 23
  807  READ INPUT TAPE 2; 810, MESH, TCHAN, TOPEN, BCHAN, BOTEN, B3(NRUNDEX)
C   T60XSEC
  1  READ INPUT TAPE 2, 817, CTSNORB
  IF( CTSNORA - 1.0 ) 837, 837, 837
  836  IF( CTSNORB - 1.0 ) 837, 837, 838
  838  SBB(NRUNDEX) = CTSNORA / CTSNORB
  WRITE OUTPUT TAPE 3, 841, SBB(NRUNDEX)
  GO TO 839

```



```

837 SBB(INRUNDEX) = 1.0
839 READ INPUT TAPE 2 TIME OF ANALYZER
C CORRECT FOR DEAD TIME
SUMD = 0.0
DO 808 I = 2, MESH
FLI = 1
SUMD = SUMD + SPEC(I) * (SPEC(I), I=1,MESH)
DO 809 I = 1, MESH
SPEC(I) = SPEC(I)/(I-1) - SUMD
TF = 1.0 + T60XSEC/TBACK
KITER = 0
C CALCULATE INPUT SPECTRUM CHANNEL ENERGIES
SPSLOPE = (TOPEN - BOTEN)/(TCHAN - BCCHAN)
DO 855 I = 1, MESH
ENDEX = 1
XS(I) = BOTEN + (ENDEX - BCCHAN) * SPSLOPE
WRITE OUTPUT TAPE 3, 860, NRUNDEX
WRITE OUTPUT TAPE 3, 865, RES, NPHOTO
IF (NPRINT - 1) 853 883
WRITE OUTPUT TAPE 3, 870,
WRITE OUTPUT TAPE 3, 875, (XS(J), J=1, MESH)
WRITE OUTPUT TAPE 3, 880
WRITE OUTPUT TAPE 3, 885, (SPEC(J), J=1, MESH)
C ADJUST ENERGY CALIBRATION AND CHANNELS OF BACKGROUND IF USED
IF(NBAOPT - 1) 884, 884
DO 10 I = 1, MINTB
TEMP(I) = BTGT(I) * (T60XSEC/TBACK)
DO 675 K = 1, MESH
CHX = BPHOT + (XS(K) - BEN) * BSL
IF (CHX - 1.0) 520, 525
VAL = TEMP(1)
GO TO 675
IF(CHX - FMINTB) 535, 530
VAL = TEMP(MINTB)
GO TO 675
525
530
535

```

1

```

FLNCHX = CHX
FRACT = CHX - FLNCHX
IF (I - 1) 381, 381, 374
DEL1 = TEMP(I) - TEMP(I-1)
DEL2 = TEMP(I+1) - TEMP(I)
DEL3 = TEMP(I+2) - TEMP(I+1) - TEMP(I-1) - DEL1
U1 = 1.0 + FRAC
U2 = U1 * (U1 - 1.0) / 2.0
U3 = U2 * (U1 - 2.0) / 3.0

```



```

380 VAL = TEMP(I-1) + U1*DEL1 + U2*DEL2 + U3*DEL3
390 IF (VAL - 0.0 ) 390, 391, 391
391 GO TO 675
381 I = 1 + TEMP(I) - TEMP(I-1)
DEL1= TEMP(I+1) - TEMP(I) - DEL1
DEL2= TEMP(I+2) - TEMP(I+1) - DEL2 - DEL1
DEL3= FRACT(U1*U1 - U2*(U1 - 1.0)/2.0
U1 = U2*(U1 - 2.0)/3.0
GO TO 380
675 BACKGND(K) = VAL
SUBTRACT BACKGND FROM SPECTRUM FOR NBAOPT = 1 OR GREATER
C 887 DO 886 I = SPEC(I) - BACKGND(I)
886 SPEC(I) = SPEC(I) - BACKGND(I)
IF (NPRINT - 1) 888, 885, 885
888 WRITE(UNIT - 1) 888, 889
WRITE(UNIT - 1) 889, 889
WRITE(UNIT - 1) 889, 881 (BACKGND (J), J = 1, MESH )
WRITE(UNIT - 1) 881, 875 (SPEC(J), J=1, MESH )
WRITE(UNIT - 1) 875, 881 (SPEC(J), J=1, MESH )
C 885 CALCULATE RESPONSE MATRIX
C SOLVE LEAST SQUARES PROBLEM FOR RELATIVE INTENSITIES
DO 854 WT(I) = 1.0
854 WT(I) = 1.0, NPKS
102 DO 105 I = 1, NPKS
102 DO 105 K = 1, NPKS
DO 110 SUM = 0.0
110 SUM = SPEC(J) * WT(J) * R(J,K) + SUM
105 STORE(I,K) = SUM
CALL PGAUSS3( A(I,K) , NPKS, ZERO , STORE , AINV , NERFL )
DO 115 K = 1, NPKS
SUM = 0.0
DO 110 J = NLOWEST, NHIEST : R(J,K) + SUM
110 SUM = SPEC(J) * WT(J) * R(J,K) + SUM
B(K) = SUM
115 DO 125 L = 1, NPKS
SUM = 0.0
DO 120 K = 1, NPKS
120 SUM = AINV(K, L) * B(K) + SUM
125 FINTEEN(L) = SUM

```



```

135 DO 135 I = 1, NPKS
135 SIGTENS(I) = SQRTF(ABSF(AINV(I,I)))
C 136 ITERATE SOLUTION TIL CONVERGENCE WITH WTS FROM COMPUTED SPECTRUM
C 136 DO 137 I = 1, MESH
136 DIF(I) = 0.0
137 DO 137 J = 1, NPKS
137 DIF(I,J) = R(I,J) * FINTENS(J) + DIF(I)
137 SUMB = 0.0
137 DO 260 I = NLOWEST, NHIEST
137 SUMB = DIF(I) + SUMB * BACKGND(I)
137 FMESSH = MESH
137 RMESH = NHIEST - NLOWEST + 1
137 SUMB = 50.0 * RMESH/SUMB
137 DO 275 I = 1, MESH
137 IF((I-1.0/(DIF(I)+TF)*BACKGND(I)) - SUMB)265,265,270
137 WT(I) = 1.0/(DIF(I) + TF * BACKGND(I))
137 GO TO 275
137 WT(I) = SUMB
137 CONTINUE
137 SUMC = 0.0
137 DO 280 I = 1, NPKS
137 SUMC = ((SAVE(I) - FINTENS(I))/FINTENS(I))*2 + SUMC
137 SAVER(I) = FINTENS(I)
137 KITER = KITER + 1
137 IF(KITER - 10) 281, 290, 290
137 IF(SUMC - 1.0E-06) 290, 290, 102
137 WRITE OUTPUT TAPE 3, 295, KITER
137 CHECK MATRIX INVERSION
137 IF(NERFL - 1) 219, 219, 1056
137 DISCARD ANY NEGATIVE INTENSITIES AND RESOLVE
137 I = 1 - NPKS
137 IF(FINTENS(I) - 224, 224, 298
137 I = 1 +
137 DO TO 220
137 IF(PKEN(I) - 9.999 ) 232, 225, 225
137 NPKS = NPKS - 1
137 LPKS = LPKS - 1
137 PKENP = PKEN(I)
137 DO 235 J = 1, NPKS
137 DO 235 K = 1, MESH
137 R(K,J) = R(K,J+1)
137 PKEN(J) = PKEN(J+1)
137 E(J) = PKEN(J)
137 PKEN(NPKS + 1) = PKENP
137 E(NPKS + 1) = PKENP

```



```

KITER = 0
GO TO 102 READ OUT SYNTHETIC SPECTRUM
C 298 IF (NPRINT - 1) 291, 299
291 WRITE OUTPUT TAPE 3, 300
WRITE OUTPUT TAPE 3, 875, (DIF(I), I = 1, MESH )
C 299 NORMALIZE AND STORE INTENSITIES
SUMA = 0.0
DO 140 I = 1, LPKS
140 SUMA = FINTENS(I) + SUMA
DO 141 I = NSETDEX, NRUNDEX
FIN(I, NSETDEX) = FINTENS(I) * SBB(NRUNDEX)
DOG = SIGTENS(I)**2 + (FIN(I, NSETDEX)*2)*(EL**2+ECC**2)
SIN(I, NSETDEX) = NRUNDEX * SQRTF(DOG) * SBB(NRUNDEX)
SINTENS(I) = FINTENS(I)/SUMA * 100.0
SIGTENS(I) = SIGTENS(I)/SUMA * 100.0
141 WRITE OUTPUT TAPE 3, 143, SUMA
WRITE OUTPUT TAPE 3, 145, NHWEST, NHWEST
WRITE OUTPUT TAPE 3, 150

9 C CALCULATE COUNTS AND NORMALIZE TO MONITOR COUNTS RATIO
DO 168 K = 1, NPKS
CUTCOUNT = E(K) * 0.4
IA = 1
XSDEF = XS(2) - XS(1)
XSP = XS(IA) - XSDEF/2.0
155 IF ( XSP - CUTCOUNT ) 162, 166, 166
162 IA = IA + 1
160 TO 155
166 FRACT = (XSP - CUTCOUNT)*XSDEF
SUM = R(IA - 1, K) * FRACT
DO 170 I = IA, MESH
170 SUM = R(I, K) + SUM * FINTENS(K) * (0.01 * SUMA)
COUNTS(K) = (0.01 * SUMA) * SIGTENS(K) * SUM
SIGCTS(K) = SIGCTS(K)**2 + (COUNTS(K)*2)*(EL**2 + ECC**2)
SIGCTS(K) = SQRTF(SIGCTS(K))
STORE INTENS SIGNS AND ENERGIES FOR ALL RUNS
OMEGA(NRUNDEX) = (FINTENS(K)/100.0) * SBB(NRUNDEX)
DOG = SIGTENS(K)**2 + (FINTENS(K)*2)*(EL**2+ECC**2)
STW(NRUNDEX, K) = (SQRTF(DOG)/100.0) * SBB(NRUNDEX)
C 168 READ CUT AND GRAPH RESULTS TAPE 3, 165; PKEN(K), COUNTS(K), SIGTENS(K),
1IA, CUTCOUNT, TOTP = 0.0

```



```

TOT = 0.0
FITMEAS = 0.0
DO 310 I = NLOWEST(NHIEST(SPEC(I)) - DIF(I)) + FITMEAS
    TOT = SPEC(I) + TOT
    TOTP = DIF(I) + TOTP
    FITMEAS = (FITMEAS/TOT)* 100.0
    WRITE OUTPUT TAPE 3,320, TOT, TOTP, FITMEAS
    DO 311 I = NCUT
        SPEC(I) = SPEC(I)/SCALE
        DIF(I) = DIF(I)/SCALE
        IF (NPRINT - 2) = 314, 309, 309
        IF (NPRINT - 4) = 316, 314, 316
        CALL GRAPH(MESH,XS,SPEC, 8)
        CALL GRAPH(MESH,XS,DIF, 8)
        DO 312 I = 1, MESH
            DIF(I) = DIF(I) + 1000.0
        CALL GRAPH(MESH,XS,DIF,8)
        GO TO 292
316 READ INPUT TAPE 2,1085, ID1, ID2, ID3, ID4, ID5
C IF CONST VECTOR USED(9.999 MEV) RESTORE AS LAST ELEMENT OF PKEN
292 DO 292 I = 1, MPKS2
    I2 = I + 1
    IF (PKEN(I) - 9.999) 293, 297, 293
293 CONTINUE
    GO TO 289
297 DO 294 I = 1, MPKS
    PKEN(I - 1) = PKEN(I),
294 E(I - 1) = PKEN(I - 1),
    PKEN(MPKS) = 9.999
    E(MPKS) = PKEN(MPKS)
    DO 296 I = 1, MESH
        RI,I, MPKS) = 0.0C2
296 MPKS = MPKS
    LPKS = MPKS2
    IF (NRUNDEX - NRUNNS) 940, 945, 945
940 NRUNDEX = NRUNDEX + 1
    WRITE OUTPUT TAPE 3, 1090
    GO TO 807
945 DO 7004 J = 1, MPKS
    DO 7004 I = 1, NRUNNS
        A(I,J) = FIN(J, NSETDEX, I)/100.0
7004 AINV(I,J) = SIN(J, NSETDEX, I), 1035
    WRITE OUTPUT TAPE 3,1035
C GO TO COEFFICIENT ANALYSIS
    CALL LEASTSQ

```



GO TO 1058  
 1056 WRITE OUTPUT TAPE 3, 1057, NSETDEX, NRUNDEX  
 WRITE OUTPUT TAPE 3, 1051, ((A(I,J), I=1,NPKS), J=1,NPKS)  
 GO TO 316  
 1058 IF SETDEX - NSETDEX = 1053, 1052, 1052  
 1053 NSETDEX = NSETDEX + 1  
 NRUNDEX = 1  
 GO TO 806  
 1052 IF (NSETS - 1) 106, 106, 1054  
 C FIND WEIGHTED MEAN DISTRIBUTION COEFFICIENTS

1054 WRITE OUTPUT TAPE 3, 1060  
 DO 1064 I = 1, 3  
 KB = I\*\*2 - 3\*I + 4  
 DO 1064 J = 1, NPKS  
 SUM1 = 0.0  
 SUM2 = 0.0  
 DO 1063 K = 1, NSETS  
 SUM1 = SUM1 + STC(I, J, K) \* (1.0/(STSC(I, J, K)\*\*2))  
 SUM2 = SUM2 + 1.0/(STSC(I, J, K)\*\*2)  
 TS(I) = SUM1/SUM2  
 TSC(I) = SQRTF(1.0/SUM2)  
 IF (I - 2) 1068, 1067, 1068  
 IF (J - 1) 1067, 1067, 1068  
 1067 WRITE OUTPUT TAPE 3, 1061  
 1068 CHISQ = 0.0  
 SSQ = 0.0  
 DO 1065 K = 1, NSETS  
 IF (STSC(I, J, K) - 0.00001) 1071, 1080, 1080  
 CHISQ = CHISQ + ((STC(I, J, K) - TSC(I))\*\*2 + SSQ)/STSC(I, J, K) )\*\*2  
 1080 SSQ = (STC(I, J, K) - TSC(I))\*\*2 + SSQ  
 1065 FNS = NSETS  
 SSQ = SQRTF( SSQ/((FNS - 1.0)\*FNS) )  
 GO TO 1064

1071 CHISQ = 999.99

1064 WRITE OUTPUT TAPE 3, 1062, E(J), KB, TC(I), TS(I), CHISQ, SSQ  
 C SUMMARY OF DISTRIBUTION COEFFICIENTS  
 WRITE OUTPUT TAPE 3, 1093  
 WRITE OUTPUT TAPE 3, 1094  
 DO 1096 J = 1, NPKS  
 DO 1096 I = 1, 3  
 KB = I\*\*2 - 3\*I + 4  
 NP = (-1\*\*2 + 5\*I + 4)/2  
 1095 PKEN(J), NP, KB, (STC(I, J, K), STSC(I, J),



```

      K)   K = 1, 5
      DO 1120 I = 1, 20
      B1(I) = 0.0
      B2(I) = 0.0
      FIND WEIGHTED MEAN INTENSITIES
      DO 1115 K = 1, NRUNS
      SUM = 0.0
      DO 1102 I = 1, LPKS
      SUM1 = 0.0
      SUM2 = 0.0
      DO 1101 J = 1, NSETS
      IF(SIN(I,J) - 1.0/NSETS) 1106, 1100
      WRITE OUTPUT TAPE 3, 1107, I, J
      GO TO 1108
      SUM1 = SUM1 + SIN(I,J,K)/(SIN(I,J,K)**2)
      SUM2 = SUM1/SUM2
      B(I) = 1.0/SQRT(F(SUM2))
      E(I,K) = B(I)/100.0
      AINV(K,I) = E(I)/100.0
      B1(I) = B1(I) + B(I)*SINF(B3(K)/57.2958)
      B2(I) = B2(I) + (E(I)**2)*(SINF(B3(K)/57.2958))
      .
      SUM = SUM + B(I)
      SUM = SUM/100.0
      WRITE OUTPUT TAPE 3, 1103, B3(K)
      DO 1105 I = 1, LPKS
      B(I) = B(I)/SUM
      E(I) = E(I)/SUM
      WRITE OUTPUT TAPE 3, 1104, PKEN(I), B(I), E(I)
      SUM = 0.0
      DO 1130 I = 1, LPKS
      SUM = SUM + B(I)
      WRITE OUTPUT TAPE 3, 1140
      WRITE OUTPUT TAPE 3, 1170
      DO 1150 I = 1, LPKS
      B1(I) = (B1(I)/SUM)*100.0
      B2(I) = 100.0*SQRT(F(B2(I))/SUM)
      WRITE OUTPUT TAPE 3, 1160, PKEN(I), B1(I), B2(I)
      CALL LEASTSQ FORMAT(3F10.3)
      WRITE OUTPUT TAPE 3, 1113
      FORMAT(31HCHECK MATRIX INVERSION ERFL = 12./)
      106 FORMAT(31HOSUM OF ABSOLUTE INTENSITIES = F15.4 /)
      143 FORMAT(31HOSUM OF ABSOLUTE INTENSITIES FROM LEAST SQUARE FITTING CHANNEL
      145 FORMAT(58HRELATIVE INTENSITIES FROM 13 THROUGH 13 /)
      150 FORMAT(92HO ENERGY COUNTS SIGMA COUNTS CUTOFFCHAN. CUTO

```



```

1 IFHEN INTENS SIGMA //)
2 FORMAT( F8.3, F12.1, F9.1, 117, F11.3, F10.2, F9.2 )
3 FORMAT( 25HNUMBER OF ITERATIONS = 12 /)
4 FORMAT( 23HSYNTHESIZED SPECTRUM = 12 /)
5 FORMAT( 16HINPUT COUNTS = F9.1, 18H FITTED COUNTS = F9.1,
6 HFIT MEASURE = F7.1, 10H PERCENT /)
7 FORMAT( F10.3, 110 )
805 FORMAT( F10.3, 6F10.3 )
810 FORMAT( 7F10.3 )
814 FORMAT( 7F10.3 )
815 FORMAT( 7F10.3 )
816 FORMAT( 7F10.3 )
817 FORMAT( 7F10.3 )
820 FORMAT( 2F10.3 )
825 FORMAT( 2F10.3 )
831 FORMAT( 2F10.3 )
832 FORMAT( 18H1 DATA SET NUMBER 13 /)
833 FORMAT( 3F10.3, 110, F10.3 )
8340 FORMAT( 110, F10.3 )
841 FORMAT( 26HNORMALIZING COEFFICIENT = F7.4 /)
842 FORMAT( 110, F10.3, 3110, F10.3, E10.1 )
848 FORMAT( 3F10.3, 110, F20.3 )
860 FORMAT( 42HUNFOLDING OF GAMMA SPECTRUM - RUN NUMBER 13 /)
865 FORMAT( 11HRESONANCE F10.3, 20H KEV PHOTO NUMBER 14 /)
870 FORMAT( 18HCHANNEL ENERGIES /)
875 FORMAT( 8F12.3 )
880 FORMAT( 18HINITIAL SPECTRUM /)
881 FORMAT( 38H0 INITIAL SPECTRUM MINUS BACKGROUND /)
889 FORMAT( 12H0 BACKGROUND /)
895 FORMAT( 14H0 CURVE NUMBER 12, 15H TO BE STRIPPED /)
900 FORMAT( 40H0 SPECTRUM AFTER STRIPPING CURVE NUMBER 12 /)
920 FORMAT( / 80H COMPONENT NUMBER ENERGY / )
1 INTENSITY COUNTS /)
930 FORMAT( 110, F22.3, F28.3 )
1020 FORMAT( 10H0 ENERGY = F6.3, F28.5H MEV /)
1030 FORMAT( 1P10E1.2 )
1035 FORMAT( 40H1ANALYZE FOR DISTRIBUTION COEFFICIENTS /)
1051 FORMAT( 1P13E9.1 )
1057 FORMAT( 29HOMATRIX X A SINGULAR DATA SET 14, 7H RUN 14 /)
1060 FORMAT( 7HOFIND WEIGHTED MEAN COEFFICIENTS FOR CORRECTED 14 /)
1 SERIES NP=2 /)
1061 FORMAT( 25HOSAME FOR NP=3 SERIES /)
1062 FORMAT( F6.3, 8H MEV - A 11, 7H/A0 = F6.4, 9H SIGMA = F6.4 , 14H C )
1 HISQUARE = F7.2
1073 FORMAT( 26 HOWED 3 } JUL AM DECK 1 /)

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085 FORMAT (A1)
090 FORMAT (2HO /////
093 FORMAT (3SHOSUMMARY OF DISTRIBUTION COEFFICIENTS /)
094 FORMAT (103H MEV NP COEF SIGMA SET1 SIGMA /)
095 FORMAT (3H SET3 SIGMA SET4 SIGMA SETS SIGMA /)
103 FORMAT (19HMEAN INTENSITIES A II, 3H/AO F9.3 9F8.3 /)
104 FORMAT (1H INTENSITY SIGMA /) F5.1, 9H DEGREES /)
107 FORMAT (5HSIN( I1, 1H, I1, 6H) = 0 /)
113 FORMAT (4HEND OF PROGRAM SPECFIT /)
140 FORMAT (3HOVERAGE INTENSITY OVER 4 PI /)
160 FORMAT (3F10.3 /)
170 FORMAT (3H ENERGY INTENS 12 SIGMA /)
5500 FORMAT (15HOHANGUP PRINT /)
END

SUBROUTINE RESPONS
DIMENSION SPEC(130), XS(130), CALB(530), 09), PHOT(09),
1 HOLD(530), Z(20), PKEN(20), PKCTS
2 BACK(130), FINTENS(20), AINV(20), STOR
3 E(20), B(20), SIGCTS(20), COUNTS(20), WT(130), SA
4 V(20), POLY(20, 4), OMEGA(20, 20), CEE(4, 4), PART1(9),
5 XI(4), COEFT(4), SCOEFT(4), ARGP(20), B1(20), B2(20),
6 SUM7(4), COANISO(2), CSBB(20), STSW(20, 20), B3(20),
7 FMU(20), EMU(20), PKSEP(10), BACKND(130), SIC(4, 20, 6),
8 TSC(4, 20, 6), CALB, PHOT, ENCALB, TEMP, BACK, HOLD, Z, DIFTS,
9 COMMON SPECS, XS, FINTENS, R, AINV, STORE, B, SIGTS, SIGCTS,
10 PKCTS, WT, PART1, CEE, INV, XI, COEFT,
11 COUNTS, POLY, OMEGA, SUM7, ANISO, SBB, STSW, BIGT,
12 COEFT, ARGP, B1, B3, MINT, GAME, N, ZERO, NERFL, NRUNS, PKSEP, LPKS
13 MESH, SETDEX, STC, SISCS, NPRINT, HCM
14 DO 675 J = 1, MINT
15 DO 602 I = 1, 0.0
16 HOLD(I) = 0.0
17 TEMP(I) = 25
18 JP = 2
19 JJJ = 1

005 IF (ENCALB(JJ)-E(J) + 1 ) 610, 620, 620
610 JJJ = JJ + 1
620 GO TO 605
620 SHIFT = PHOT(JJ) - PHOT(J-1)
SHIFT LOWER UNDER UPPER
KK = MINT - NSHFT

```



```

DO 630 I=1, KK
  TEMP(I+NSHFT) = CALB(I, JJ-1)
  KK = NSHFT + 1
  SLC = CALB(2, JJ - 1) - CALB(1, JJ - 1)
  DO 640 I = 1, KK
    FLX = I - 1 - NSHFT
    TEMP(I) = CALB(1, JJ - 1) + SLC * FLX
    IF (I-JP) 641 641 641
    WRITE OUTPUT TAPE 3, 400 , E(J)
  400 PH = NSHFT
  641 FRACT = (PHOT(JJ) - PHOT(JJ-1)) - PH
    HOLD(1) = TEMP(1)
    HOLD(2) = TEMP(2)
    KK = MINT - 2
    GFACCTR = (PKSEP(JJ - 1) - PKSEP(JJ)) / PKSEP(JJ - 1)
    DO 315 II = 3, KK
    I = II
    X = I
    FRACT2 = FRACT - GFACTOR*( X - PHOT(JJ))
  642 Y = I - II
    IF(Y + C.0 ) 643 643 643
    IF(Y - 1.0 ) 648 648 648
    643 I = I + 1
    644 GO TO 642
    647 I = I - 1
    648 FRACT2 = Y
    IF(I - 2) 649 649 649
    IF(I - MINT + 1) 652 652 652
    652 HOLD(II) = TEMP(II)
    GO TO 315
  653 CONTINUE
    DEL1 = TEMP(I-1) - TEMP(I-2)
    DEL2 = TEMP(I) - TEMP(I-1) - DEL1
    DEL3 = TEMP(I+1) - TEMP(I) - DEL2 - DEL2 - DEL1
    U1 = 2.0 - FRACT2
    U2 = U1 * (U1 - 1.0)/2.0
    U3 = U2 * (U1 - 2.0)/3.0
    HOLD(II) = TEMP(I-2) + U1*DEL1 + U2*DEL2 + U3*DEL3
  315 CONTINUE
    DO 320 I = 1, MINT
      IF ((HOLD(I) - 0.0 ) 319, 320, 320
  319 HOLD(II) = 0.0
  320 TEMP(II) = HOLD(II)
  C INTERPOLATE BETWEEN CURVES

```



```

646 DD 65C I = (E(J) - ENCALB(JJ-1))/(ENCALB(JJ) - ENCALB(JJ-1))
650 TEMP(I) = TEMP(I) + CALB(I,JJ)/TEMP(I) * EFRAC T
406 AMP = 0.165 + 0.00667 * HCM
PKHTFAC = (1.0/(7.62/(4.0 * HCM)))*XS(2)-XS(1)
SOLIDA = 0.9935 * 0.866 * 0.089 * SQRTF(PKEN(J)))*XS(2)-XS(1)
ATTEN = EXP = 0.9699 * PKHTFAC * 0.00918 * HCM
RED = AMP * PKHTFAC * SOLIDA * ATTEN * (PKEN(J) ** (-EXP))
DO 675 K=1,MESH
CHX = PHOT(JJ)+ (XS(K) - E(J)) * SL
NPHOB = PHOT(JJ)
NCHX = CHX
I = NCHX
FLNCHX = NCHX
FRACT = CHX - FLNCHX
IF(I-1) 390, 381, 374
374 IF(I-MINT(I)) 371, 390, 390
IF(I-DEL1) = TEMP(I) - TEMP(I-1)
DEL2 = TEMP(I+1) - TEMP(I+1) - DEL1
DEL3 = TEMP(I+2) - TEMP(I+1) - DEL2 - DEL1
U1 = 1.0 + FRAC T
U2 = U1*(U1 - 1.0)/2.0
U3 = U2*(U1 - 2.0)/3.0
380 VAL = TEMP(I-1) + U1*DEL1 + U2*DEL2 + U3*DEL3
IF(VAL - 0.0) 391, 391
390 VAL = 0.0
391 GO TO 675

```

```

381 I = I + 1
DEL1 = TEMP(I) - TEMP(I-1)
DEL2 = TEMP(I+1) - TEMP(I) - DEL1
DEL3 = TEMP(I+2) - TEMP(I+1) - DEL2 - DEL1
U1 = FRAC T
U2 = U1*(U1 - 1.0)/2.0
U3 = U2*(U1 - 2.0)/3.0
GO TO 380
675 R(K,J) = VAL * REDREADOUTS , F5.2 , 4 H MEV /
410 FORMAT(20HRESPONSE READOUTS
679 FORMAT(2HO /1P8E12.3)
5500 FORMAT(15HORANGUP PRINT 12 /)
RETURN
END
SUBROUTINE PGAUSS3(N,EP,A,X,KER)

```



```

DIMENSION A(20,20),X(20,20)
DO 1 I=1,N
DO 1 J=1,N
1 X(I,J)=0.0
1 DO 2 K=1,N
2 X(K,K)=1.0
10 DO 34 L=1,N
KP=0
Z=0.0
DO 12 K=L,N
12 IF(Z-ABSF(A(K,L))>0)11,12,12
11 Z=ABSF(A(K,L))
KP=K
12 CONTINUE
12 IF(L-KP)>13,20,20
13 DO 14 J=L,N
14 Z=A(L,J)
A(L,J)=A(KP,J)
A(KP,J)=Z
14 DO 15 J=1,N
15 X(K,P,J)=Z
Z=X(L,J)
X(L,J)=X(KP,J)
15 IF(ABSF(A(L,L))-EP)>50,50,30
20 IF(L-N)>31,34,34
30 LP=L+
31 LP=L+
32 DO 36 K=LP1,N
36 IF(A(K,L))32,36,32
32 RAT10=A(K,L)/A(L,L)
30 DO 33 J=LP1,N
33 A(K,J)=A(K,J)-RAT10*A(L,J)
30 DO 35 J=1,N
35 X(K,J)=X(K,J)-RAT10*X(L,J)
36 CONTINUE
34 CONTINUE
40 DO 43 I=1,N
43 I=N+1-J=1,N
DO 43 J=1,N
43 S=0.0
41 IF((I-N)>41,43,43
41 IIP1=IIP1+1
40 DO 42 K=IIP1,N
42 S=S+A(IIP1,K)*X(K,J)
43 X(IIP1,J)=(X(IIP1,J)-S)/A(IIP1,IIP1)
43 KER=1
RETURN

```



50 KER=2

END  
SUBROUTINE LEASTSQ  
DIMENSION SPEC(130), CALB(530), PHOT(09), ENCALB(C9)  
TEMP(530), HOLD(530), DIF(530), PKEN(20), PKCTS  
BACK(09), HNSL(20), A(20), NV(20), STQR  
T(20), R(130), INTENS(20), COUNTS(20), W(20), SA  
E(20), B(20), SIGCTS(20), COUNTS(20), PART(9),  
V(20), POLY(20), SIGENTS(20), CEEINV(4),  
X(4), COEFT(4), OMEGA(20), CEE(4), B(20), B(20),  
SUM(4), COEF(4), SCOEFT(4), ARGP(20), B(20), B(20),  
FMU(20), ANISO(2), SBB(20), STW(20), BTGT(530),  
STSC(4,20,6), EMU(2), PKSEP(10), BACKGND(130), STCW(4,20,6),  
COMMON SPECS, XS, CALB, PHOT, ENCAB, TEMP, BACK, HOLD, Z, DIFFS,  
PKEN, PKCTS, R, INTENS, A, AINV, STORE, B, SIGTENS, SIGFS,  
1,2 COUNTS, WT, PART, CEE, COEFT,  
1,2 COEFT, ARGP, SAVE, OMEGA, SUM, CEEINV, XI, COEFT,  
3 SCOEFT, ARGP, B1, B2, B3, SUM, ANISO, SBB, STW, STSW, BTGT,  
4 MESH, MINT, GAME, ZERO, NERFL, NPKS, NRUNS, PKSEP, LPKS  
5,6 NSETDX, STC, SIS, STC, STC, NPRINT, HCM  
NP = 3  
JNN = 1

NN = NPKS  
RA = 57.29577795  
ALPHA1 = 1.0  
ALPHA2 = 0.0  
ALPHA3 = 0.0  
ALPHA4 = 0.0  
B(1) = 1.0  
B(2) = 0.9295  
B(3) = 0.7780  
B(4) = 0.4955  
LIST GAMMA ABSORPTION COEFFICIENTS AND CORRESPONDING ENERGIES  
FMU(1) = 0.834  
FMU(2) = 0.0760  
FMU(3) = 0.0588  
FMU(4) = 0.0475  
FMU(5) = 0.0416  
FMU(6) = 0.0356  
FMU(7) = 0.0328  
FMU(8) = 0.0316  
FMU(9) = 0.0309  
FMU(10) = 0.0306  
FMU(11) = 0.0309  
FMU(12) = 0.0309



```

    EMU(1) = 0.5
    EMU(2) = 0.6
    EMU(3) = 0.8
    EMU(4) = 1.0
    EMU(5) = 1.5
    EMU(6) = 2.0
    EMU(7) = 3.0
    EMU(8) = 4.0
    EMU(9) = 5.0
    EMU(10) = 6.0
    EMU(12) = 10.0
    IF (NPRINT-3) 201,199,199
    WRITE OUTPUT TAPE 3,711
    CLEAN STORAGE
    DO 7021 I = 1, 20
    B1(I) = A(I,JNN)
    B2(I) = AINV(I,JNN)
    DO 7021 J = 1, 20
    OMEGA(I,J) = 0.0
    STW(I,J) = 0.0
    STSW(I,J) = 0.0
    GAMEN = PKEN(JNN)
    JK = 1
    CORRECT FOR ATTENUATION IN COPPER TARGET BACKING
    IF (NP-L) 1243, 1210
    1243 I = 1
    IF( I - 13) 1245, 1251
    1244 IF( I - EMU(I)-GAMEN) * 1251, 1250
    1245 I = 1 + 1
    GO TO 1244
    1251 CNUM = FMU(I) - FMU(I-1)
    CDEM = FMU(I) - FMU(I-1) + (CNUM/CDEM) * (GAMEN - EMU(I-1))
    FMUE = 8.89 * FMUE
    IF (NPRINT-3) 1252, 1249, 1249
    WRITE OUTPUT TAPE 3,1235, GAMEN
    DO 1260 I = 1, L
    THICK = 0.046 / (COSF((B3(I) - 45.0)/57.29578))
    ATTENC = EXPF(-FMUE * THICK)
    B1(I) = B1(I) / ATTENC
    B2(I) = B2(I) / ATTENC
    IF (NPRINT-3) 1254, 1260
    WRITE OUTPUT TAPE 3,1253, B1(I), B2(I), B3(I)
    1254 CONTINUE
    1260 FUNCIF(A,B,C,G,D) = A + (B/4.0)*(3.0*COSF(2.0*D/RA) + 1.0) +
```



```

1 (0.015625*C)*(35.0*COSF(4.0*D/RA) + 20.0*COSF(2.0*D/RA) + 9.0) +
2 (6.012.0)*COSF(6.0*D/RA) + 126.0*COSF(4.0*D/RA) + 9.0) +
3 SET UP POLYNOMIAL MATRICES
DO 203 I=1,L
ARGP(I) = COSF((B3(I)/RA))
POLY(I,1) = 1.0
POLY(I,2) = 0.5*(3.0*(ARGP(I)**2)-1.0)
POLY(I,4) = 0.0625*(231.0*ARGP(I)**6-315.0*ARGP(I)**4+105.0*
1 ARG P(I)**2-5.0)
203 POLY(I,3) = 0.125*(35.0*(ARGP(I)**4)-30.0*(ARGP(I)**2)+3.0)

IF (JK-1)206,204
204 DO 205 I=1,L
POLY(I,1) = 1.0
POLY(I,2) = ARGP(I)**2
POLY(I,4) = ARGP(I)**6
205 POLY(I,3) = ARGP(I)**4
206 IF (NP-I-3)1450,1460, 1470
1450 DO 1455 I=1,L
POLY(I,3) = 0.0
1455 POLY(I,4) = 0.0
1456 GO TO 1470
1460 DO 1465 I=1,L
1465 POLY(I,4) = 0.0
1470 CONTINUE
C COMPUTE CEE MATRIX
DO 215 I=1,L
OMEGA(I,I) = ((1.0)/B2(I))**2
215 DO 225 J=1,NP
DO 225 J=1,NP
CEE(I,J) = 0.0
DO 225 K=1,NP
CEE(I,J,K) = CEE(I,J) + OMEGA(K,K)*POLY(K,I)*POLY(K,J)
C INVERT CEE
DO 7020 I=1,20
DO 7020 J=1,20
STORE(I,J) = 0.0
7020 STORE(I,J) = 0.0
DO 7030 I=1,NP
DO 7030 J=1,NP
STORE(I,J) = CEE(I,J)
7030 STORE(I,J) = CEE(I,J)
CALL PGAUSS3(NP,ZERO,J)
DO 230 I=1,NP
DO 230 J=1,NP

```



```

230 CEEINV(I,J) = STSW(I,J)
241 IF(NERFL-1)241, 1056
241 FORMAT(3F20.12/)
246 DO 246 J=1, NP
      XI(J) = 0.0
      DO 246 I=1,L
      XI(J) = POLY(I,J)*OMEGA(I,I)*B1(I) + XI(J)
      COMPUTE COEFFICIENTS AND THEIR SIGMAS
      DO 248 I=1, NP
      SUM81 = 0.0
      DO 247 K=1, NP
      SUM81 = CEEINV(I,K)*XI(K) + SUM81
      COEFT(I) = SUM81
      DIV = COEFT(I)
      DO 249 I=1, NP
      COEFT(I) = COEFT(I)/DIV
      SCOEFT(I) = SQRTF(ABSF(CEEINV(I,I))/DIV)
      DO 260 I=1, NP
      SCOEFT(I) = SCOEFT(I)*(CEEINV(I,I)/ABSF(CEEINV(I,I)))
      TEMP(I) = 0.0
      VEEESQ1 = (B1(J)/B2(J))*2 + VEEESQ1
      DO 993 VEEESQ1 = 0.0
      VEEESQ2 = 0.0
      DO 994 K=1,NP
      VEEESQ2 = COEFT(K)*XI(K)*DIV + VEEESQ2
      E2 = NP
      EPSSQ = (VEEESQ1 - VEEESQ2)/(E2 - EP)
      DO 8100 CO I=1, NP
      BRA = (COEFT(I)*2) * CEEINV(I,I) + CEEINV(I,I) - 2.0*COEFT(I)*
      CEEINV(I,I)
      BRA = BRA*EPSSQ
      SCOEFT(I) = (BRA/ABSF(BRA))*SQRTF(ABSF(BRA))/DIV
      COMPUTE ANISOTROPY
      IP = NP+1
      DO 259 I=1,4
      COEFT(I) = 0.0
      SCOEFT(I) = 0.0
      TEMP(I+10) = 0.0
      IF(JK-2)261, 262, 262
      W0 = WFUNCT(COEF(I),COEFT(2),COEFT(3),COEFT(4),0.0)
      W90= WFUNCT(COEF(I),COEFT(2),COEFT(3),COEFT(4),90.0)
      IF(NP-3)251, 252, 254

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```

251 STC(1, JNN, NSETDEX) = TEMP(2)
      STSC(1, JNN, NSETDEX) = TEMP(2 + 10)

252 DO 253 KE = 2, 3
      STC(KE, JNN, NSETDEX) = TEMP(KE)
253 STSC(KE, JNN, NSETDEX) = TEMP(KE + 10)
254 GO TO 264
262 WO = 0.0
W90 = COEFT(1)
DO 263 I=1, NP
      WO = W0 + COEFT(I)
263 ANISO(JK) = W0/W90 - 1.0
P1 = ((1.0/W90)**2
      SAN(JK) = 0.0
DO 7110 I = 1, NP
      Z(I) = (W90*POLY(1, 1) - W0*POLY(1, 1))**2
7110 SAN(JK) = SAN(JK) + Z(I) * (SCOEFT(I)**2)
      SAN(JK) = P1 * SQRT(F(SAN(JK)))
IF (NPRINT-3) 265, 7107, 7107,
265 IF (JK-2) 274, 339, 339
274 CONTINUE
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=275)
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=342)
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=1370)
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=342), TEMP(I+10), I=1, NP)
7107 IF (NPRINT-3) 7108, 7108
7108 WRITE(UNIT=TAPE, IOSTAT=3, FMT=7112)
      TEMP(12), TEMP(3), TEMP(13), TEMP(14), TEMP(11), TEMP(1),
      GO TO 1058
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=295)
      EPSSQ
      IF (NPRINT-1) 298, 320
298 IF (B3(3)-180.0) 297, 320
C COMPUTE POINTS ON THEORETICAL AND EXPERIMENTAL CURVES
297 WRITE(UNIT=TAPE, IOSTAT=3, FMT=296)
      SUM9 = 0.00
300 WFUNCT = WFUNCT(COEFT(1), COEFT(2), COEFT(3), COEFT(4), SUM9)/
      1 WFUNCT(COEFT(1), COEFT(2), COEFT(3), COEFT(4), 090.0)
315 VFUNCT = WFUNCT(ALPHA1, ALPHA2, ALPHA3, ALPHA4, SUM9)/WFUNCT
      1 (ALPHA1, ALPHA2, ALPHA3, ALPHA4, 090.0)
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=324)
      SUM9 + 15.0
      IF (SUM9 - 090.0) 300, 300, 314
314 WRITE(UNIT=TAPE, IOSTAT=3, FMT=316
      WRITE(UNIT=TAPE, IOSTAT=3, FMT=1316

```



```

ADIV = DIV * WFUNCTF(COEFT(1), COEFT(2), COEFT(3), COEFT(4), 090.0)
DO 318 J = 1, L
AB1 = B1(J) / ADIV
AB2 = B2(J) / ADIV
318 WRITE OUTPUT TAPE 3, 317, B3(J), AB1, AB2
60 TO 338
320 WRITE OUTPUT TAPE 3, 296
SUM10 = 270.0
325 1 WFUNCT = WFUNCTF(COEFT(1), COEFT(2), COEFT(3), COEFT(4), SUM10) /
335 VFUNCT = WFUNCTF(COEFT(1), COEFT(2), COEFT(3), COEFT(4), 270.0)
335 1 (ALPHA1, ALPHA2, ALPHA3, ALPHA4, 270.0) / WFUNCT
WRITE OUTPUT TAPE 3, 324, SUM10, WFUNCT, VFUNCT
SUM10 = SUM10 + 15.0
IF (SUM10 - 360.0) 325, 325, 634
634 WRITE OUTPUT TAPE 3, 635
WRITE OUTPUT TAPE 3, 1635
ADIV = DIV * WFUNCTF(COEFT(1), COEFT(2), COEFT(3), COEFT(4), 270.0)
DO 637 J = 1, L
AB1 = B1(J) / ADIV
AB2 = B2(J) / ADIV
637 WRITE OUTPUT TAPE 3, 636, B3(J), AB1, AB2
338 JK = 2
GD TO 204
339 WRITE OUTPUT TAPE 3, 340
WRITE OUTPUT TAPE 3, 1241, ANISO(1), ANISO(2)
WRITE OUTPUT TAPE 3, 1340
WRITE OUTPUT TAPE 3, 1241, SAN(1), SAN(2)
IF (NPRINT - 1) 343, 344, 344
343 WRITE OUTPUT TAPE 3, 341
WRITE OUTPUT TAPE 3, 342, (COEFT(1), SCOEFT(1), I = 1, NP)
344 WRITE OUTPUT TAPE 3, 336
GO TO 1058
1056 WRITE OUTPUT TAPE 3, 1057, E(JNN)
1058 JNN = JNN + 1
JK = 1
IF (JNN - NN) 201, 201, 1210
1210 IF (NP - 2) 1212, 1212, 1211
1211 NP = NP - 1

```

```

JNN = 1
60 TO 201
1212 CONTINUE
DO 351 1 = 1, 20
DO 351 J = 1, 20

```



```

STORE(I,J) = 0.0
AINV(I,J) = 0.0
351 A(I,J) = 0.0
L=0

RETURN(72)OLEAST SQUARE FIT COEFFICIENTS OF LEGENDRE POLYNOMIALS A
1ND THEIR SIGMAS
275 FORMAT(F9.6)
292 FORMAT(18H0EPSILON SQUARE =E20.6/)
295 FORMAT(45H0EXPONENTIAL FUNCTION THEORETICAL FUNCTION/)
296 FORMAT(42H0NORMALIZED EXPERIMENTAL COUNTS AND SIGMAS/)
316 FORMAT(42H0NORMALIZED EXPERIMENTAL COUNTS AND SIGMAS/)
317 FORMAT(F20.2,2F20.5)
324 FORMAT(F7.2,F10.6,F20.6)
336 FORMAT(2H0,///)
340 FORMAT(52H0 ANISOTROPY-P ANISOTROPY-C
341 FORMAT(42H0COEFFICIENTS AND SIGMAS FOR COSINE SERIES/)
342 FORMAT(2F10.5)
335 FORMAT(42H0NORMALIZED EXPERIMENTAL COUNTS AND SIGMAS /)
6336 FORMAT(F20.2,2F20.5)
1057 FORMAT(22H0MATRIX CEE SINGULAR F6.3, 11H MEV GAMMA /)
1203 FORMAT( F10.6)
1235 FORMAT(49HOSTART ANALYSIS FOR DISTRIBUTION COEFFICIENTS OF F7.3
1241 FORMAT(14H MEV GAMMARAY /14H0INPUT COUNTS
1316 FORMAT(2F20.12,/)
1316 1MA FORMAT(81H0 ANGLE COUNTS
1370 FORMAT(50H0COEFFICIENTS CORRECTED FOR SOLID ANGLE
1340 FORMAT(52H0 SIGMA ANISO-P ANGLE COUNTS
1635 FORMAT(81H0 1 SIGMA )
1253 FORMAT( 8H0COUNTS= F10.1, 7H SIGMA= F10.1, 7H ANGLE=F6.1 )
1NC, F10.7)
7111 FORMAT( 15H0ENERGY NP A0 SIGMA A2 SIGMA
1A4 SIGMA EPSSQ ANISO SIGMA
7112 FORMAT( F7.3, 18, 10F10.3 /)
END
END

```



H. Fortran Listing of Program CONTOUR

```

••JOB* PHELPS CONTOUR - GRAPHS
PROGRAM CONTOUR
DIMENSION A(4,4), B(4,4), C(4,4), Y(4,4)
1ALPHA(4,4), BETA(4,4), GAMMA(4,4), DELTA(4,4),
2YP(900), XP(900), YP2(1800), XP2(1800), LITTLE(10),
SFY = 30.0
MINOFFX=4
MINOFFY=4
LABELNO=0
MODE=1
MTITLE=4
NODCURV=2
NODCURV=3
LABEL = 0
DO 10 I = 1, 10
10 LTITLE(1) = 1, (8H PHELPS )
LTITLE(2) = 8H CONTOUR
RA=57.2957795131
PI=3.1415926536
BSINF(A) = ATANF( A/SQRTF( 1.0 - A**2 ) )
WRITE OUTPUT TAPE 3, 180
WRITE OUTPUT TAPE 3, 190
READ INPUT TAPE 2, 190 NSETS
DO 270 MA = 1, NSETS
MA = MA
READ INPUT TAPE 2, 40, JFIND, JSTEP, JFIX, IMAX, JMAX, NCONT, NPTS
NPTS = NPTS * 2
READ INPUT TAPE 2, 50, DELTA(JFIX) RHO
READ INPUT TAPE 2, 60, ((A(I,J), B(I,J), C(I,J)), J=1, JMAX), I=1,
IMAX
IF (RHO - 0.0) 24, 26, 24
24 DO 25 I = 1, IMAX
25 BI(1) = B(1,1)*RHO
26 WRITE OUTPUT TAPE 3, 200
WRITE OUTPUT TAPE 3, 210
WRITE OUTPUT TAPE 3, 220, ((N, J, A(N,J), B(N,J), C(N,J)), J=1, JMAX)
1) N=1, IMAX)
DO 270 IB = 1, IMAX
1 = 1, IMAX
1 = 1, 18
WRITE OUTPUT TAPE 3, 5500, NHANG
NHANG = NHANG + 1
MODCURV = 1
I = 1 * 2
DO 85 J = 1, JMAX

```



```

ALPHA(I,J) = (A(I,J) + C(I,J))/2.0
IF(B(I,J) = 0.0) S = 80,70,80 * ((A(I,J)-C(I,J))/ABSF(A(I,J)-C(I,J)))
70 GAMMA(I,J) = (PI/2.0) * ((A(I,J)-C(I,J))/ABSF(A(I,J)-C(I,J)))
S = 1.0
GO TO 85
80 GAMMA(I,J) = ATANF((A(I,J))- C(I,J))/B(I,J)
S = B(I,J)/ABSF(B(I,J))
85 BETAI(I,J) = S*0.5* SQRFF( B(I,J)**2 + (A(I,J)-C(I,J))**2)
90 IF(JFIX = 4) 100,100,100
100 Y(JFIX) = ATANF(DELTA(JFIX))
GFIJX = ALPHA(I,JFIX) + BETAI(I,JFIX) * SINF(2.0*Y(JFIX) + GAMMA
I(JFIX)
1 GRA = GAMMA(I,JFIX) * RA
      WRITE OUTPUT TAPE 3, 87, JFIX, DELTA(JFIX), BETA(JFIX), GRA
      WRITE OUTPUT TAPE 3, 88, I,JFIX, ALPHA(JFIX), GRA
      GO TO 120
110 GFIJX = 1.0
120 FLNC = NCONTS
X1 = (ALPHA(I,JSTEP) + BETA(I,JSTEP))*(ALPHA(I,JIND)+BETA(I,
1 JIND)
1 X12 = (ALPHA(I,JSTEP) + BETA(I,JSTEP))*(ALPHA(I,JIND)-BETA(I,
1 JINC)
1 X21 = (ALPHA(I,JSTEP) - BETA(I,JSTEP))*(ALPHA(I,JIND)+BETA(I,
1 JINC)
1 X22 = (ALPHA(I,JSTEP) - BETA(I,JSTEP))*(ALPHA(I,JIND)-BETA(I,
1 JIND))
1 AMAX = MAX(IF(X1,X12,X21,X22)* GFIX
AMIN = -MAX(IF(-X1,-X12,-X21,-X22)* GFIX
DEL1 = RA * (PI/2.0 - GAMMA(I,JSTEP)) * 0.5
DEL2 = RA * (PI/2.0 - GAMMA(I,JIND)) * 0.5
      WRITE OUTPUT TAPE 3, 125, I, AMAX, DEL1
      WRITE OUTPUT TAPE 3, 170, JIND, JSTEP, JFIX, DELTA(JFIX), II
DO 265 K = 1, NCONT
LDEX = 1
FLK = K
K2 = FLK/2.0
COEFT(I) = AMIN + ((FLK - 0.5)/FLNC) * (AMAX - AMIN)
FLNPTS=NPTS
DO 16C N = 1, 2
AY1 = 2 - N
AY2 = N - 1
LNA = NPTS*(N - 1) + 2 - N
LNB = NPTS * N
...
```



```

DO 160 L = LNA, LNB
FLL=L - LNA
Y(JSTEP) = (FLL/FLNPTS - 0.5) * PI
GSTEP = ALPHAI(JSTEP) + BETA(I,JSTEP) * SIN(2.0*Y(JSTEP)) +
1 GAMMA(I,JSTEP)
GFIND=COEFT(I)/(GFI X*(GSTEP))
YY=(GFIND-ALPHA(I,JFIN)) /BETA(I,JFIN)
IF(ABSF(YY)-1.0)130,140,AY1*BSINF(YY) + AY2*(PI - BSINF(YY))
130 YY = AY1*BSINF(YY) + AY2*(PI - BSINF(YY))
GO TO 150
140 XP(L) = -95.0
      YP(L) = -95.0
GO TO 160
150 YY = TAN(0.5 * (YY - GAMMA(I, JFIN)))
      YP(L) = ATANF(YY) * RA
      XP(L)=Y(JSTEP)*RA
160 CONTINUE
      WRITE(UNIT=1, IOSTAT=ERR) TAPE3,230, JSTEP
      WRITE(UNIT=1, IOSTAT=ERR) TAPE3,240, (XP(L), L=1,LNB, 10)
      WRITE(UNIT=1, IOSTAT=ERR) TAPE3,250, JFIN,I, COEFT(I)
      WRITE(UNIT=1, IOSTAT=ERR) TAPE3,260, (YP(L), L=1,LNB, 10)
L2 = 265
      IF(YP(L) = 95.0, NGPS, 300
      300 YP2(L2) = YP(L)
      XP2(L2) = XP(L)
      IF((YP(L+1)+95.0) 318, YP(L), 330, 318
      318 IF((ABSF((YP(L+1)-YP(L))-5.0) 320, 330, 330,
      319 IF((ABSF((XP(L+1)-XP(L))-5.0) 319, 330, 330
      320 L2 = L2 +
      320 GO TO 265
      330 IF((L2 - 1) 265, YP2, 312
      331 CALL GRAPH2(L2,XP2,YP2,8,MODCURV, LABLE,LITLE,SFX,SFY,
      332, MINOFFX,MINOFFY,LABLE,NODE,NTITLE,M)
      MODCURV = 2
      L2 = 1
      265 CONTINUE
      WRITE(UNIT=1, IOSTAT=ERR) TAPE3, 5500, NHANG
      NHANG = NHANG + 1
      XP(1) = -95.0
      XP(2) = -95.0
      YP(1) = -95.0
      YP(2) = -95.0
      L2 = 2
      CALL GRAPH2(L2,XP2,YP2,8,NODCURV, LABLE,LITLE,SFX,SFY,

```



```

1 MINOFFX,MINOFFY,LABLEENO,MODE,MTITLE,M)
5001 WRITE OUTPUT TAPE 3, 5500, NHANG
NHANG = NHANG + 1
CALL GRAPH2(L2,XP2,YP2,8,LODCURV,LABLE,LTITLE,SFX,SFY,
1 MINOFFX,MINOFFY,LABLEENO,MODE,MTITLE,M)
5004 WRITE OUTPUT TAPE 3, 5500, NHANG
NHANG = 0
CONTINUE
WRITE OUTPUT TAPE 3,280
30 FORMAT(110)
40 FORMAT(710)
50 FORMAT(2F10.3)
60 FORMAT(3F10.3)
87 FORMAT(45HOSINE CURVE FOR FIXED MIXING PARAMETER DELTA 11,
1 23H WHICH HAS BEEN SET TO F6.3, 18H FOR COMPUTATION /)
1 23H WHICH HAS BEEN SET TO F6.3, 18H FOR COMPUTATION /
88 FORMAT(2H0G 211, 3H + F7.3, 1H) /
125 FORMAT(2H0A 11, 6HMAX = F7.4, 3H A 11, 6HMIN = F7.4, 1H X AND
1 Y ANGLES = 2F9.4 /)
170 FORMAT(7HDELTA 11,21H IS INDEPENDENT DELTA 11,23H IS INDEPENDENT
1 DELTA 11,10H FIXED AT F6.4, 18H LEGENDRE ORDER = 11 /)
180 FORMAT(120H1PROGRAM CONTOUR - P A PHELPS - DELTA 1 INDICES AS FOLLOW
1S - 3 REFERS TO OBSERVED GAMMA MIXING, 1 REFERS TO PROTON PARTIAL
190 FORMAT(116HOWAVE MIXING, 4 REFERS TO CHANNEL SPIN MIXING, 2 REFERS
1 TO UNOBSERVED GAMMA MIXING AND 5 MEANS P-GAMMA TRANSITIONS /
200 FORMAT(35H0
210 FORMAT(90HLEGENDRE ORDER DELTA INDEX
1 B
1 A
1 C
1 ) )
220 FORMAT(18, 116, 3F20.3 /)
230 FORMAT(37H0 VALUES IN DEGREES OF ARCTAN DELTA 11 /)
240 FORMAT(10F10.3)
250 FORMAT(26H0 VALUES OF ARCTAN DELTA 11, 2H A 11, 5H/A0= F5.3 /)
260 FORMAT(10F10.3)
268 FORMAT(A1)
280 FORMAT(25H0END OF PROGRAM--CONTOUR
5500 FORMAT(20H0HANG UP PRINT OUT
13
1 ) )
5001 FORMAT(614)
END
END

```



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